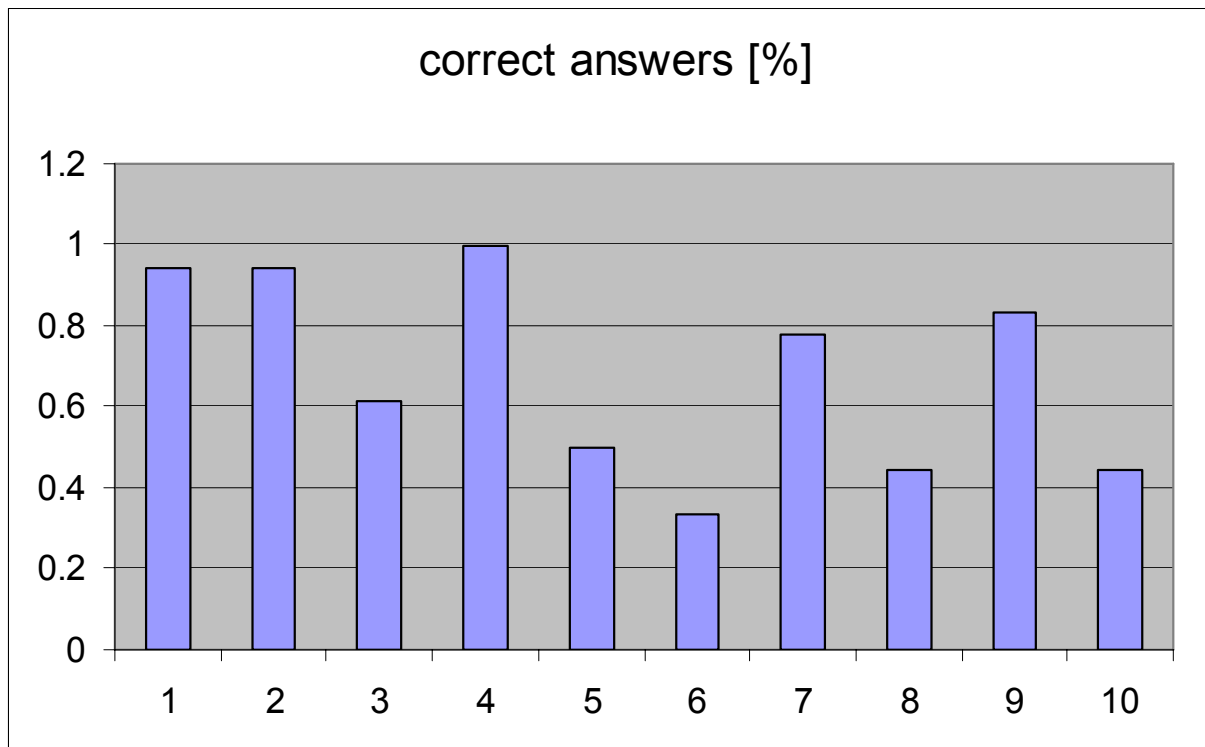




# The art of calorimetry part II

Erika Garutti  
DESY

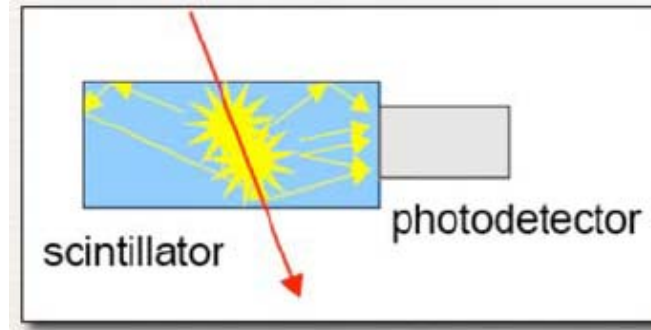
# Result from yesterday starting test



# How to “look” at the signal

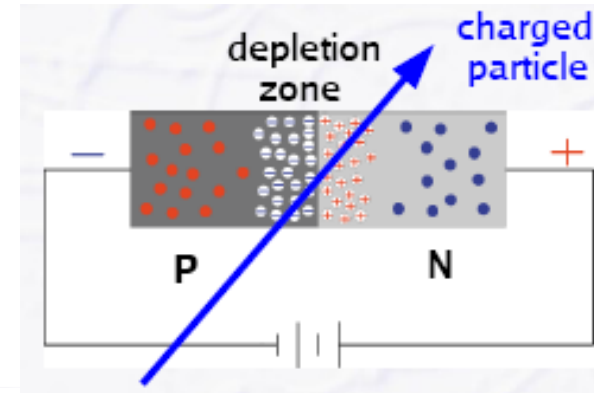
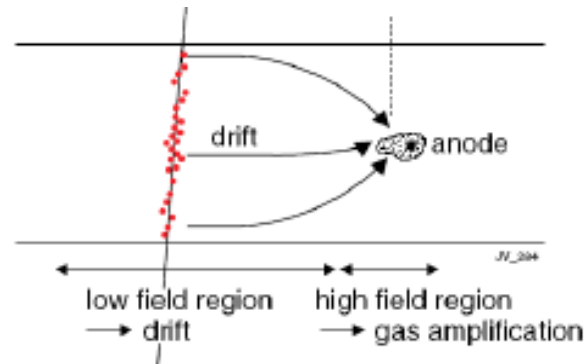
1) Convert particle energy to **light**:  
scintillator (org. / in-org.)

& measure light:  
PMT / APD / HPD / SiPM ...



2) Measure ionization E:  
gas  
noble liquids  
semiconductors

& measure charge signal



3) Measure temperature:

specialized detectors for: DM, solar vs, magnetic monopoles, double  $\beta$ -decay  
very precise measurements of small energy deposits  
phenomena that play a role in the 1 Kelvin to few milli-Kelvin range



# The measurement of showers

or “from energy to signal”

Step 1: Convert energy to light

Step 2: Convert light to electrical signal

Step 3: Reading an electrical signal



# The measurement of showers

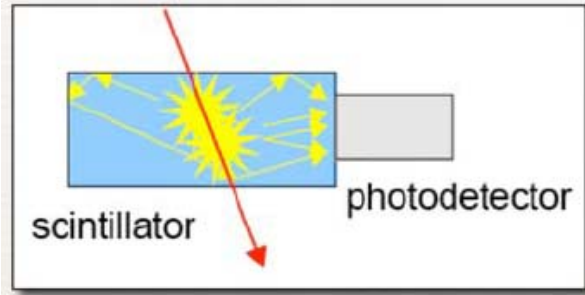
or “from energy to signal”

**Step 1: Convert energy to light**

Step 2: Convert light to electrical signal

Step 3: Reading an electrical signal

# Introduction to scintillators



Energy deposition by a ionizing particle  
→generation  
→transmission  
→detection

} of scintillation light

Two categories: Inorganic and organic scintillators

Inorganic  
(crystalline structure)

Up to 40000 photons per MeV  
High Z  $>2.5 \text{ eV} / \text{photon}$   
Large variety of Z and  $\rho$   
Undoped and doped  
ns to  $\mu\text{s}$  decay times  
Expensive

→ E.m. calorimetry (e,  $\gamma$ )  
Medical imaging  
Fairly Rad. Hard (100 kGy/year)

Organic  
(plastics or liquid solutions)

Up to 10000 photons per MeV  
Low Z  $>10 \text{ eV} / \text{photon}$   
 $\rho \sim 1 \text{ gr/cm}^3$   
Doped, large choice of emission wavelength  
ns decay times  
Relatively inexpensive

→ Tracking, TOF, trigger, veto counters,  
sampling calorimeters.  
Medium Rad. Hard (10 kGy/year)

# Scintillators in brief

**Convert energy** deposited by charged particles or high energy photons **into light**: atoms or molecules of the scintillating medium are excited and decay emitting photons which are detected and converted into electric signals (photo-detector).

Scintillating materials:

**organic**: aromatic hydrocarbon compounds, solid crystals, plastics or liquids. Typically **faster** but have **lower light yields**

**inorganic**: ionic crystals doped with activator centres or glasses.  
Typically **larger light yields**

Two types of light emission:

**Fluorescence**: prompt ns  $\rightarrow$   $\mu$ s in visible wavelength range, temperature independent (component useful for particle detection)

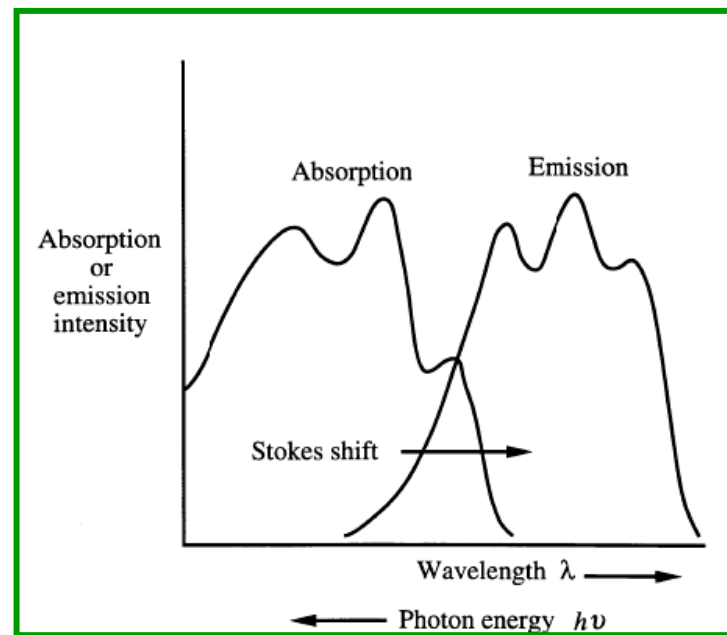
**Phosphorescence**: emission over longer period  $\mu$ s  $\rightarrow$  ms, hrs with longer wavelength and temperature dependent

# Properties of scintillators

Parameters characterizing scintillators:

- **Efficiency**  $R_s$  = average n. of emitted photons/energy of incident radiation
- **Scintillation yield** =  $R_s h\nu$  with  $h\nu$  = energy of emitted photons
- **Time response**: depends on decay time of fast component
- **linearity**

• the scintillator should be transparent to its own scintillation light  
(**Stokes' shift**: the emission wavelength is longer than absorption one)





# Organic scintillators

## Organic scintillators

(plastics or liquids) are composed of aromatic hydrocarbon compounds. Typically consist of solvent + scintillator and a secondary fluor as wavelength shifter.  
**No crystal structure is needed.**

The emission of light is due to **excitation of molecular levels in a primary fluorescent material that emits UV light during de-excitation.**

This light is absorbed in most organic materials with an absorption length of ~mm. The extraction of a light signal becomes possible only by introducing a second fluorescent material in which the UV light is converted into visible light (**wavelength shifter**).

Organic scintillators are typically made of **low Z** materials and have **low density**. Hence the **main interaction >20keV** process is not photoelectric absorption (such as in the case of inorganic scintillators) but **Compton scattering**. Typically, because of the low density more volume is required to obtain a reasonable detection efficiency, but they have low cost.

# Reminder : Photons

Increasing Z

• *Photons* interact by:

1) Photoelectric effect

$$\sigma \propto Z^5, E^{-3}$$

2) Compton scattering

$$\sigma \propto Z, E^{-1}$$

3) Conversion into  $e^+e^-$

$\sigma$  increases with  $E, Z$ , asymptotic at  $\sim 1$  GeV

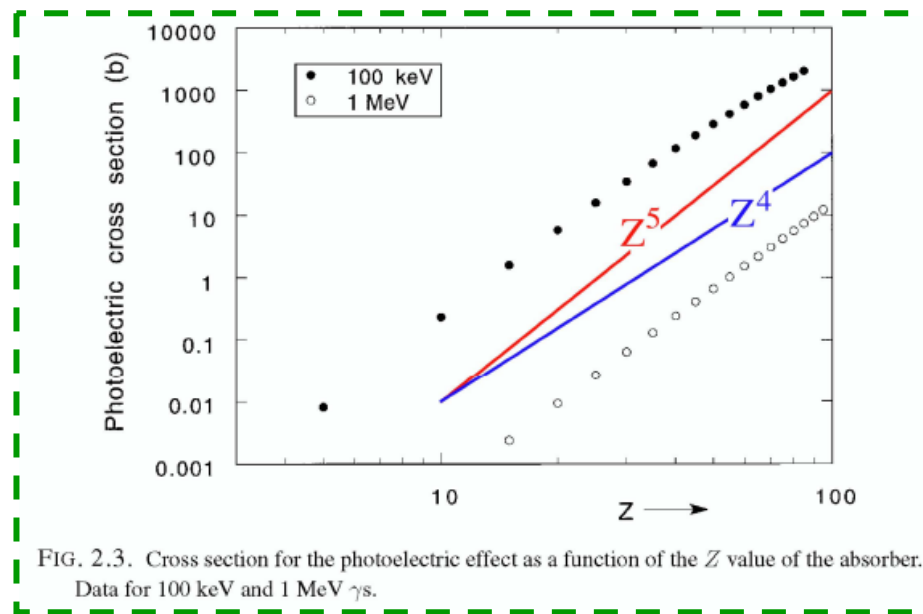
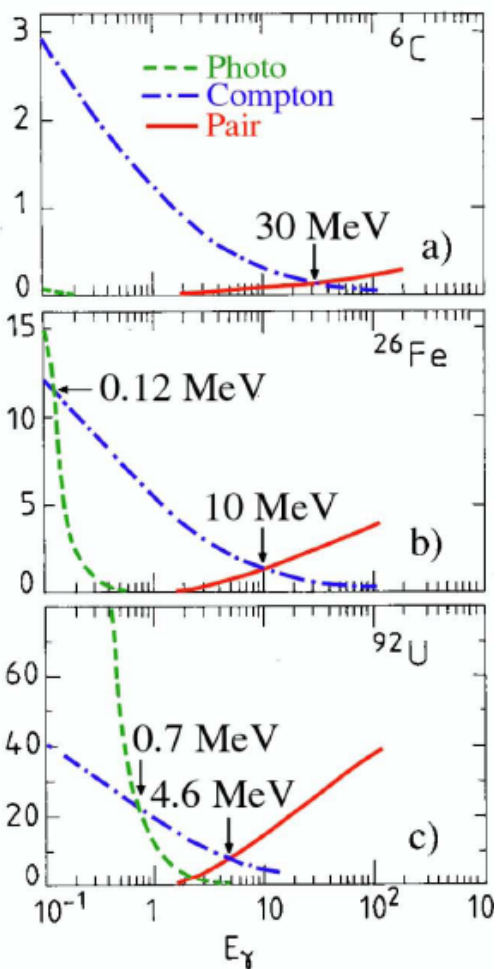


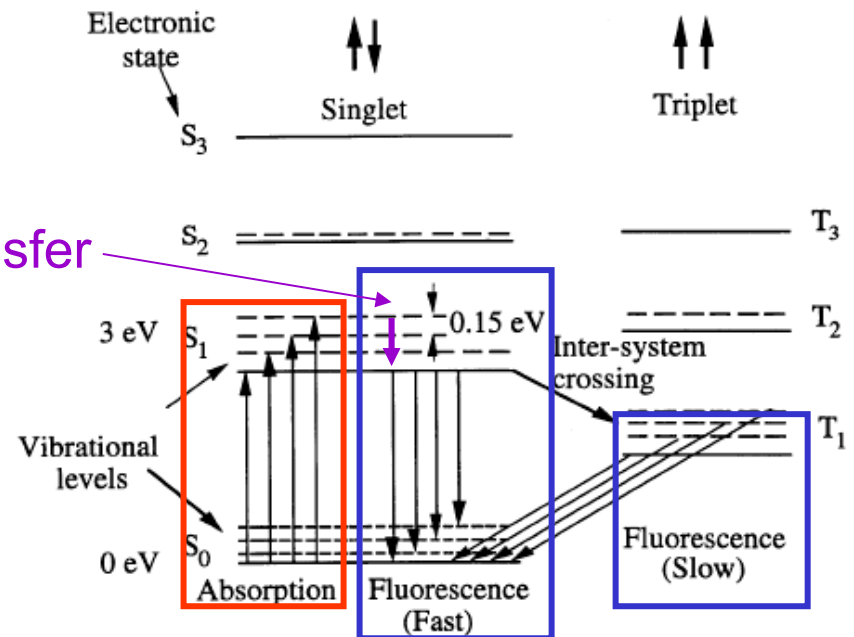
FIG. 2.3. Cross section for the photoelectric effect as a function of the  $Z$  value of the absorber. Data for 100 keV and 1 MeV  $\gamma$ s.

# Organic scintillators

Energy levels of organic scintillators: at room  $T_a$  ( $KT_a = 0.025$  eV) electrons on ground state, incident radiation excites electrons to  $S_1$  states, radiation-less decays to base  $S_1$  state, emission of light to  $S_0$ .  $S_1$  can decay to a triplet state with lower energy and longer decay time. The fluorescence UV light (250-370 nm) is absorbed by most organic materials, so the light signal can be extracted using a second fluorescent material (wavelength shifter) that converts UV in visible light (320-500 nm)

Non-radiative transfer

# photons produced  
 $\sim 2 \cdot 10^4 \text{ cm}^{-1}$  [for a MIP]  
 $\sim 10 \text{ eV / photon}$



# Time dependence of emitted light

Non-radiative transfer of energy from vibrational states to fluorescence state S1: **0.2-0.4 ns**

Decay of fluorescent state: **1-3 ns**

$$I(t) \propto 1 - e^{-t/\tau_r}$$

$\tau_r$  = rise time constant

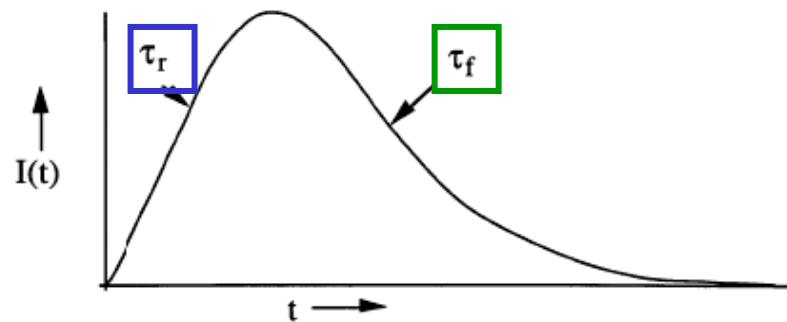
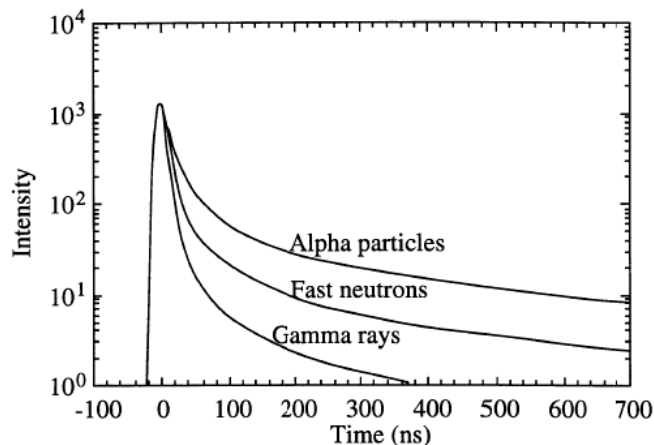
Fall with time constant  $\tau_f$

$$I(t) \propto e^{-t/\tau_f}$$

total pulse shape

$$I(t) = I_0 (e^{-t/\tau_f} - e^{-t/\tau_r})$$

Decay time in stilbene for various particles



The decay time depends on the ionization density

Material	State	$\lambda_{\max}$ [nm]	$\tau_f$ [ns]	$\rho$ [g/cm <sup>3</sup> ]	photons/MeV
Anthracene	crystal	447	30	1.25	$1.6 \cdot 10^4$
Pilot U	plastic	391	1.4	1.03	$1.0 \cdot 10^4$
NE104	plastic	406	1.8	1.03	$1.0 \cdot 10^4$
NE102	liquid	425	2.6	1.51	$1.2 \cdot 10^4$

# Birk's law

For organic scintillators the relation between emitted light and energy loss is not linear. Deviations from linearity are due to quenching interactions between excited molecules created along the ionizing particle path absorbing energy

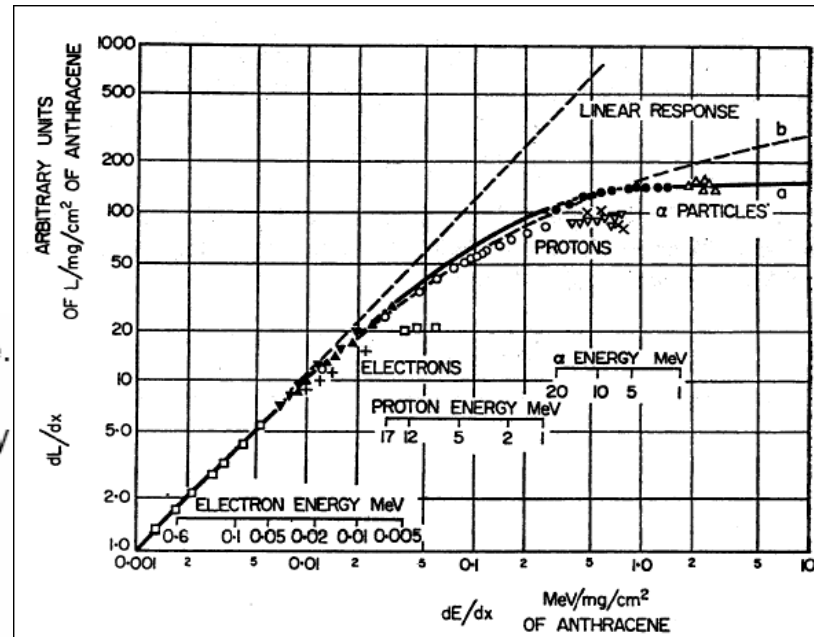
For an ideal scintillator and low ionization density Luminescence  $\propto$  Energy dissipated in scintillator

$$L = SE \quad \text{or} \quad \frac{dL}{dr} = S \left( \frac{dE}{dr} \right)$$

$$\frac{dL}{dr} = \frac{S \left( \frac{dE}{dr} \right)}{1 + kB \left( \frac{dE}{dr} \right)}$$

Density of ionized and excited molecules along track  
Quenching parameter

The light output depends on the ionization density



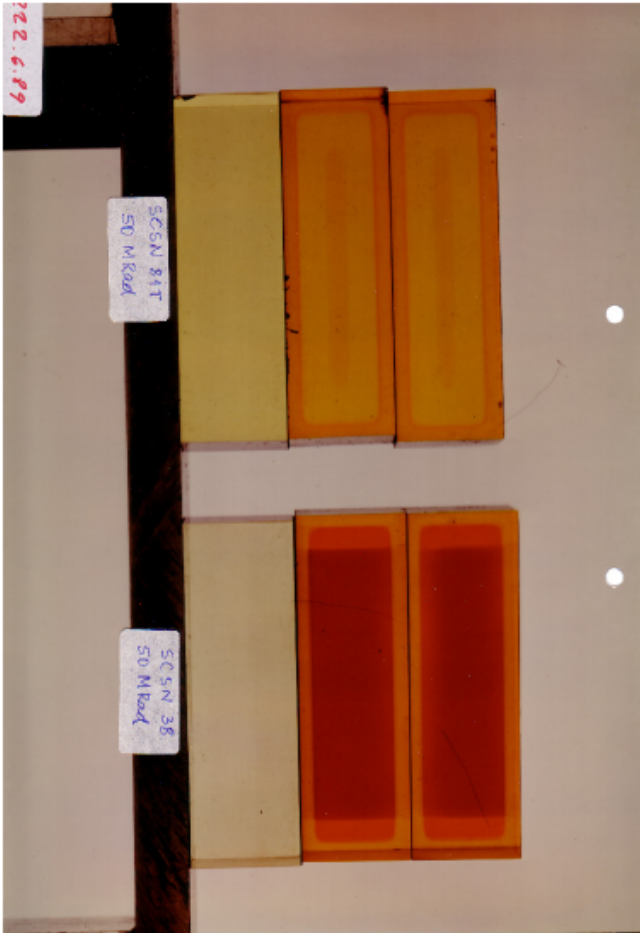
For small  $dE/dr$  this yields the luminescence yield postulated above.

For large  $dE/dr$  the specific luminescence saturates, as indicated by the data.

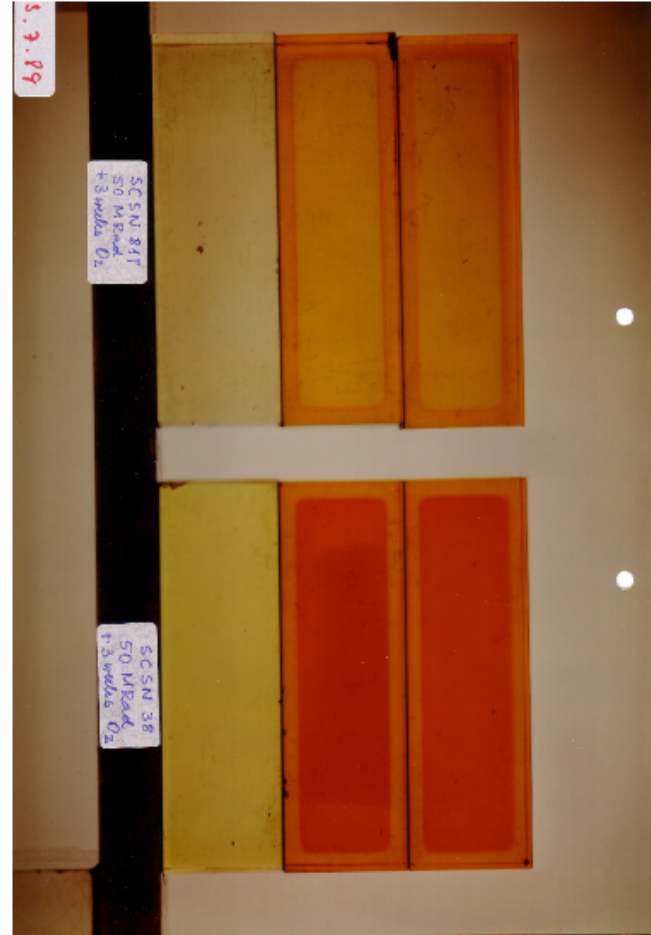
$$\frac{dL}{dr} = \frac{S}{kB} = const$$

# Radiation hardness in plastic

500 kGy irradiation of SCSN81T and SCSN38 from Kuraray

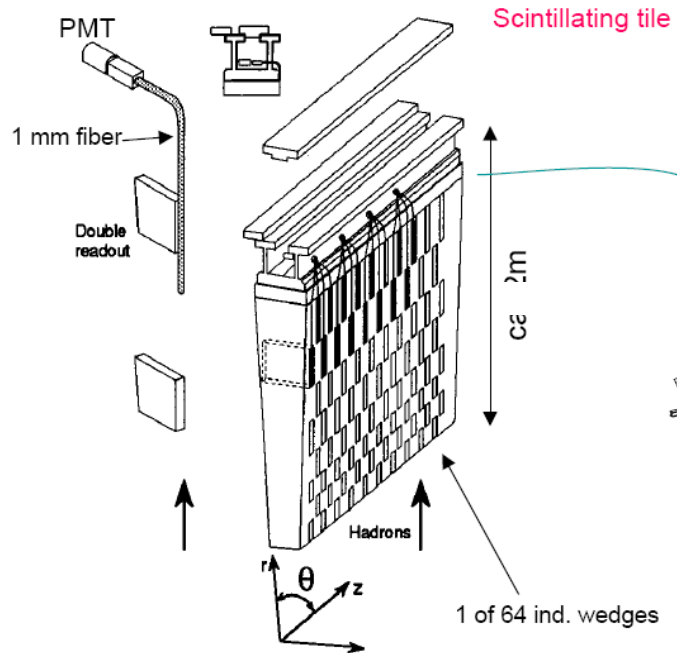


and after 3 weeks recovery in an oxygen-rich atmosphere



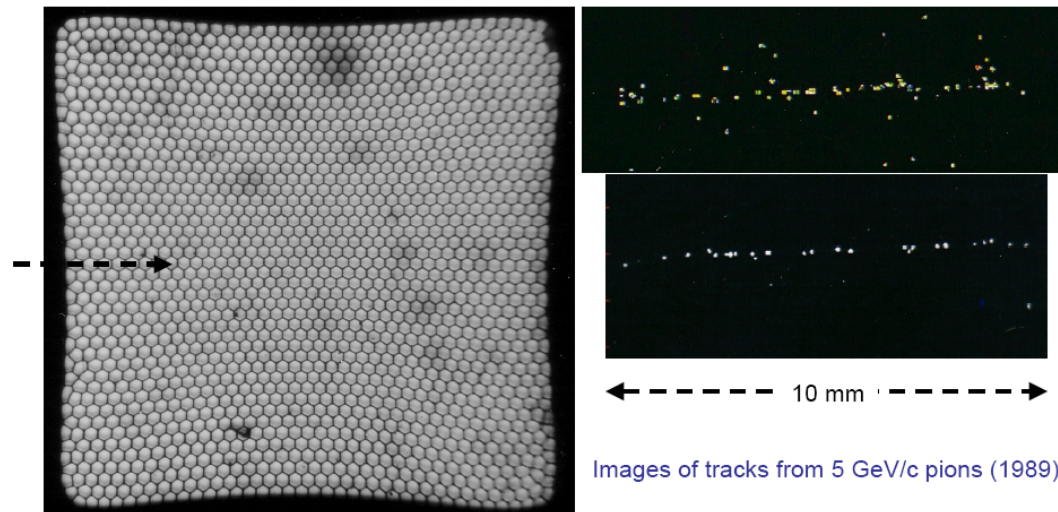
# Most common applications of organic scintillators

Large volume liquid or solid detectors (in form of tiles): underground experiments, sampling calorimeters (HCAL in CMS or ATLAS, etc.), counters, light guides.



High precision, small volume active targets and fibre tracking (UA2, D0, CHORUS)

RD7 development: bundles of hexagonal fibres (typ. 60  $\mu\text{m}$  dia., 2.5 mm bundle size) for tracking

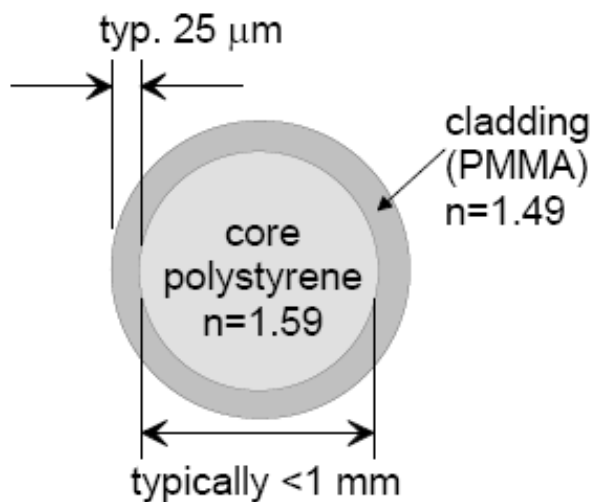


ATLAS hadron calorimeter

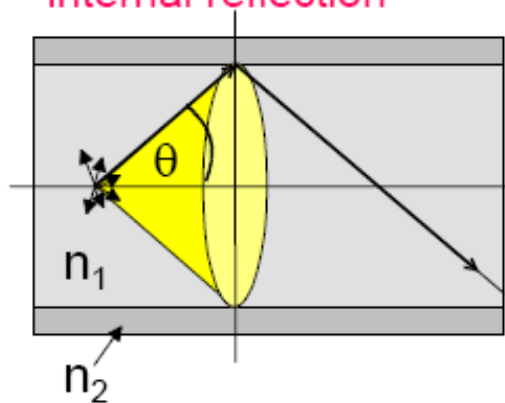
Images of tracks from 5 GeV/c pions (1989)

# Scintillator fibers

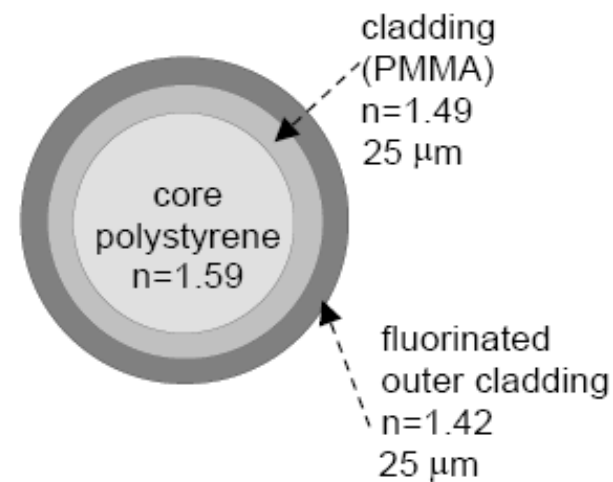
scintillating plastic fibre working principle:



light transport by total internal reflection



Double cladding system (developed by RD7)



$$\frac{d\Omega}{4\pi} = 0.5 (1 - \cos^2 \theta) = 3\%$$

$$\theta \leq \arccos \frac{n_2}{n_1} \approx 69.6^\circ$$

$$\frac{d\Omega}{4\pi} = 0.5 (1 - \cos^2 \theta) \approx 5.3\%$$



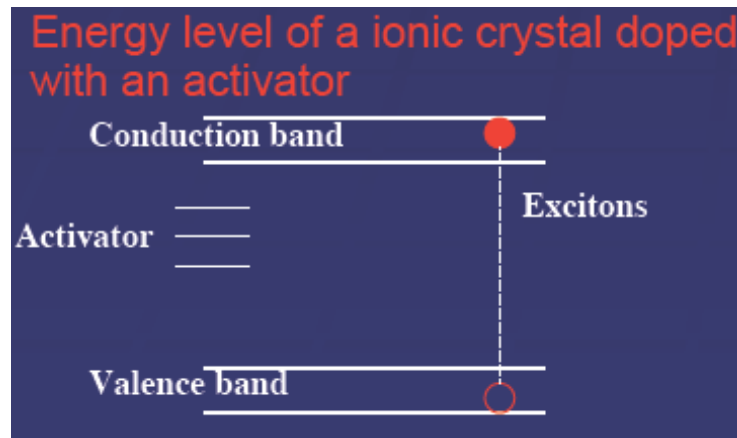
# Main R&D on organic scintillators

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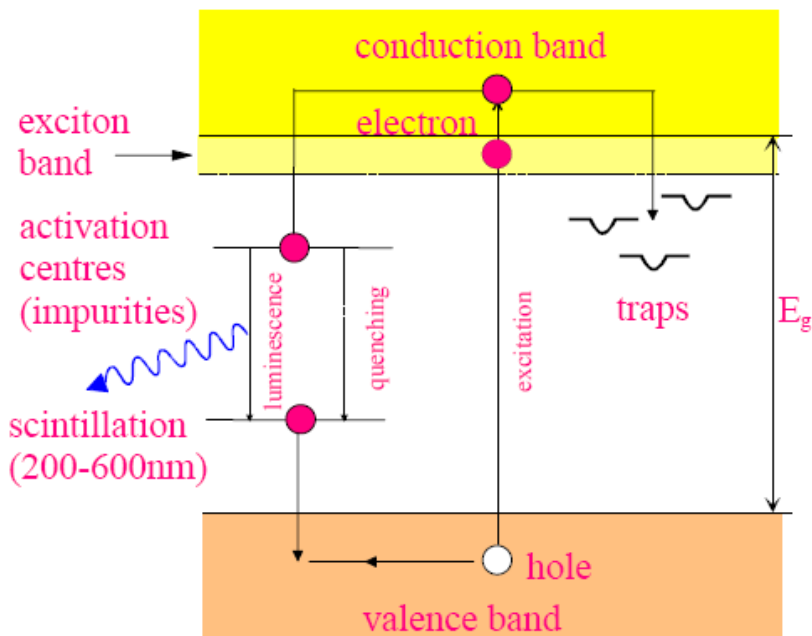
- **New dopants** with better light yield and larger Stokes shift
- **High granularity** readout of fibres
- Larger **attenuation lengths** in plastic fibres
- New **radiation hard** plastics to stand 100 kGy/year dose

# Inorganic scintillators

- Usually made of **high Z materials** → high density
- High Z enhances the **photoelectric interaction** contribution and high density increases the interaction efficiency.
- Crystals are grown in high temperature furnaces and are made of Alkali Halides (ie. NaI, CsI) or Oxides (eg. BGO).
- The crystalline structure creates energy bands between which electrons can jump → scintillation light
- Some crystals need activators to enable emission in the visible (Thallium in NaI(Tl))



# Inorganic scintillators II



Ionizing particles produce free electrons, holes and couples of electron-holes (excitons). These move around the crystal lattice until they meet an activation centre that they transform into an excited state  $A^*$  of energy  $E_1$  that can decay emitting light. **The decay time depends on the temperature as  $\exp[-E_1/(KT)]$**

**Warning**, sometimes  $\geq 2$  time constants

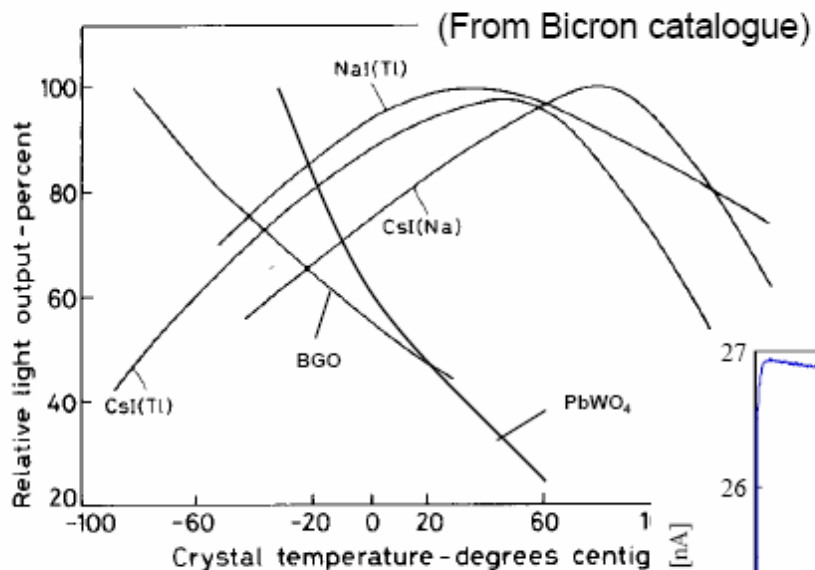
- fast recombination (ns- $\mu$ s) from activation centers
- delayed recombination due to trapping ( $\mu$ s-ms)

→ full control of growth, doping and impurities is imperative to optimize light yield, transmission and decay time

Material	Form	$\lambda_{\max}$ (nm)	$\tau_f$ (ns)	$\rho$ (g/cm <sup>3</sup> )	Photons per MeV
NaI(Tl) (20°C)	crystal	415	230	3.67	38,000
pure NaI (-196°C)	crystal	303	60	3.67	76,000
Bi <sub>4</sub> Ge <sub>3</sub> O <sub>12</sub> (20°C)	crystal	480	300	7.13	8,200
Bi <sub>4</sub> Ge <sub>3</sub> O <sub>12</sub> (-100°C)	crystal	480	2000	7.13	24,000
CsI(Na)	crystal	420	630	4.51	39,000
CsI(Tl)	crystal	540	800	4.51	60,000

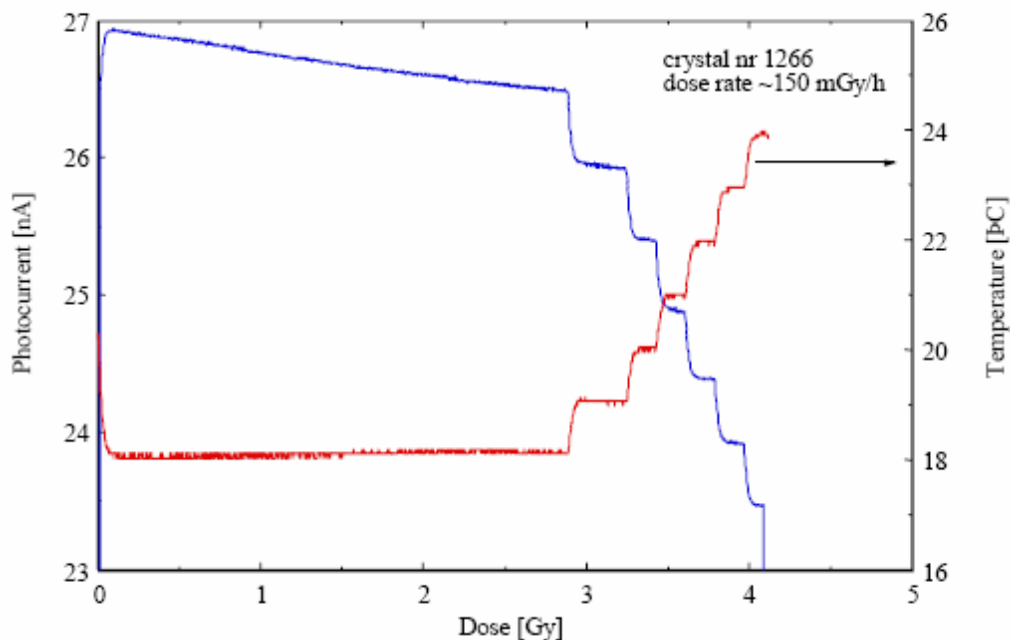
# Temperature dependence of crystals

Light output of crystals depends on temperature



Low dose irradiation of an **old** PbWO<sub>4</sub> crystal and check of the temperature dependence of its light yield (1996). The blue curve is the photocurrent generated by the irradiation and the red curve is the temperature of the sample.

The PbWO<sub>4</sub> crystal is used in CMS ECAL and its ~2% light yield decrease per °C asks for temp. control and monitoring



# For your reference ...

Scintillator composition	Density (g/cm <sup>3</sup> )	Index of refraction	Wavelength of max.Em. (nm)	Decay time Constant (μs)	Scinti Pulse height <sup>1)</sup>	Notes
Nal(Tl)	3.67	1.9	410	0.25	100	2)
CsI	4.51	1.8	310	0.01	6	3)
CsI(Tl)	4.51	1.8	565	1.0	45	3)
CaF <sub>2</sub> (Eu)	3.19	1.4	435	0.9	50	
BaF <sub>2</sub>	4.88	1.5	190/220 310	0,0006 0.63	5 15	
BGO	7.13	2.2	480	0.30	10	
CdWO <sub>4</sub>	7.90	2.3	540	5.0	40	
PbWO <sub>4</sub>	8.28	2.1	440	0.020	0.1	
CeF <sub>3</sub>	6.16	1.7	300 340	0.005 0.020	5	
GSO	6.71	1.9	430	0.060	40	
LSO	7	1.8	420	0.040	75	
YAP	5.50	1.9	370	0.030	70	

1) Relative to Nal(Tl) in %; 2) Hygroscopic; 3) Water soluble

# Most common applications of inorganic scintillators

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- Calorimetry
- X-ray and gamma spectroscopy
- Imaging
  - Positron Emission Tomography (PET) in medical imaging
  - Gamma Imaging (Anger camera)
- Monitoring in nuclear plants
- Oil wells, Mining, etc.

# Positron Emission Tomography

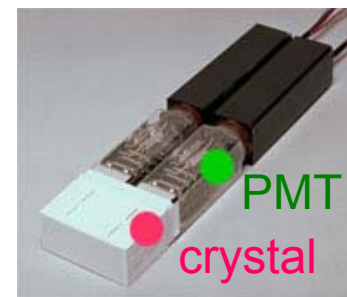
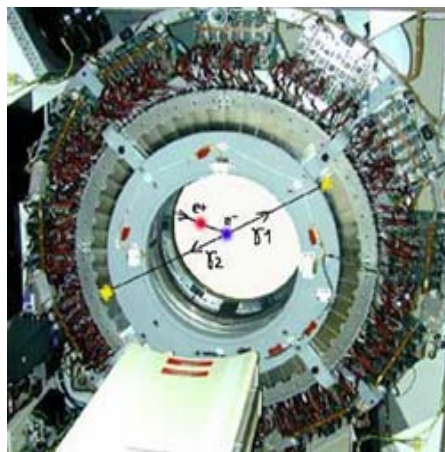
How can a calorimeter save your life?

→ PET

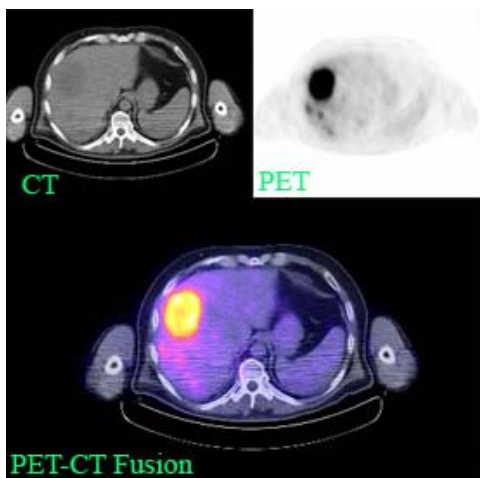
a commercial PET system for hospital treatment



the same system without cover doesn't it look like something familiar?



basic unit of a PET:  
crystal (LSO, BGO) + PMT

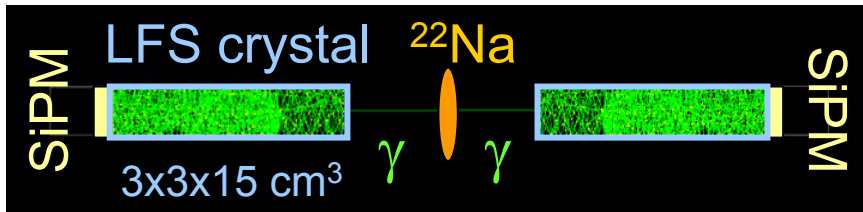
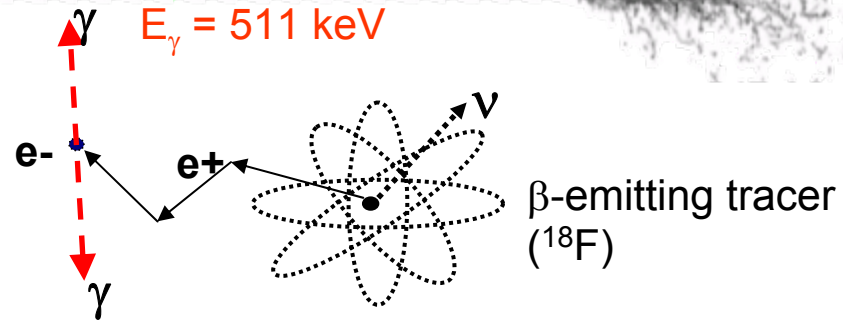


**PET scan** is a powerful tool for detecting several types of cancer. PET works by having the ability to detect sites of high metabolic activity. Since many cancers have significantly higher metabolism than normal tissues or noncancerous masses, PET allows sensitive detection of even small cancers.

**PET-CT Fusion** is a refinement of the technique that allows the most accurate correlation of anatomic information (from the CT) and metabolic information (from the PET scan) and helps to ensure the highest degree of accuracy for the exam.

# PET performance

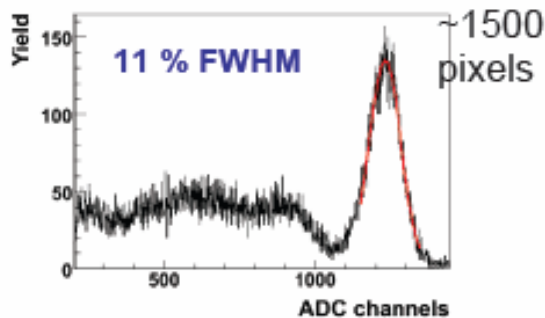
The physics process to detect:  
 $e^+e^- \rightarrow \gamma\gamma$  ( $E_\gamma = 511$  keV)



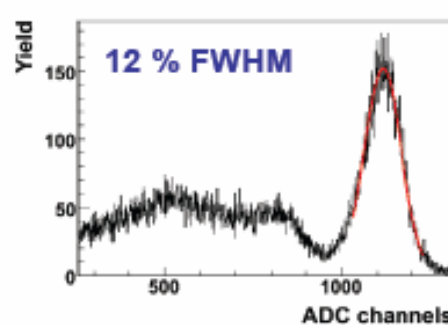
Simple system for prove of principle studies:  
 Two crystals in coincidence to detect back to back 511keV photons

Energy resolution:

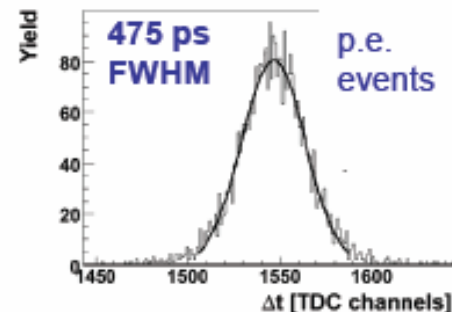
Channel 1:



Channel 2:



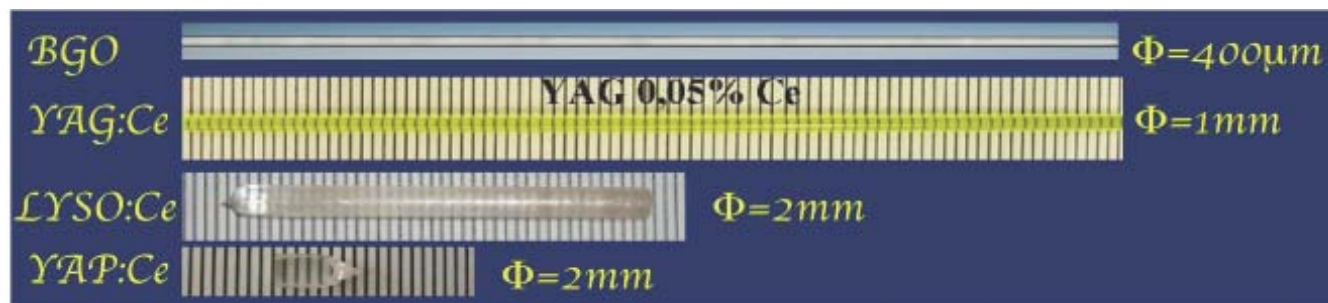
Coincidence time resolution:





# Main R&D on inorganic scintillators

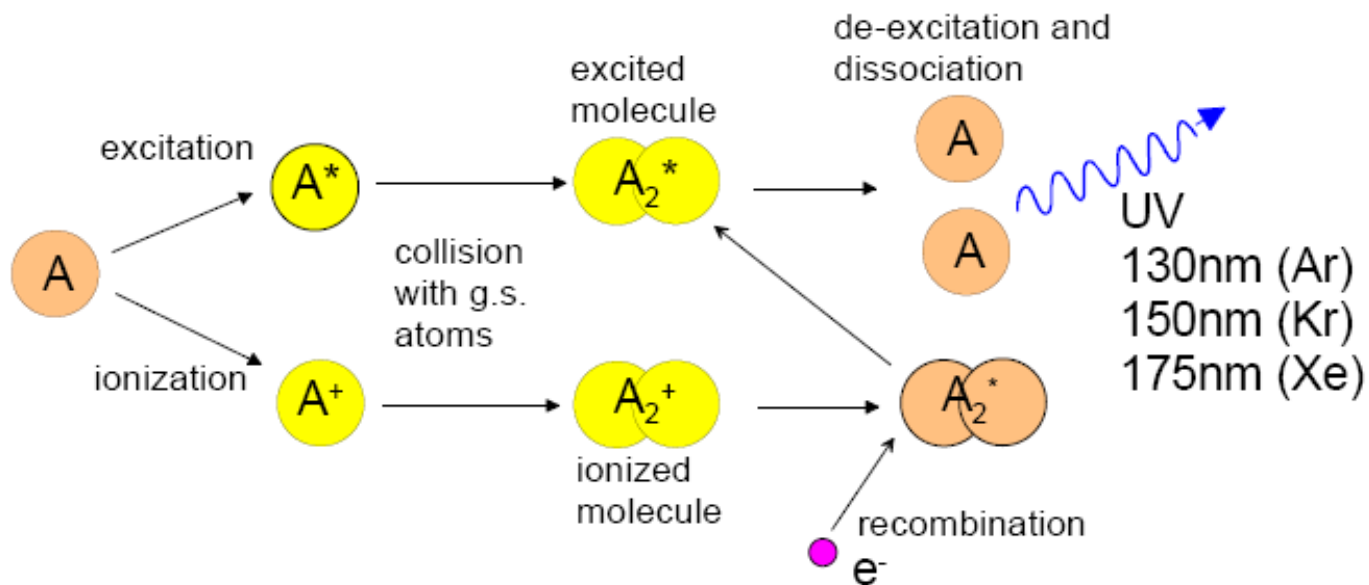
- Higher densities for higher Z (improve photoabsorption)
- High light yield (NaI(Tl) light yield still unchallenged)
- Short decay time (improve time resolution)
- Improve light coupling with photon detector
- More radiation hard
- Inexpensive, “easy” to manufacture, reproducible
- Large size, easy handling and “machinable”
- Smaller size → fiber crystals



The micro-pulling down crystal growth technology (Courtesy Fibercryst)

# Liquefied noble gases

Liquefied noble gases: LAr, LXe, LKr



Also here one finds 2 time constants: from a few ns to 1  $\mu$ s.

# Cherenkov effect

A charged particle traveling in a dielectric medium with  $n > 1$  radiates **Cherenkov light** if its velocity is larger than the phase velocity of light  $v > c/n$  or  $\beta > 1/n$  (**threshold**)

The emission is due to an asymmetric polarization of the medium in front and at the rear of the particle, giving rise to a varying electric dipole momentum.

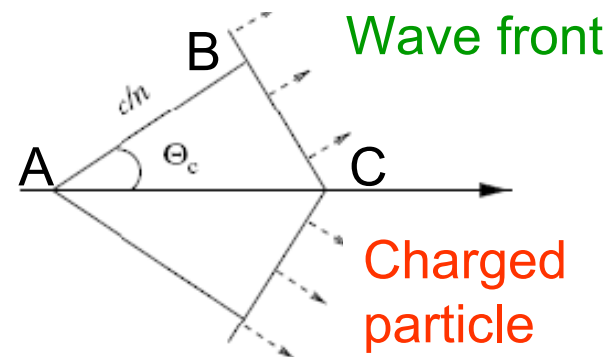
Some of the particle energy is converted into light. A coherent wave front is generated moving at velocity  $v$  at an angle  $\Theta_c$ .

If the media is transparent the Cherenkov light can be detected.

If the particle is ultra-relativistic ( $\beta \sim 1$ )  $\Theta_c = \text{const}$  and has max value

$$\cos \Theta_c = \frac{AB}{AC} = \frac{c}{n} t \cdot \frac{1}{\beta c t} = \frac{1}{\beta n}$$

In water  $\Theta_c = 43^\circ$ , in ice  $41^\circ$



# Cherenkov effect

The intensity of the **Cherenkov radiation** (number of photons per unit length of particle path and per unit of wave length) **depends on charge and velocity** of particle

$$\frac{d^2 N}{dx d\lambda} = \frac{4\pi^2 z^2 e^2}{hc\lambda^2} \left(1 - \frac{1}{n\beta^2}\right) = \frac{2\pi z^2}{\lambda^2} \alpha \sin^2 \Theta_c$$

$$\alpha = \frac{2\pi e^2}{hc}$$

In the range 400-700 nm:

$$\frac{dN_\gamma}{dx} = \int_{\lambda_1}^{\lambda_2} d\lambda \frac{d^2 N_\gamma}{dx d\lambda} = 2\pi z^2 \alpha \sin^2 \Theta_c \int_{\lambda_1}^{\lambda_2} \frac{d\lambda}{\lambda^2} = 2\pi z^2 \alpha \sin^2 \Theta_c \left(\frac{1}{\lambda_1} - \frac{1}{\lambda_2}\right) = 490 z^2 \sin^2 \Theta_c \text{ photons/cm}$$

for electrons in quartz  $z=1$ ,  
at  $\Theta_c=45^\circ$   $\sin^2 \Theta_c=0.5$

$$E_{\text{thr}} = m_{\text{particle}} \gamma_{\text{thr}} = m_{\text{particle}} \frac{1}{\sqrt{1 - \beta_{\text{thr}}^2}}$$

for quartz  $E_{\text{th}}(e) > 0.5 \text{ MeV}$

in water energy loss by Ch. is about  $10^4$  less than by ionization (ion: 2 MeV/cm)

→ **but directional effect!!!**

$E_{\text{th}}(e) > 0.26 \text{ MeV}$  and  $E_{\text{th}}(\mu) > 54 \text{ MeV}$  (neutrino exper.)

# PID with a threshold cherenkov detector

The Cherenkov threshold can be used to separate particles of momentum  $p$  and masses  $m_1$  and  $m_2 > m_1$ .

The radiator medium can be chosen such that the heavier particle is just below threshold:  $\beta_2 \approx 1/n$

Calculate the number of produced photo-electrons / cm of a given Cherenkov radiator material by a 1 GeV pion when discriminating  $K/\pi$  (assume photo-detector efficiency = 20%)

# Solution

Choose a radiator medium such that  $K=m_2$  is below threshold:

$\beta_2 \approx 1/n$  while  $\beta_1 > 1/n$

$$\sin^2 \Theta_c = 1 - \frac{1}{\beta_1^2 n^2} \approx 1 - \frac{\beta_2^2}{\beta_1^2} \approx 1 - \frac{\beta_2^2}{\beta_1^2} = 1 - \frac{E_1^2}{E_2^2} \approx \frac{c^2(m_2^2 - m_1^2)}{p^2}$$

$$p \gg m_2$$

For 20% PDE one gets:

$$N/L = 100 \sin^2 \Theta_c = 100 \frac{c^2(m_2^2 - m_1^2)}{p^2} \quad [\# \text{ photoelectrons / cm}]$$

$$m_2 = m_K = 493.7 \text{ MeV}$$

$$m_1 = m_\pi = 139.6 \text{ MeV}$$

for  $K/\pi$  separation at 1 GeV  $N_{pe}/L \approx 22 \text{ pe/cm}$  for  $m_1 = m_\pi$   
and by design 0 for  $m_1 = m_K$



# The measurement of showers

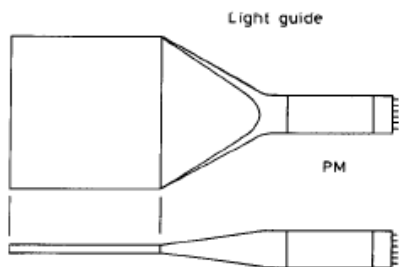
or “from energy to signal”

Step 1: Convert energy to light

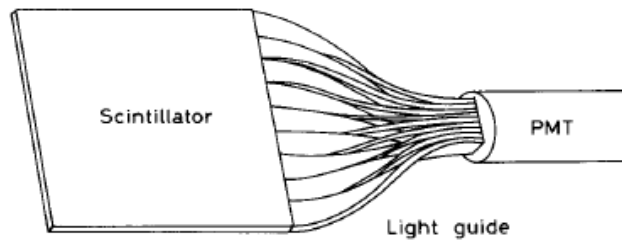
**Step 2: Convert light to electrical signal**

Step 3: Reading an electrical signal

# Reading out light from a scintillator



“fish tail”



adiabatic

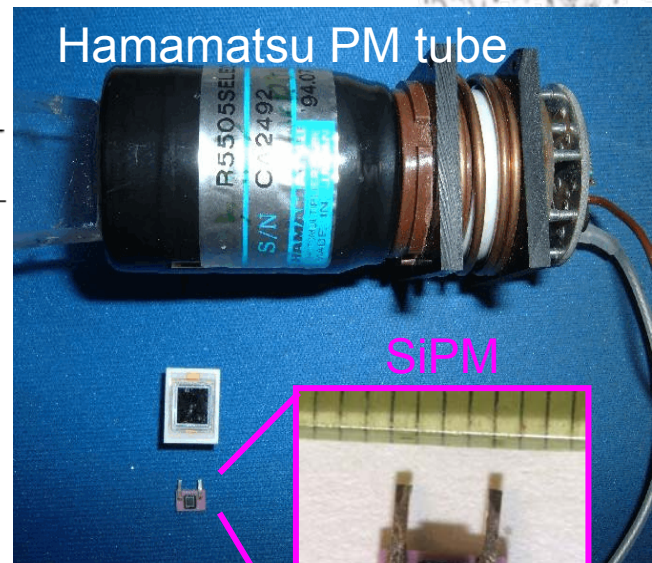
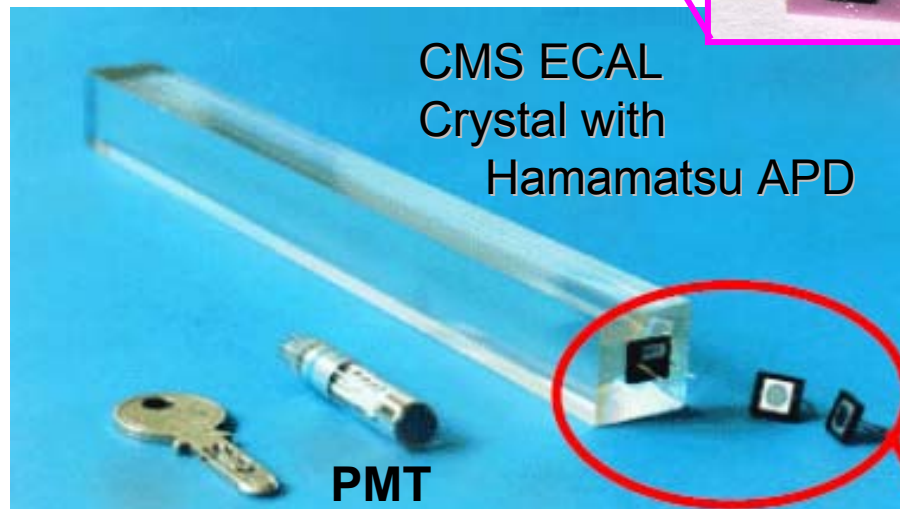


Photo-multiplier tubes are the past!

Be modern: go silicon!



PMT

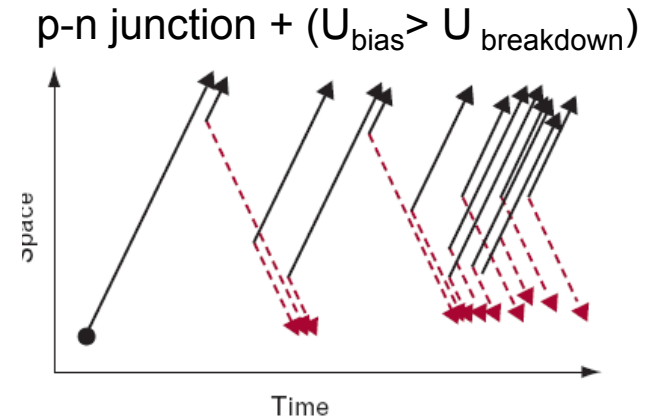
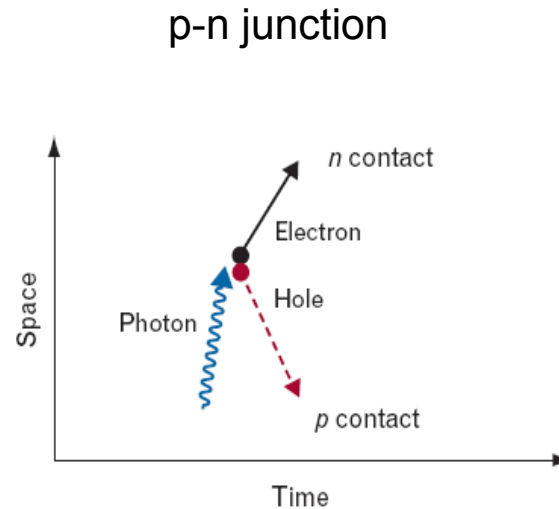
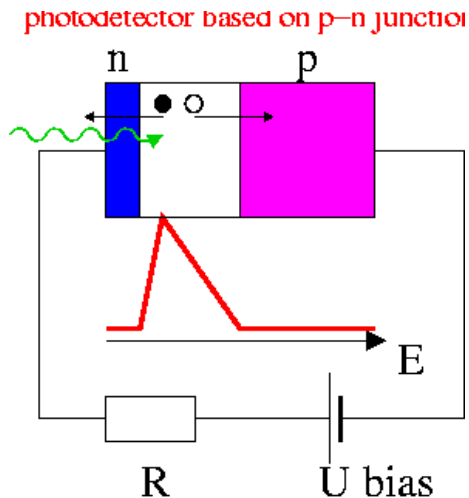


# New photo-detectors

Main drawbacks of PMTs: bulky shape, the high price and the sensitivity to magnetic fields.

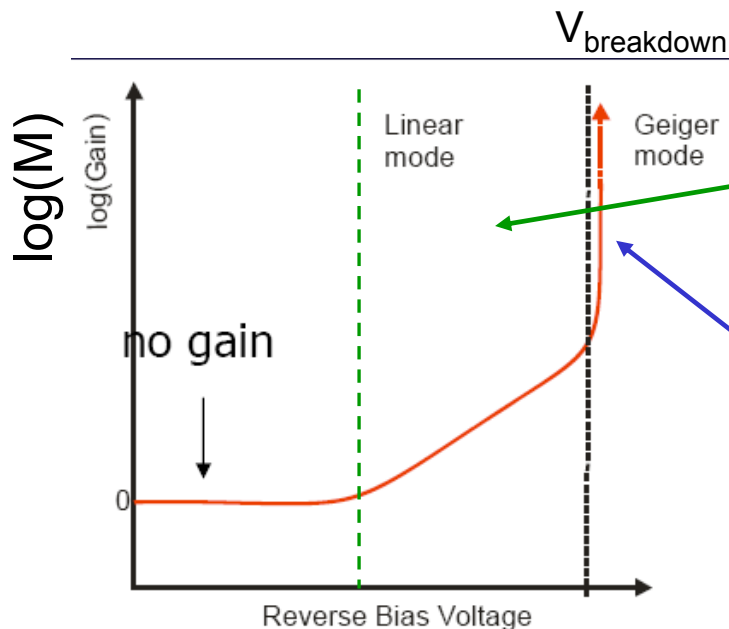
Photodiodes are **semiconductors light sensors** that generate a current or voltage when light illuminates the p-n junction.

→ Allow detection of light in 200-1150 nm



# Photon detection in Semiconductors

- PIN photodiodes are Near IR detectors (InGaAs or Si) have excellent linearity with light, high speed, low noise and dark current but major drawback: low gain  $\sim 1$
- Avalanche Si photodiodes (APDs) have internal gain so are more sensitive and have much higher QE  $> 70\%$ .



## Avalanche Photo-Diode (APD)

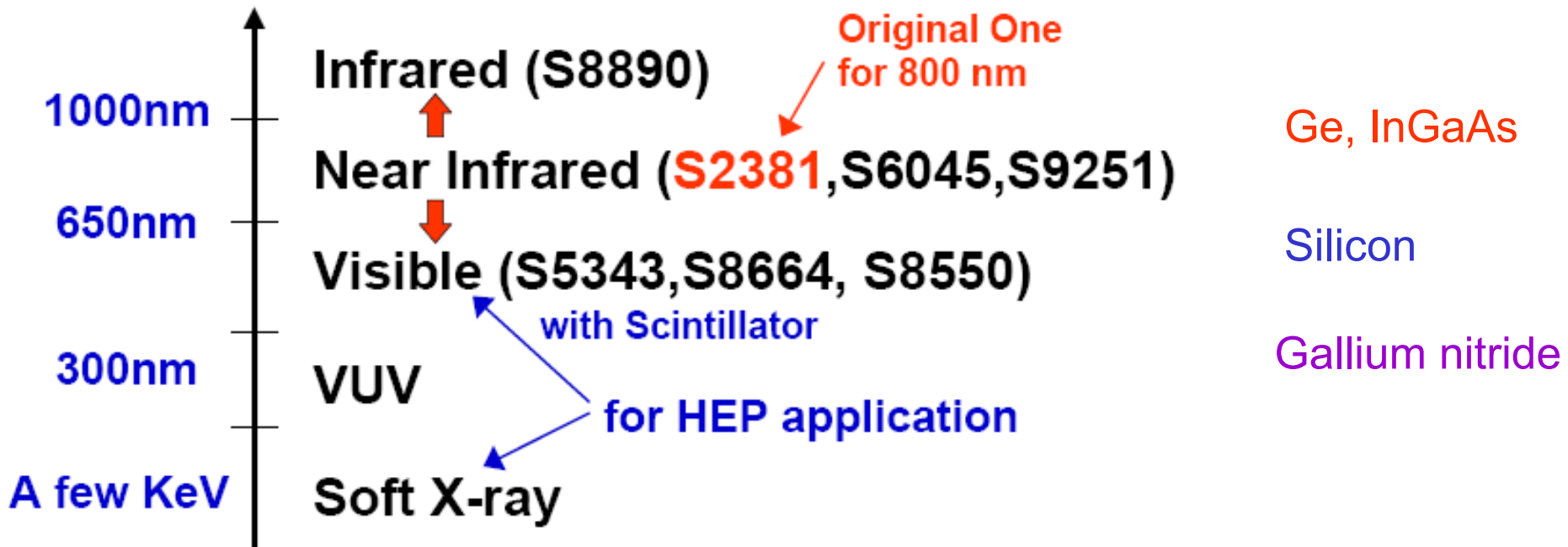
operated in linear mode  $\sim$  AMPLIFIER (Gain  $M \sim 50-500$ )  
!! signal proportional to energy deposited  
 $\rightarrow$  used in CMS ECAL

## Geiger mode Avalanche Photo-Diode

operate above breakdown voltage (Gain  $M \sim 1 \cdot 10^6$ )  
!! It's a BINARY device  
 $\rightarrow$  for practical application use ARRAY of single Geiger mode Avalanche Photo-Diodes:  
the Silicon Photo-Multiplier

# APD for various applications

Pick the right one for your application:



# Excess noise

Noise due to the multiplication process at a gain,  $M$ , denoted by  $F(M)$  often expressed as:

$$F(M) = \kappa M + \left(2 - \frac{1}{M}\right)(1 - \kappa)$$

For an electron multiplication device  $\kappa$  is given by the hole impact ionization rate divided by the electron impact ionization rate.

It is desirable to have a large asymmetry between these rates, in order to minimize  $F(M)$  since  $F(M)$  is one of the main factors which limit the energy resolution obtainable.

Typical values of  $\kappa$  are: 0.02-0.06 for Si, 0.9 for Ge, 0.45 for InGaAs

Typical values of  $F(M)$  are: 5-8 for Si, 9 for Ge, 5 for InGaAs

# Geiger-mode APD

In principle an APD could be operated in a supercritical state above breakdown and could detect single photons but it would not stay long in this state and the recovery time would be much longer than the time between consecutive generation of free carriers (dark counts).

Way out:

Subdivide the APD into many cells and connect them all in parallel via an individual limiting (quenching) resistor.

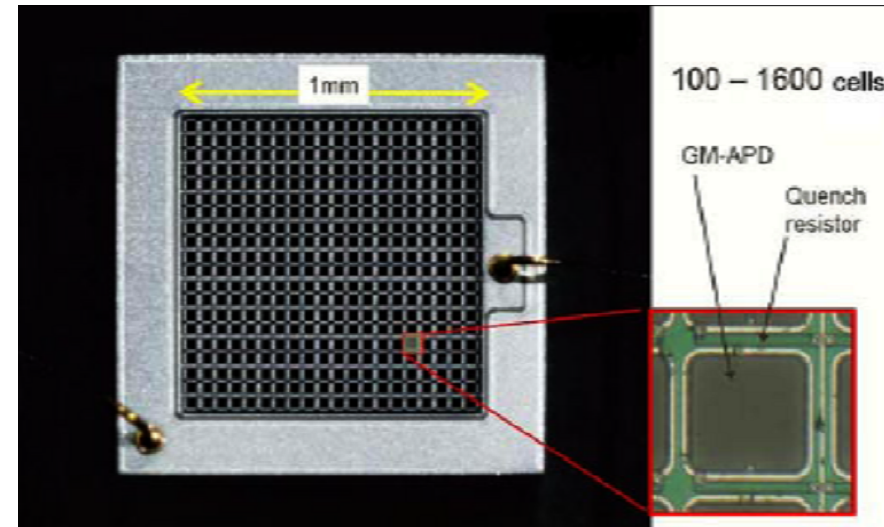
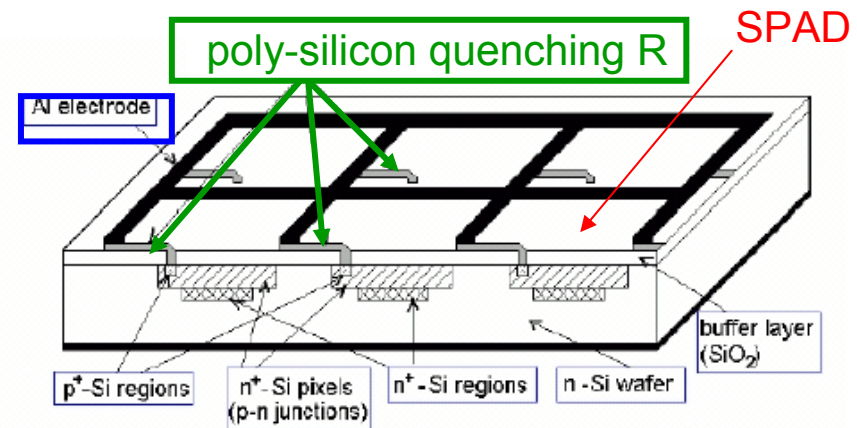
→ typically 100-1000 pixels / mm<sup>2</sup>

Some typical pixel parameter:

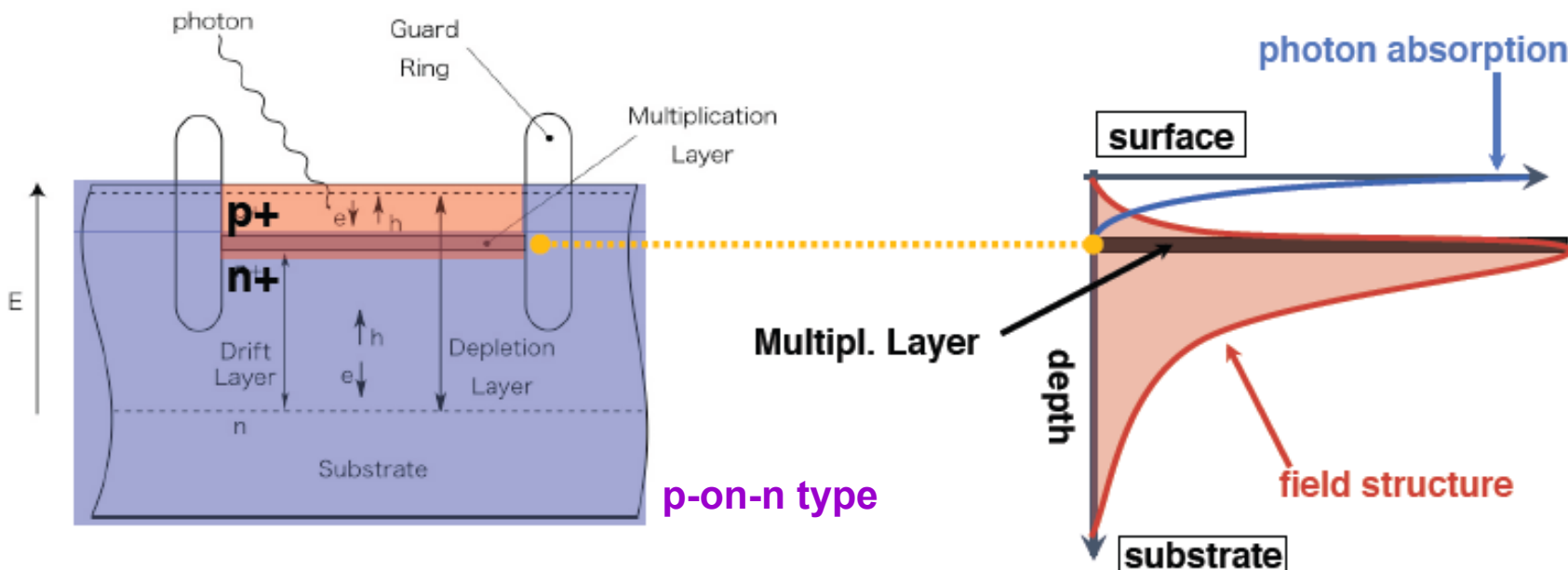
-pixel size ~20-30μm

-pixel capacitance  $C_{\text{pixel}} \sim 50\text{fmF}$

-quenching resistor  $R_{\text{pixel}} \sim 1-10\text{M}\Omega$



# Working principle



- small depletion region  $\sim 2\mu\text{m}$
- strong electric field  $(2-3)\times 10^5 \text{ V/cm}$
- carrier drift velocity  $\sim 10^7 \text{ cm/s}$

-very short Geiger discharge development  $< 500 \text{ ps}$

Photoelectric conversions occur **above** the multiplication layer

→ **electrons** drift to the multiplication layer

Random excitations occur mainly **below** the multiplication layer

→ **holes** drift to the multiplication layer

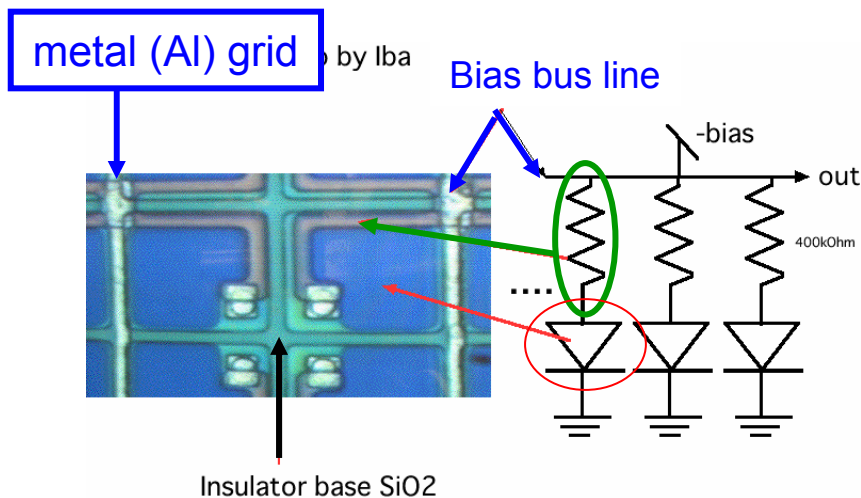
# High gain

G-APDs produce a standard signal when any of the cells goes to breakdown. The amplitude  $A_i$  is proportional to the capacitance of the cell divided by the electron charge times the overvoltage:

$$A_i \sim C/q \cdot (V - V_b) \quad (V - V_b) \text{ we call "overvoltage"}$$

$V$  is the operating bias voltage and  $V_b$  is the breakdown voltage

When many cells are fired at the same time, the output is the sum of the standard pulses.



all pixels connected in parallel  
only one signal line

→ output  $A = \sum A_i$  (individual pixel signals)  
sometimes non-linear because of crosstalk

typical Bias voltage  $\sim 2$  V above breakdown

# SiPM properties

G-APDs behave like PMTs, thereby the name: **Silicon Photomultiplier (SiPM)**

The **gain is in the range of  $10^5$  to  $10^7$** . Single photons produce a signal of several mV on a 50 Ohm load. No or at most a simple amplifier is needed!

Pickup noise is no more a concern (no shielding).

There is no nuclear counter effect – even a heavily ionizing particle produces a signal which is not bigger than that of a photon.

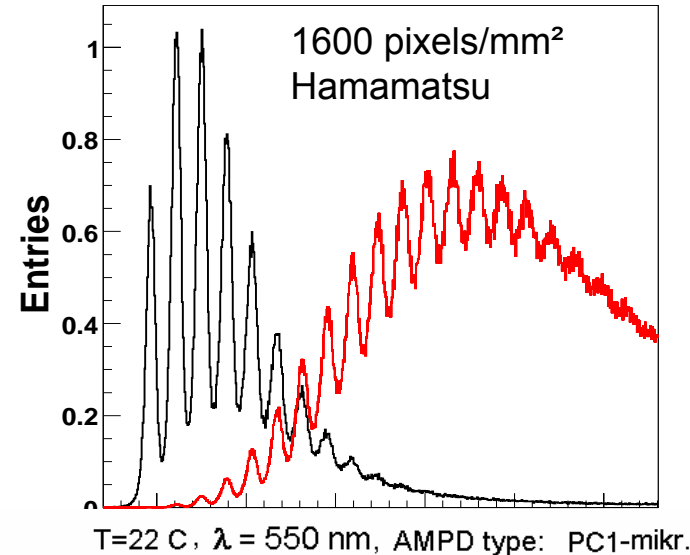
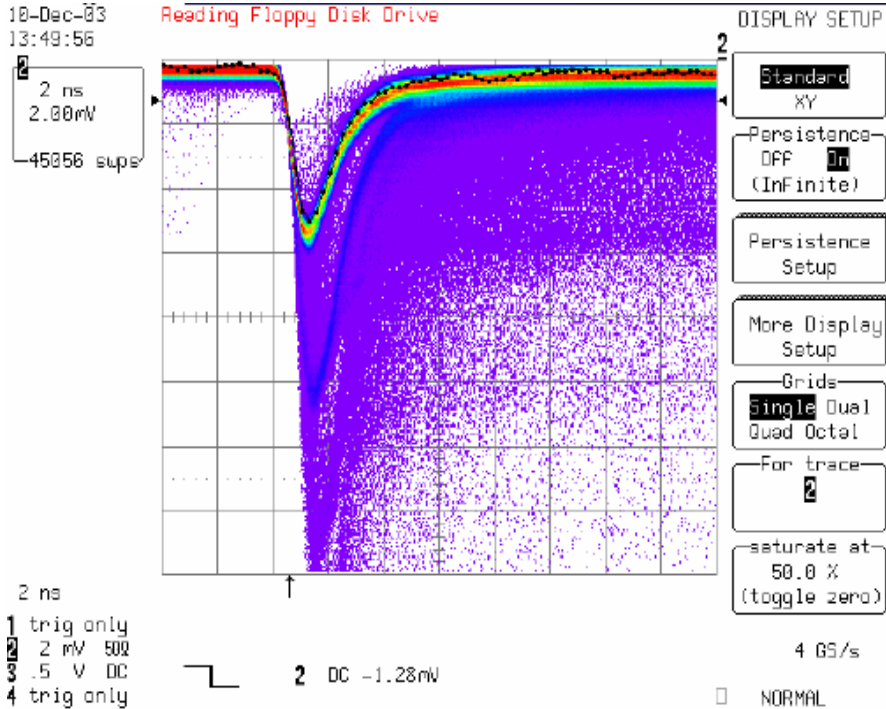
Since there are no avalanche fluctuations (as we have in normal APDs) the **excess noise factor is very small**, could eventually be one.

Grooms theorem (the resolution of an assembly of a scintillator and a semiconductor photodetector is independent of the area of the detector Nucl. Instr. and Meth. 219 (1984) 141) is no more valid.



# SiPM properties: single pixel resolution

SiPM output is the analog sum of all pixel signals

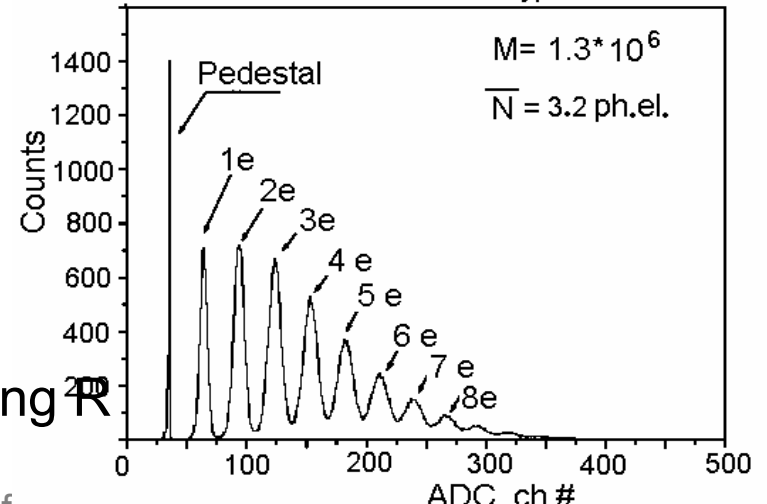


high gain → pixel signal visible on scope

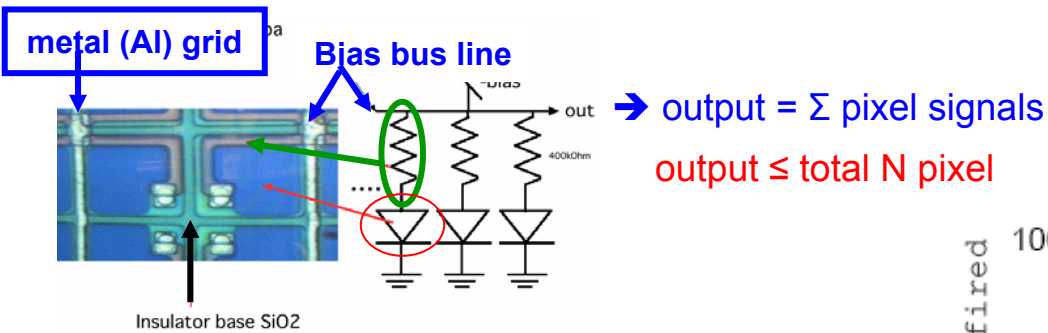
- signal rise time < 1 ns
- fast fall ~ 5-10 ns

recovery time tunable by choice of quenching R

$$\tau \sim R_{\text{pixel}} C_{\text{pixel}} \sim 20 - 500 \text{ ns}$$



# Dynamic range



Dynamic range naturally limited by number of pixels

Optimal working condition:  
number of photo-electrons < N pixels

from probability considerations:

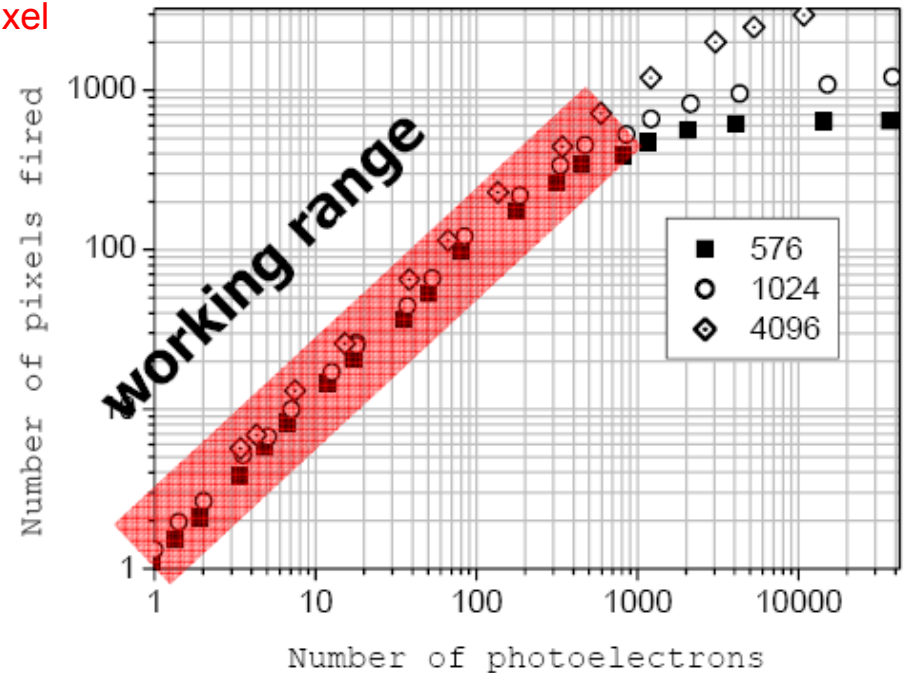
$$A \approx N_{\text{pixels}} = N_{\text{total}} \cdot \left[ 1 - e^{-\frac{N_{\text{photon}} \cdot \text{PDE}}{N_{\text{total}}}} \right]$$



~20% deviation from linearity if  
50% of pixels are fired

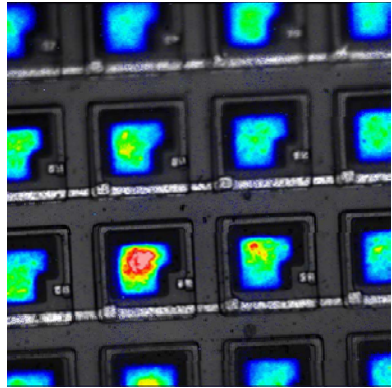
2 or more photons in one cell look exactly like 1 single photon

Measured on SiPM MEPHI/Pulsar, from B. Dolgoshein



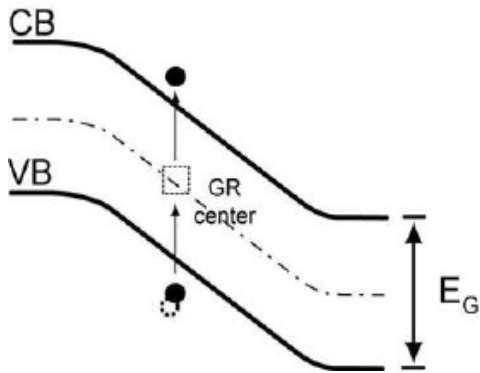
# Dark count rate

Only the very basics:



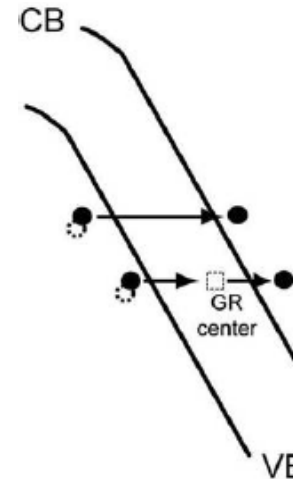
a SiPM pixel can be fired by an incoming photon but free carriers can be generated also by **thermal effects** or **tunneling** (field-assisted generation)

**these** lead to a dark count rate of 100 kHz – 10 MHz / mm<sup>2</sup> (@25°C) with threshold at half of one photo-electron amplitude ( $\sim 0.5 \times 10^6$ )



Free carrier generation  
by thermal effects

Depends on temperature  
(can be cooled away)



Tunneling

Depends on operation voltage (E field)  
Influenced by technological design

# Dark count rate

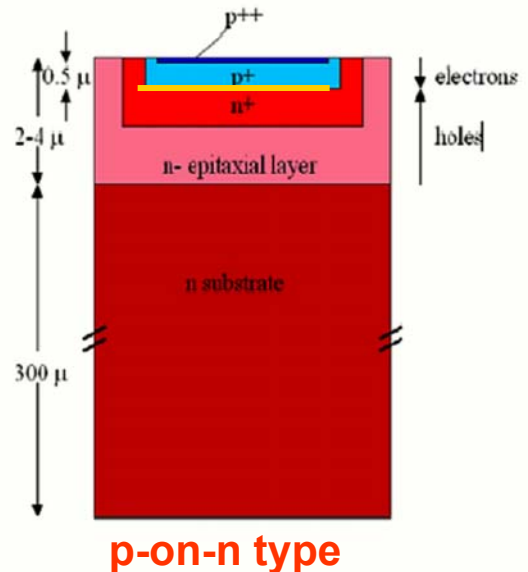
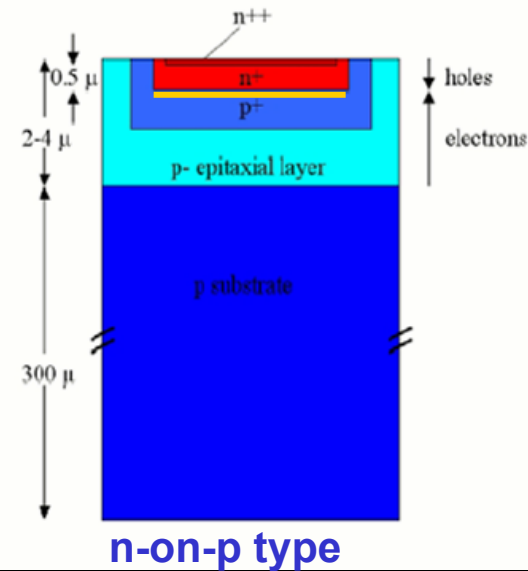
In first order the thermal generation of carriers is **proportional to the depleted volume** which is for every cell the area times the thickness of all the layers on top of the low ohmic substrate.

In the n-on-p type layers the electrons (in the p-on-n type layers the holes) drift towards the high field region of the junction.

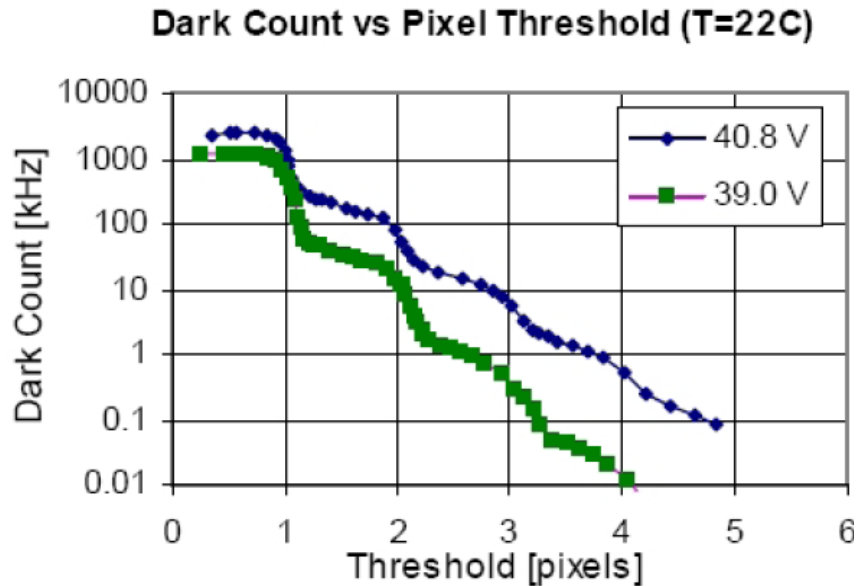
**The electrons will trigger there a breakdown with higher probability than the holes.**

In a G-APD with a p-on-n type substrate the volume of the p-layer is much smaller than in the G-APD on a n-on-p type substrate.

**Devices with a p-on-n-type substrate show smaller dark count rates.**

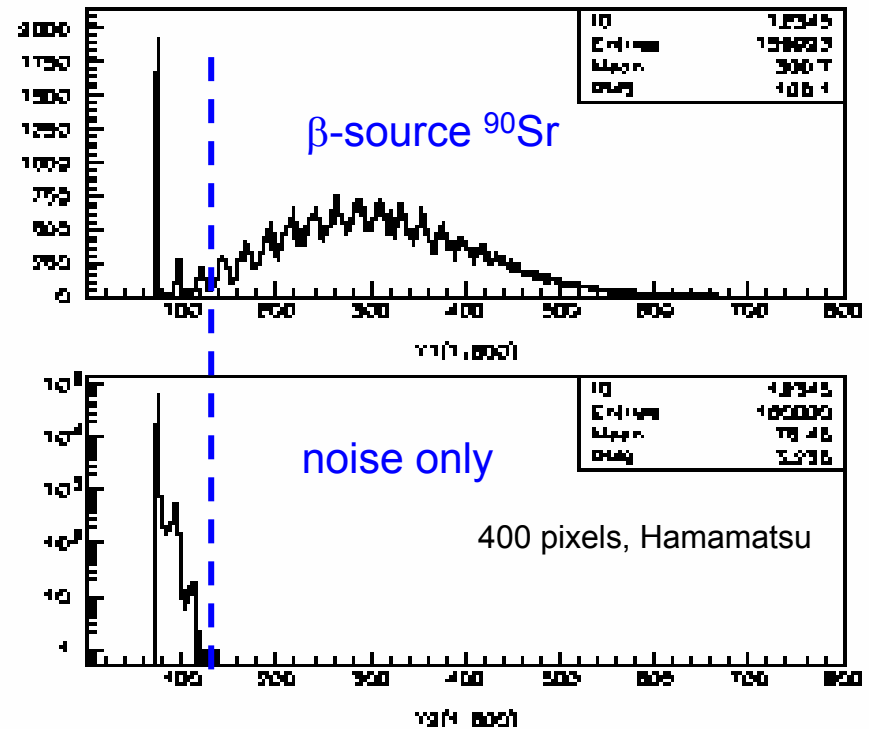


# Dark count rate



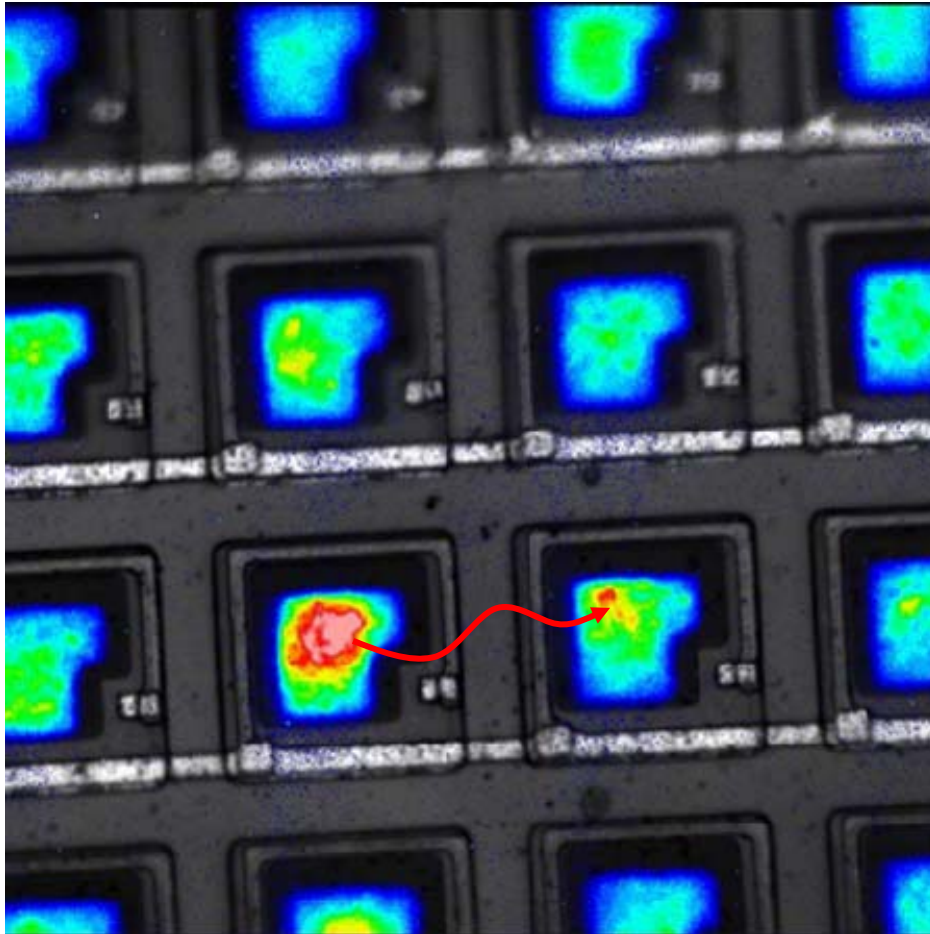
dark count rate > 0.5 pixel ~ MHz  
decrease rapidly with threshold

→ what is the relevant threshold  
for physics?



Blue sensitive SiPM directly coupled on  
3x3x0.5 cm<sup>3</sup> scintillator tile

# Pixel: a closer look

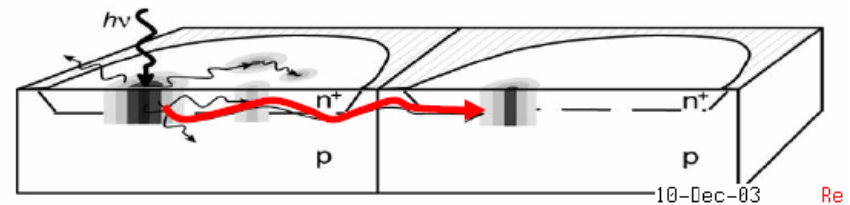


Emission microscopy picture, MPI

Optical inter-pixel cross talk:

during Geiger avalanche  
 $\sim 3$  emitted  $\gamma$  /  $10^5$  carriers  
with  $E_\gamma > 1.14$  eV

*A. Lacaita et al, IEEE TED (1993)*



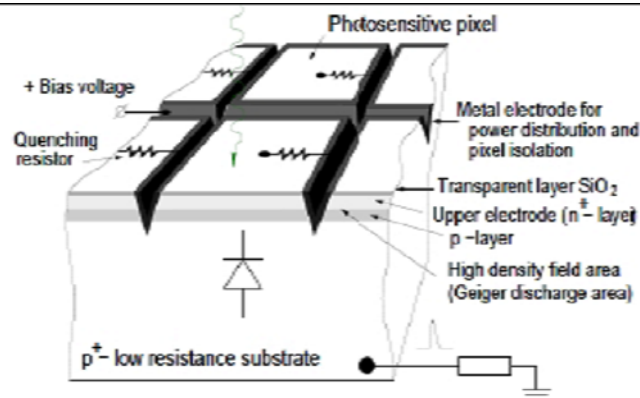
➔ Leads to artificial increase of signal

# Suppress optical cross talk

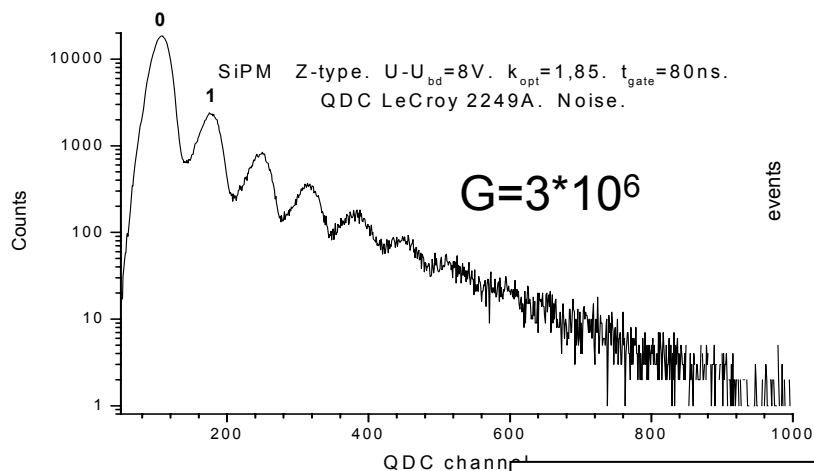
Possible counter measures:

- lower  $V_{Bias}$   $\rightarrow$  lower breakdown probability (lower PDE)
- optical insulation between pixels  $\rightarrow$
- technological modifications: i.e. smaller  $C_{pixel}$

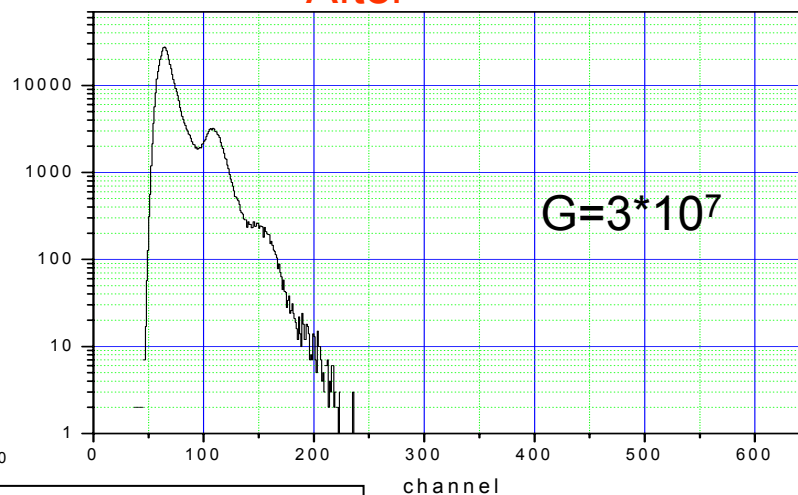
Inter-pixel trench to block  $\gamma$



Before



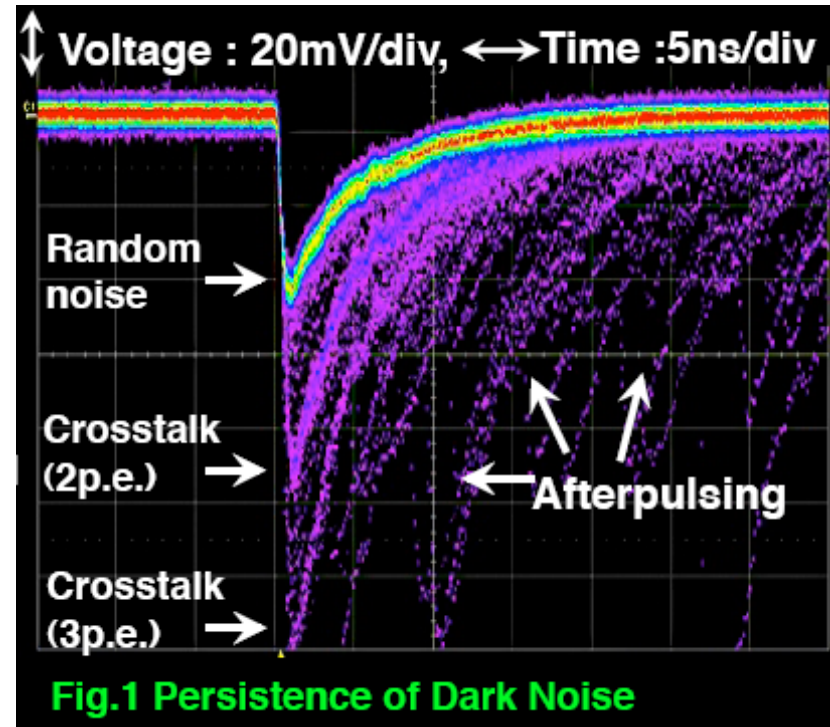
After



# After pulse

In the silicon volume, where a breakdown happened, a plasma with high temperatures (few thousand degree C) is formed and deep lying traps in the silicon are filled. Carrier trapping and delayed release causes afterpulses during a period of **several 100 ns after a breakdown**.

The probability for afterpulses increases with higher overvoltage (higher gain).



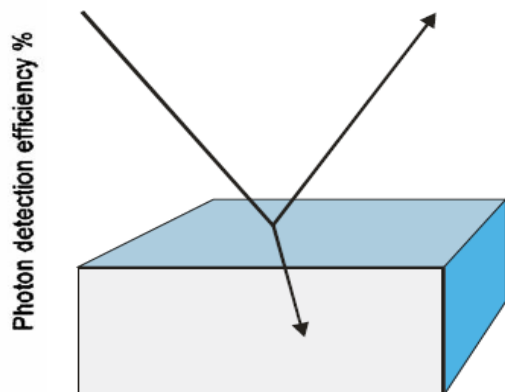


# Photon Detection Efficiency (PDE)

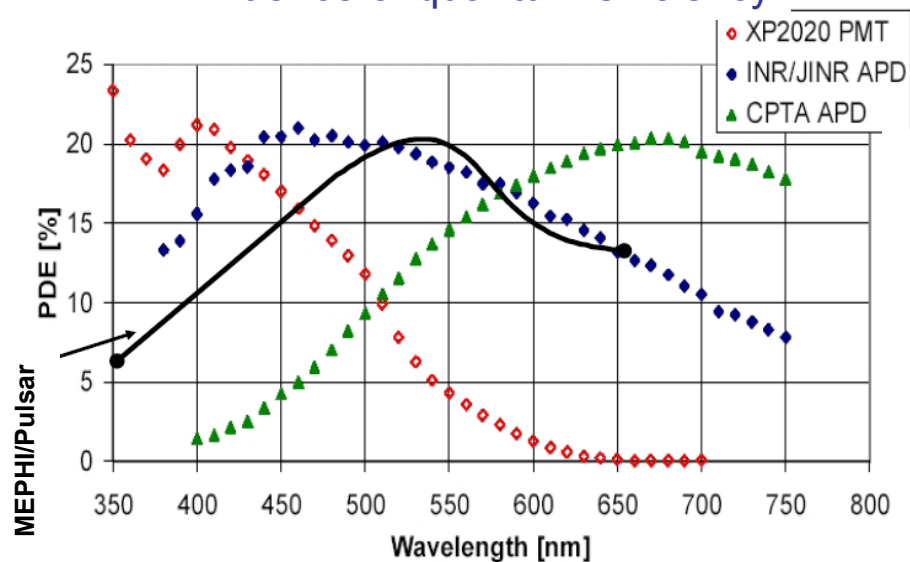
PDE is the product of mainly three components:

$$\text{PDE} = \text{QE} \cdot \varepsilon \cdot P_{\text{Geiger}}$$

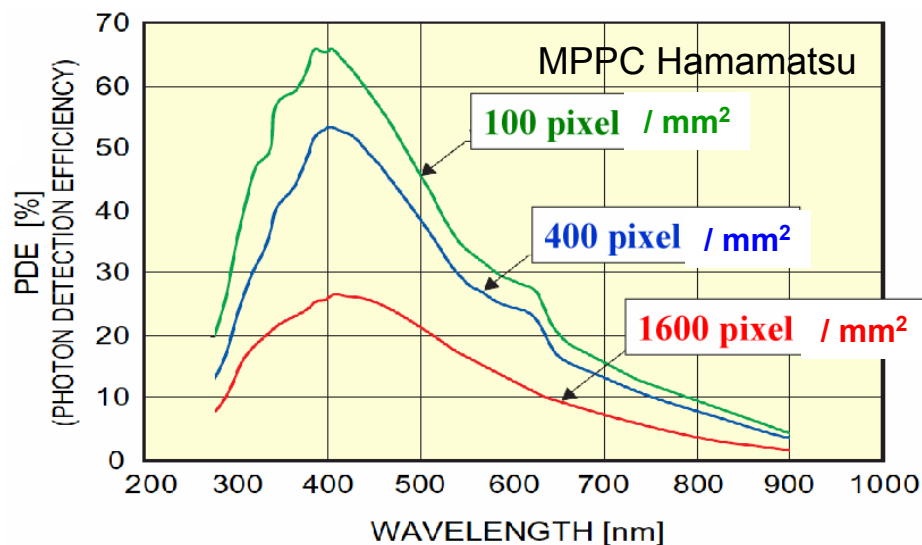
- intrinsic **Q**uantum **E**fficiency ( $\sim 80\%$  for Si)  $\rightarrow \lambda$  dependent
- fraction of sensitive area (20-80%)  $\rightarrow$  technology
- probability of Geiger breakdown ( $\sim 100\%$ )  $\rightarrow V_{\text{Bias}}$  (or E) dep.
- surface reflection losses
- pixel recovery time ( $R_{\text{pixel}}$ )



influence of quantum efficiency



influence of sensitive area

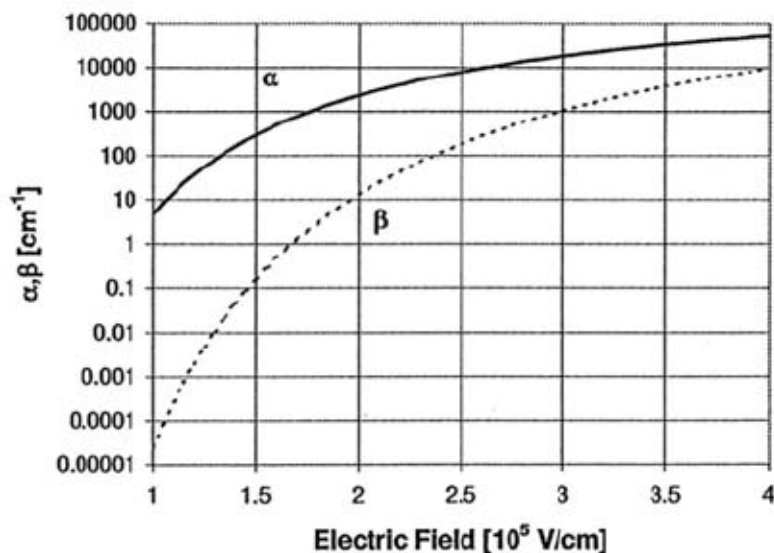


=400nm, including the cross-talk and after pulse

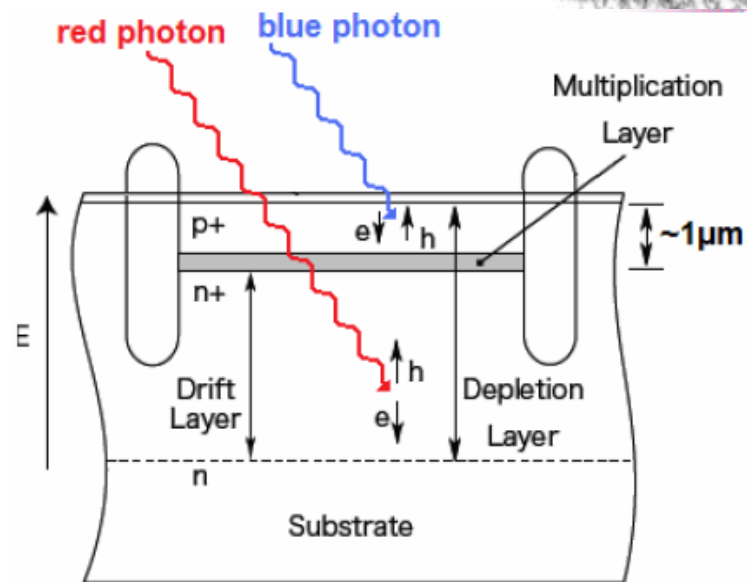
# Photon Detection Efficiency (PDE)

The triggering probability depends on the position where the primary electron-hole pair is generated and it depends on the overvoltage.

Electrons have in silicon a better chance to trigger a breakdown than holes. Therefore a conversion in the p+ layer has the highest probability to start a breakdown.



Ionization coefficients for electrons ( $\alpha$ ) and holes ( $\beta$ ) in silicon



Wavelength dependence of PDE linked to depth of penetration of photon

Blue (470nm)	0.6 $\mu\text{m}$
Green (525nm)	1.2 $\mu\text{m}$
Yellow (590nm)	2.2 $\mu\text{m}$
Red (625nm)	2.9 $\mu\text{m}$

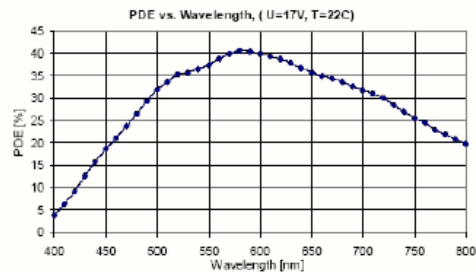
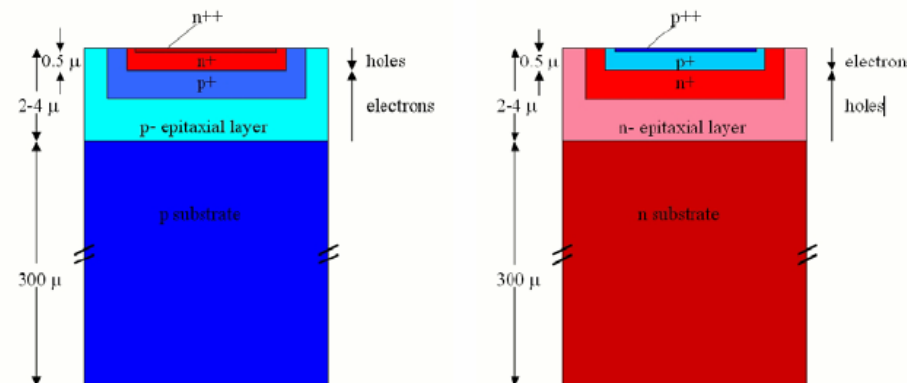
# Photon Detection Efficiency (PDE)

Photons with short wavelengths will be absorbed in the very first layer of Si and create an electron-hole pair.

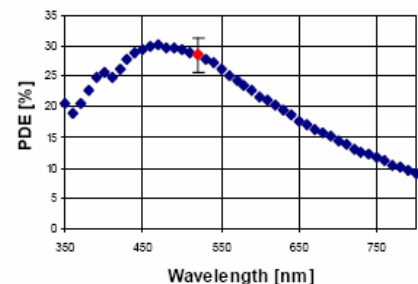
In a structure with a n-type substrate (right) the electrons drift towards the high field of the p-n junction and trigger with high probability a breakdown.

A G-APD made on a n-type substrate will be preferential sensitive for blue light.

A G-APD made on a p-type substrate (left) needs long wavelengths for the creation of electrons in the p-layer behind the junction and will have the peak sensitivity in the green/red.



Photonique/CPTA  
(SSPM\_0710G9MM)



Hamamatsu  
(PSI-33-050C)

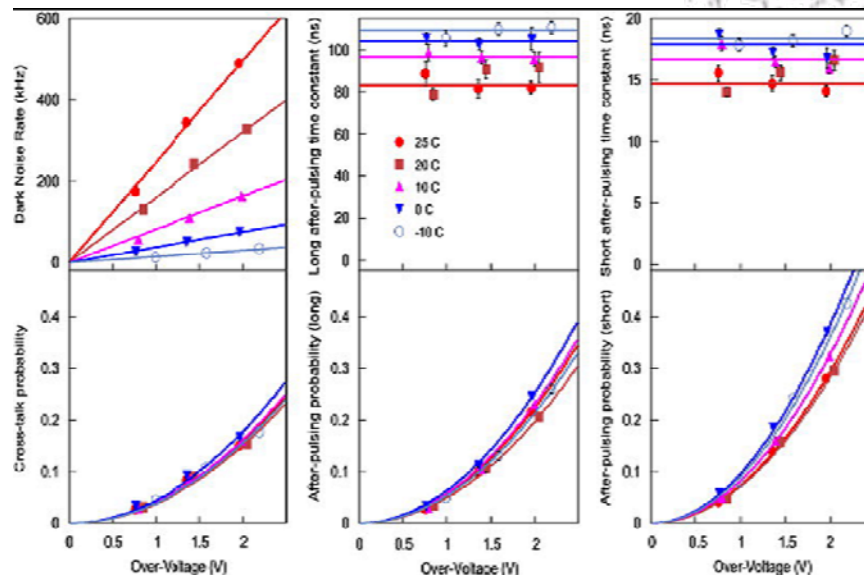
# Consequences of the basic properties

We want the highest possible PDE and the best time resolution.

→ operate G-APDs at high overvoltage (high gain)

Consequences are:

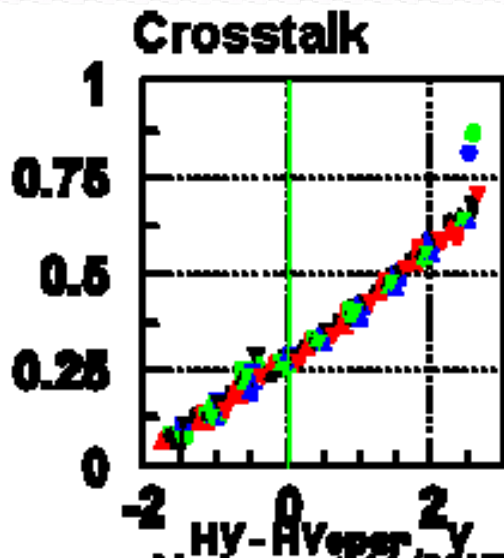
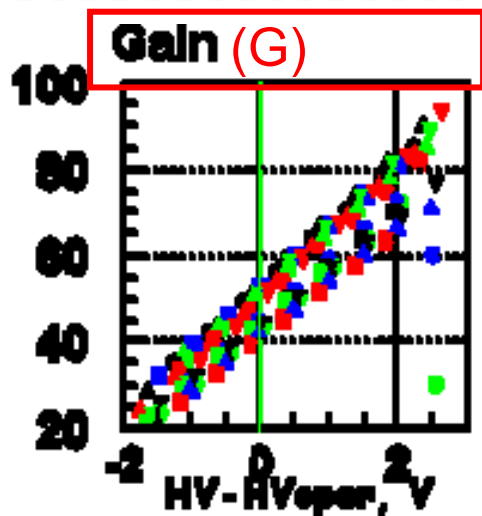
- Large number of dark counts
- High crosstalk probability
- High afterpulse probability
- Large currents and possible self-heating in a high rate environment



The operation is different compared to that of a PMT where the gain can be chosen in a wide range with no or little consequences on the PDE.

G-APDs need a re-engineering (e.g. reduction of C) when the gain has to be modified.

# SiPM voltage dependence

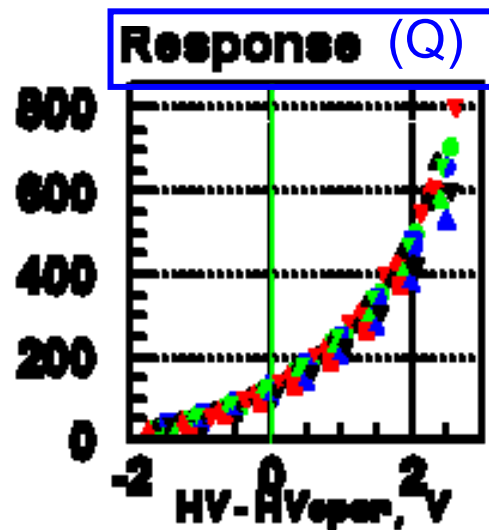
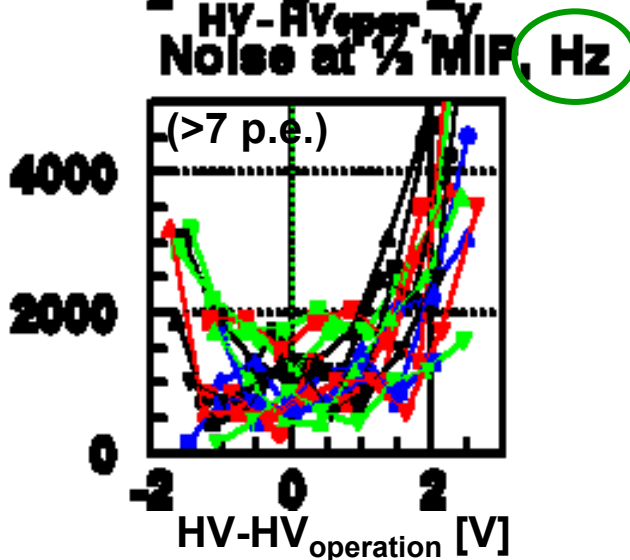
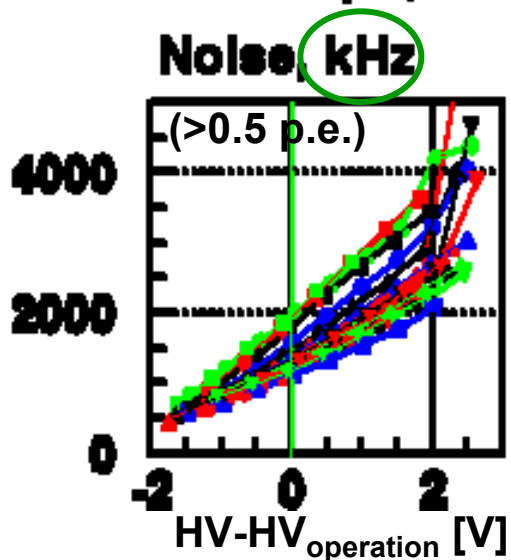


**Gain (G)**

$$\frac{dG}{dV} \sim 2.5 \frac{\%}{0.1V}$$

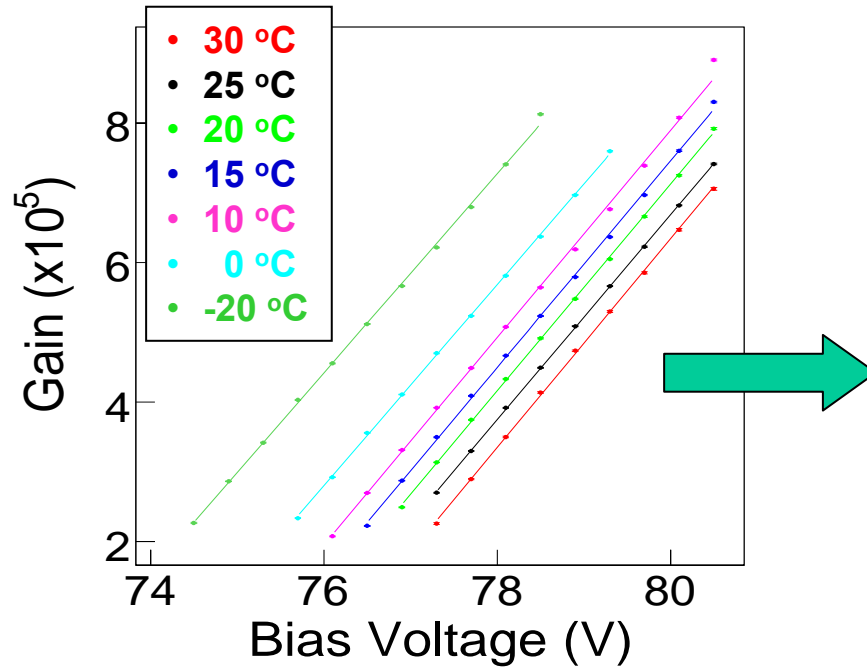
Total Charge (Q) = G \* PDE

$$\frac{dQ}{dV} \sim 7.0 \frac{\%}{0.1V}$$



characterization of 4000 SiPM MEPHI/Pulsar, from E. Tarkovski, ITEP

# SiPM temperature dependence



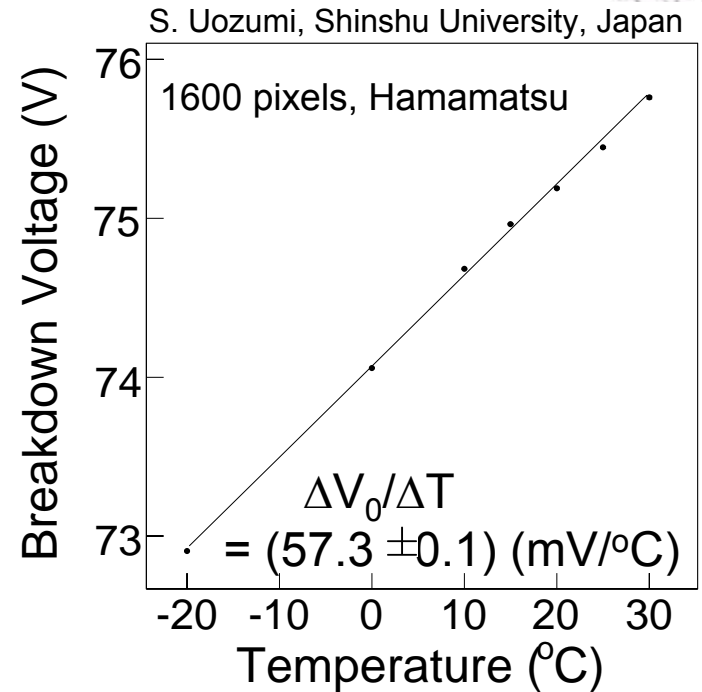
Gain =  $C (V_{\text{bias}} - V_0) / e$   
 C : Pixel capacity  
 $V_0$  : Breakdown voltage

**Gain (G)**

$$\frac{dG}{dT} \sim -1.7 \frac{\%}{\text{K}}$$

**Total Charge (Q) = G \* PDE**

$$\frac{dQ}{dT} \sim -4.5 \frac{\%}{\text{K}}$$



$V_0$  is linear to temperature.  
 → temperature changes affect the over-voltage =  $V_{\text{bias}} - V_0(T)$   
 → therefore all SiPM properties

← 1200 pixels, MEPHI/Pulsar

# Summary of SiPM features

High gain ( $\sim 10^6$ ) → simplest r/o electronics

Low electronics noise

Low bias voltage ( $\sim 50$  V)

Low power consumption ( $\leq 50$   $\mu\text{W}/\text{mm}^2$ )

Insensitivity to magnetic field → next generation of HEP detectors

Compact and light → direct couple to active material / space mission

Excellent photon counting capability → astro-particle physics

Very low charge particle sensitivity (negligible nuclear counting effect)

Very good timing ( $\leq 100$  ps) → medical applications

Small recovery time

Good temperature and voltage stability

Room temperature operation

Relatively low cost → high n

## Drawbacks and limitations:

- Small size (established up to  $3 \times 3 \text{mm}^2$ )
- Not enough PDE ( $\sim 20\text{-}40\%$ )
- High dark rate
- Limited dynamic range
  
- Optical crosstalk being reduced ( $< 10\%$ )
- Sensitivity to blue established, what about UV?

→ Fast developing technology

Solutions to the remaining open issues are coming



# The measurement of showers

or “from energy to signal”

Step 1: Convert energy to light

Step 2: Convert light to electrical signal

Step 3: Reading an electrical signal



# Signal Acquisition

## Determine energy deposited and event time in detector

- Detector signal generally a short current pulse:
  - thin silicon detector (10 –300  $\mu\text{m}$ ): 100 ps–30 ns
  - thick ( $\sim\text{cm}$ ) Si or Ge detector: 1 –10  $\mu\text{s}$
  - proportional chamber: 10 ns –10  $\mu\text{s}$
  - Microstrip Gas Chamber: 10 –50 ns
  - Scintillator+ PMT/APD: 100 ps–10  $\mu\text{s}$

$$E \propto Q_s = \int i_s(t) dt$$

- Necessary to integrate detector signal current:
  - integrate charge on input capacitance
  - use integrating (“charge sensitive”) preamplifier
  - amplify current pulse and use integrating ADC

# Relevant aspects of electronics for calorimetry

**Dynamic range:** maximum signal/minimum signal (or noise)

typically:  $10^3 - 10^5$

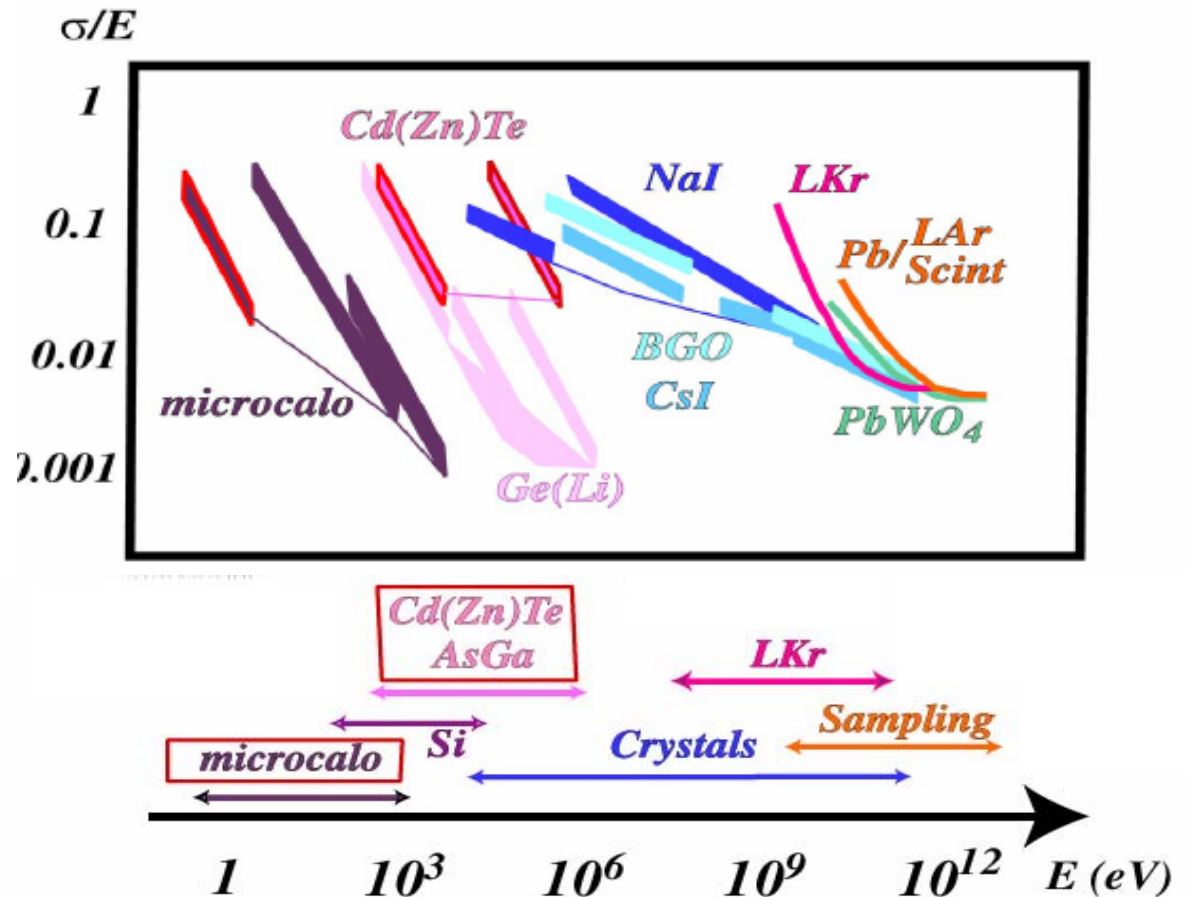
often specified **in dB**

$$= 20 \log V_{\max}/V_{\min} = 60 - 100 \text{ dB}$$

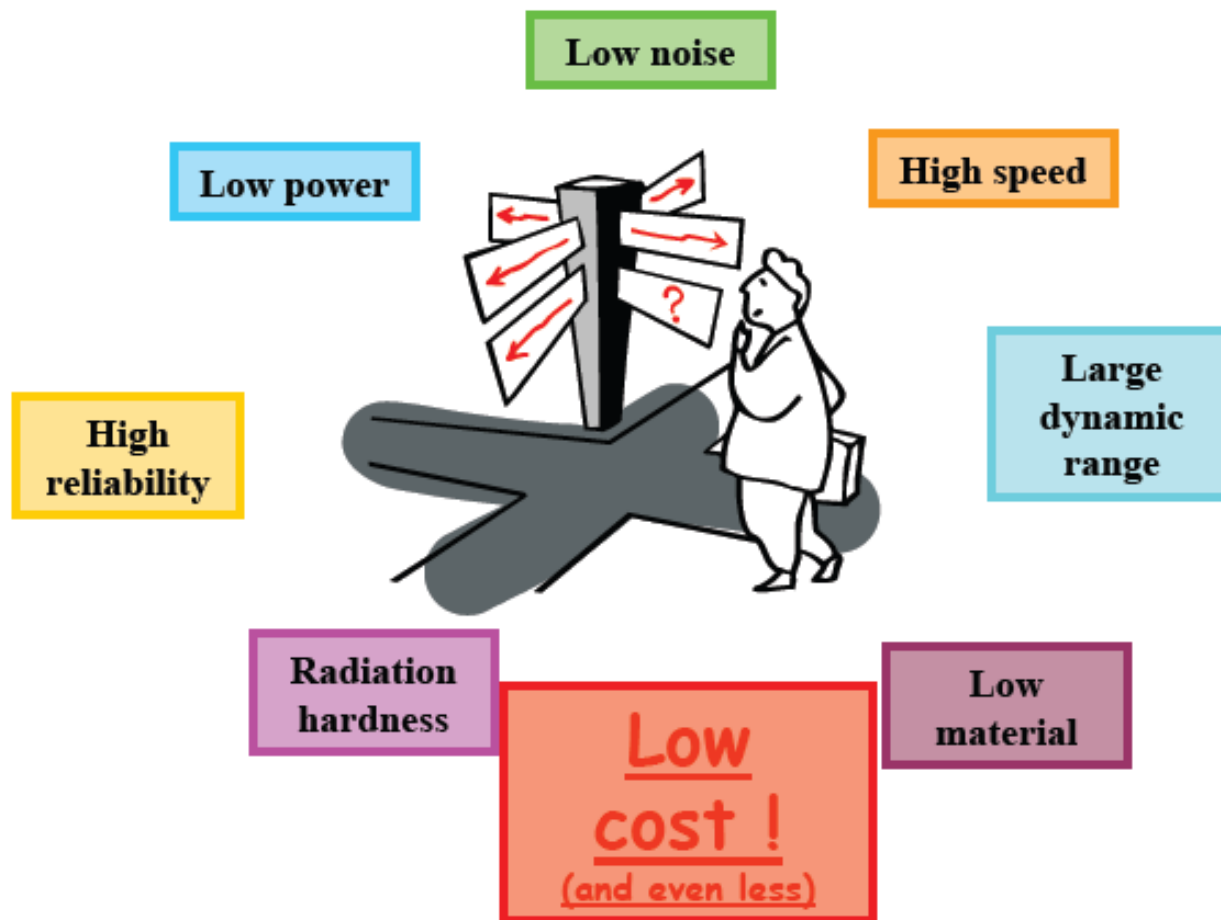
also **in bits:**

$$2^n = V_{\max}/V_{\min} = 10 - 18 \text{ bits}$$

The large dynamic range is a **key parameter** for calorimeter electronics

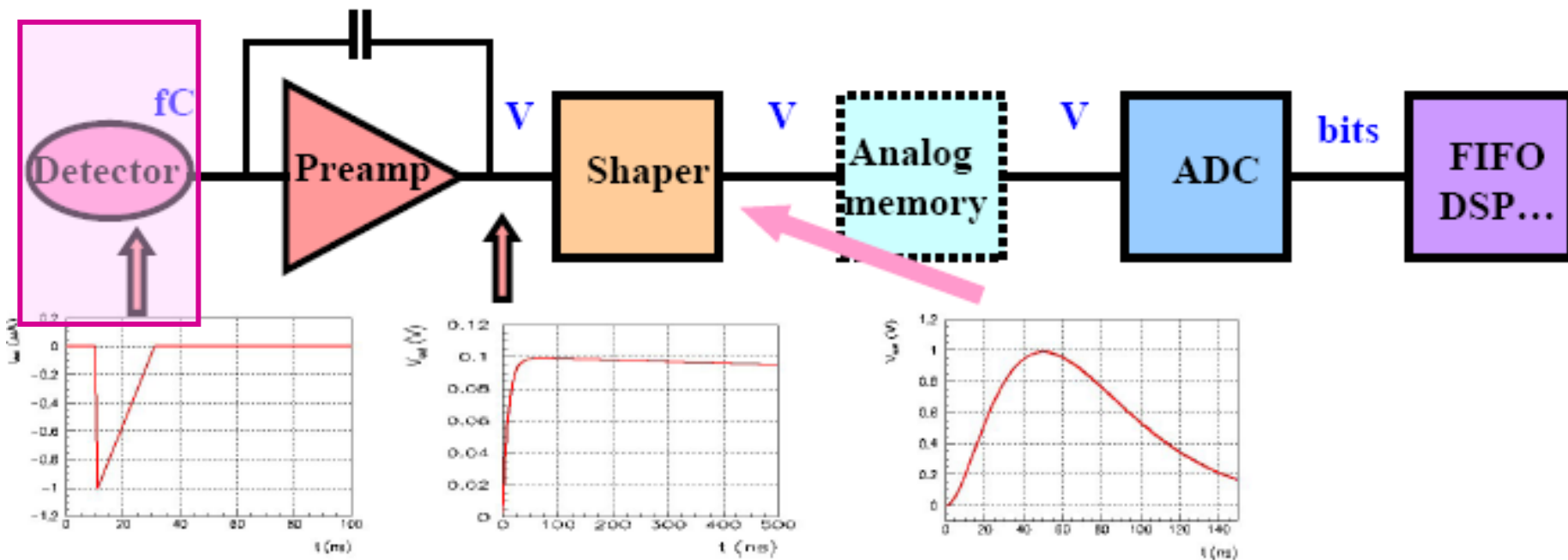


# Readout electronics requirements



# Overview of readout electronics

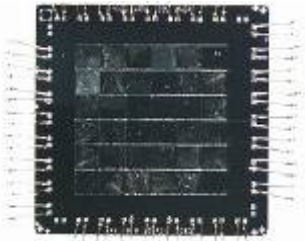
Most front-ends follow a similar architecture



- Very small signals ( $fC$ ) -> need **amplification**
- Measurement of **amplitude** and/or **time** (ADCs, discriminator, TDCs)
- Thousands to millions of channels

# Detector(s)

A large variety of detectors  
But similar modeling



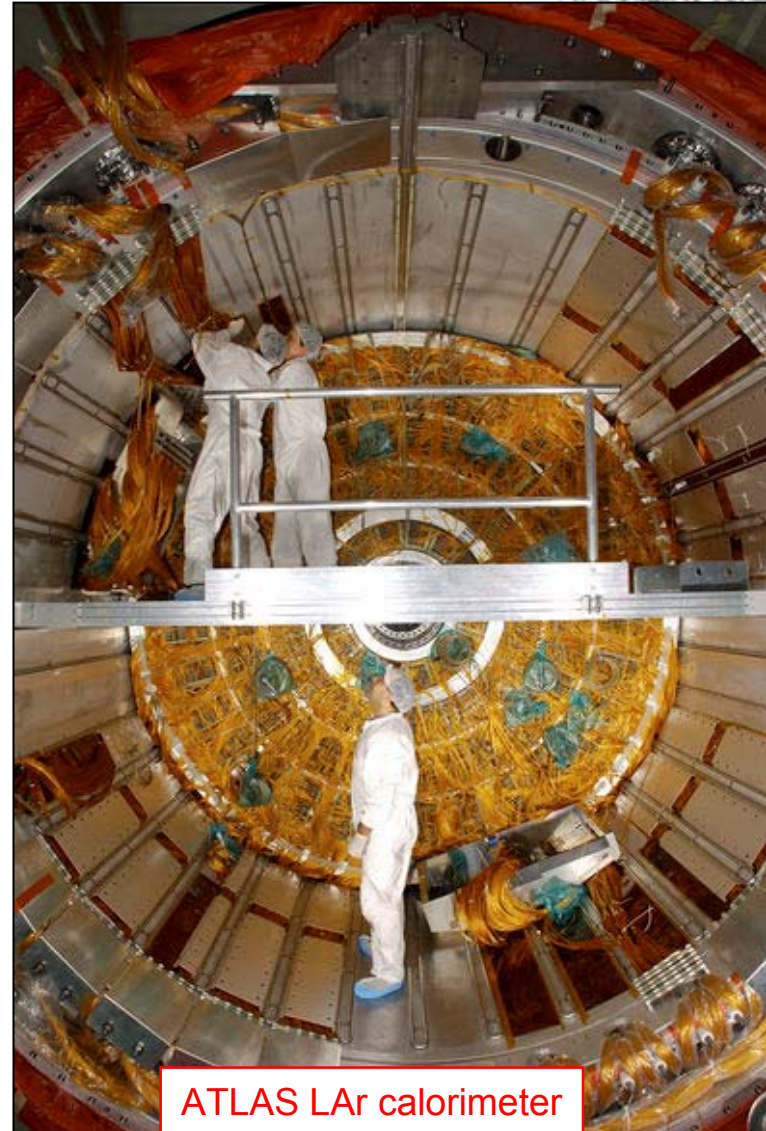
6x6 pixels, 4x4 mm<sup>2</sup>  
HgTe absorbers, 65 mK  
12 eV @ 6 keV



PMT in ANTARES



CMS pixel module



ATLAS LAr calorimeter

# Detector modeling

Detector = capacitance  $C_d$

- Pixels : 0.1-10 pF
- PMs : 3-30 pF
- Ionization chambers: 10-1000 pF
- Sometimes effect of transmission line

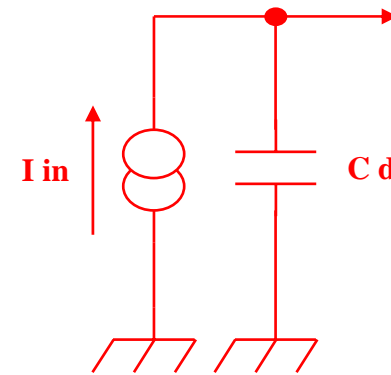
Signal : current source

- Pixels :  $\sim 100$  e $^-$ / $\mu$ m
- PMs : 1 photoelectron  $\rightarrow 10^5$ - $10^7$  e $^-$
- Modeled as an impulse (Dirac) :

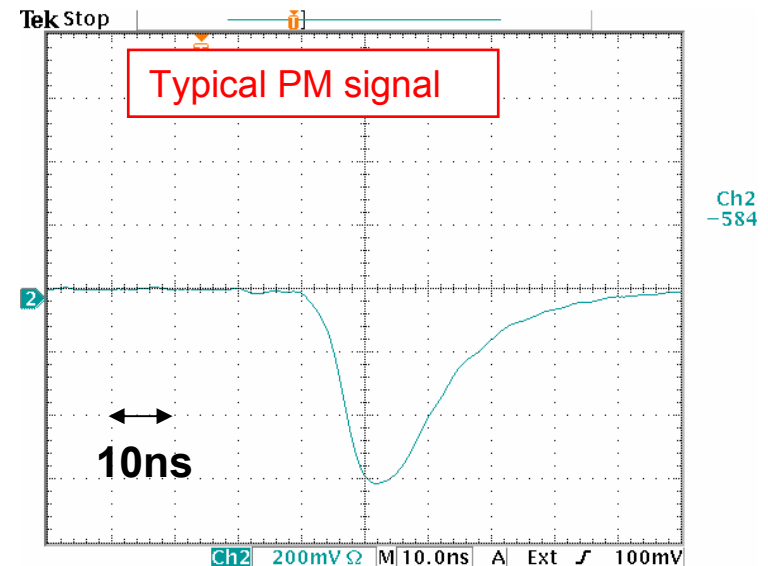
$$i(t) = Q_0 \delta(t)$$

Missing :

- High Voltage, bias
- Connections, grounding
- Neighbors
- Calibration...



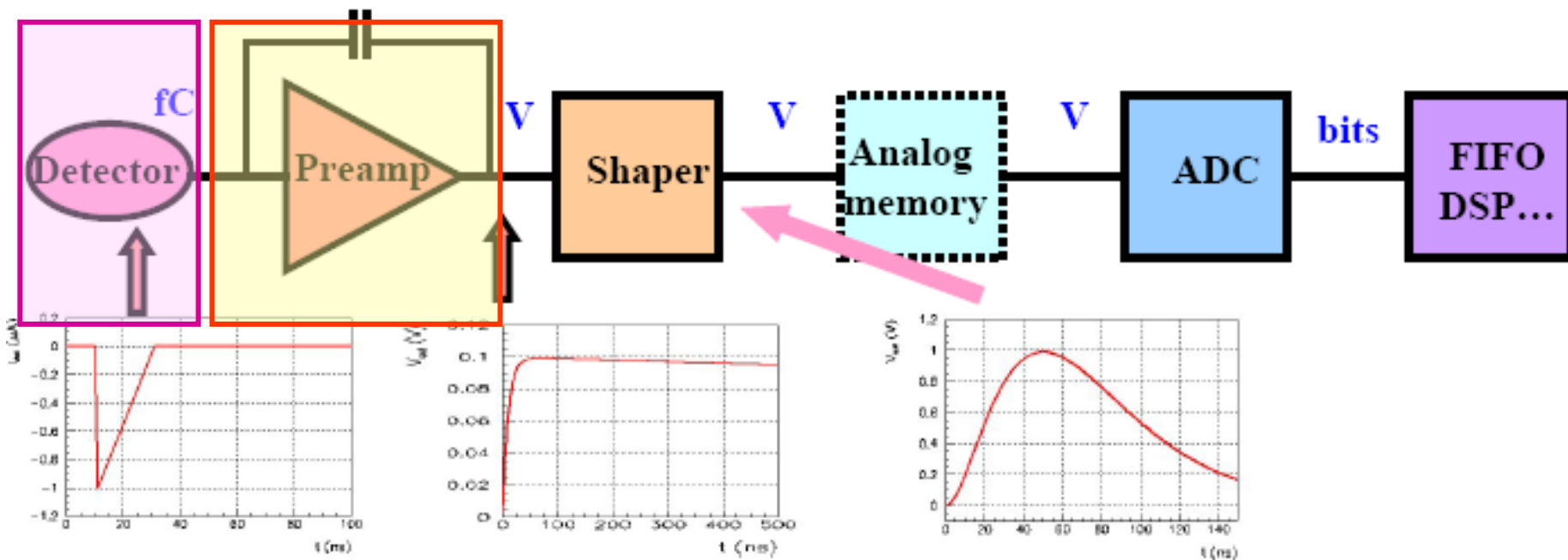
Detector model



Ch2  
-584

# Overview of readout electronics

Most front-ends follow a similar architecture

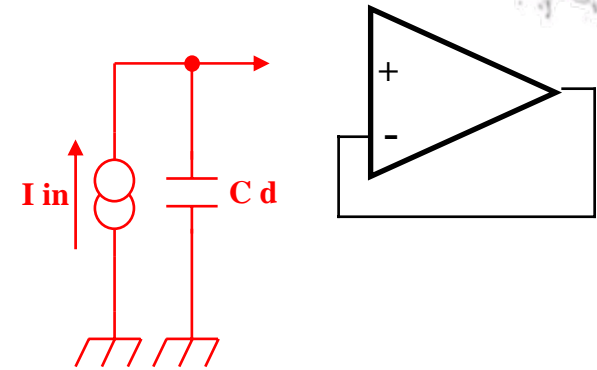


- Very small signals (fC) -> need **amplification**
- Measurement of **amplitude** and/or **time** (ADCs, discriminator, TDCs)
- Thousands to millions of channels

# Reading the signal

## Signal

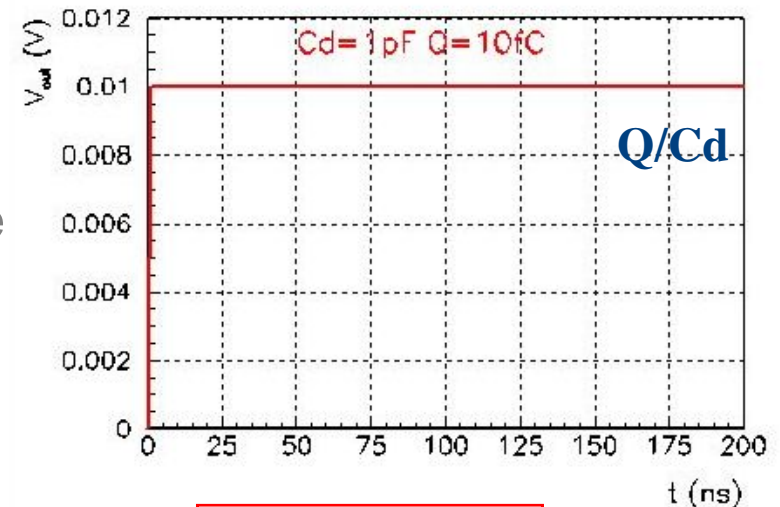
- Signal = current source
- Detector = capacitance  $C_d$
- Quantity to measure
  - Charge  $\rightarrow$  integrator needed
  - Time  $\rightarrow$  discriminator + TDC



Voltage readout

## Integrating on $C_d$

- Simple :  $V = Q/C_d$
- « Gain » :  $1/C_d$  :  $1 \text{ pF} \rightarrow 1 \text{ mV/fC}$
- Need a follower to buffer the voltage  $\rightarrow$  parasitic capacitance
- Gain loss, possible non-linearity
- crosstalk
- Need to empty  $C_d$ ...



Impulse response



# Reading the signal (II)

If the input time constant of the amplifier,  $\tau = C_i R_i$  is large compared to the duration of the current pulse of the detector,  $t_c$  the current pulse will be integrated on the capacitance  $C_i$ .

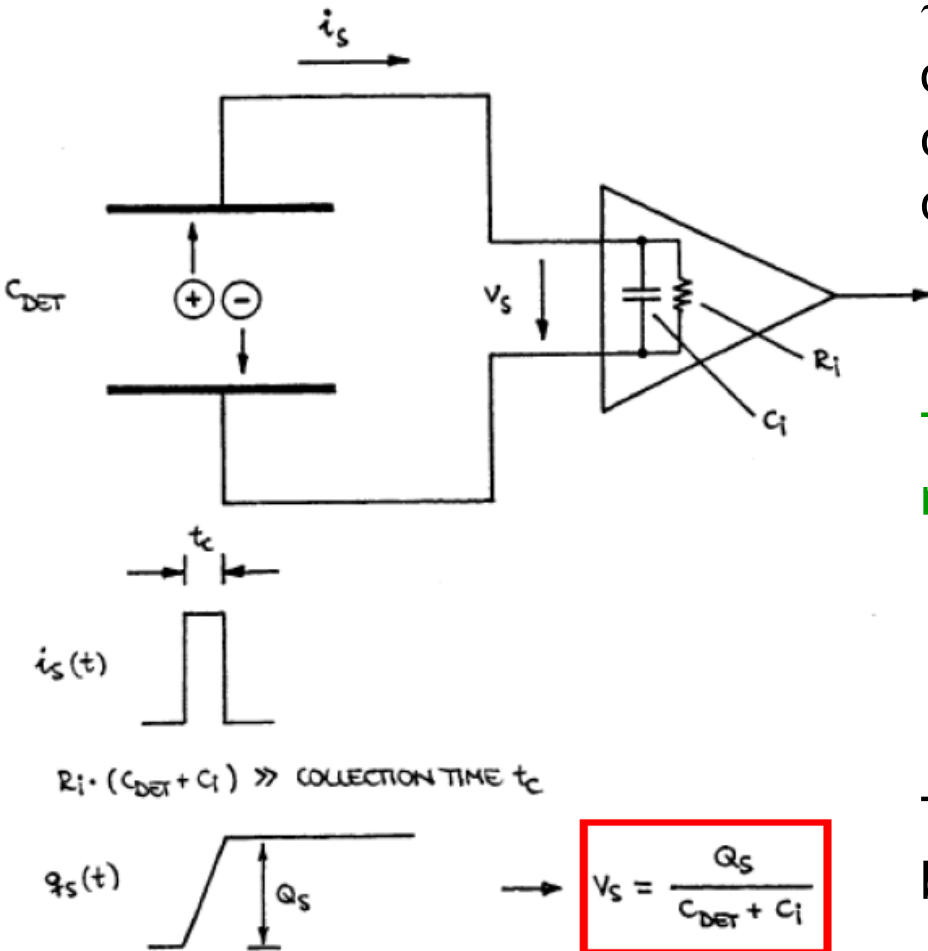
The resulting voltage at  $C_i$  and  $R_i$  is  

$$v_i = v_s = Q_s / (C_{det} + C_i)$$

The fraction of the signal charge measured is:

$$\frac{Q_i}{Q_s} = \frac{C_i v_i}{v_i (C_i + C_{det})} = \frac{1}{1 + C_{det} / C_i}$$

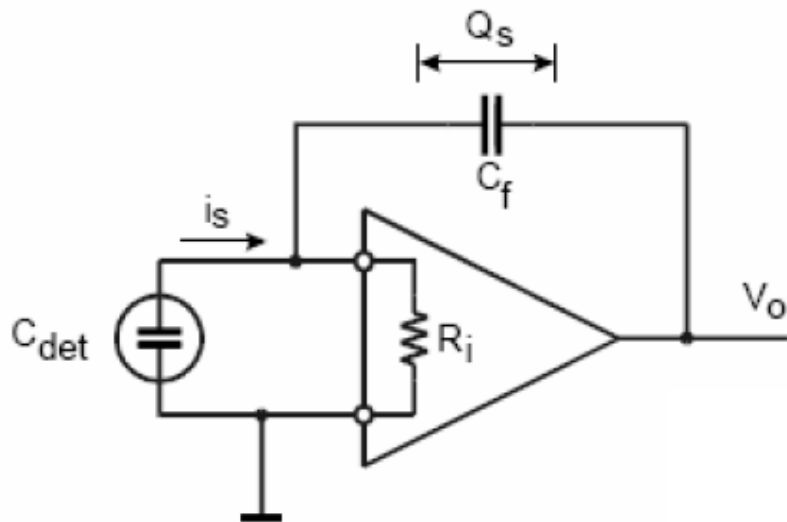
The dynamic input capacitance  $C_i$  should be  $\gg C_{det}$  to get a good ratio close to 1



Depends on the detector capacitance

# Charge sensitive amplifier

Add feedback capacitor  $C_f$ :



Voltage gain  $dV_o/dV_i = -A \rightarrow v_o = -Av_i$   
 Input impedance =  $\infty$  (no signal current flows into amplifier input)

Voltage diff. across  $C_f$ :  $v_f = (A+1)v_i$   
 $\rightarrow$  Charge deposited on  $C_f$ :  $Q_f = C_f v_f$   
 $Q_i = Q_f$  (since  $Z_i = \infty$ )  
 $\rightarrow$  Effective input capacitance

$$C_i = Q_i/v_i = C_f(A+1)$$

“dynamic input capacitance”

Amplifier gain:

$$A_Q = \frac{dV_o}{dQ_i} = \frac{A \cdot v_i}{C_i \cdot v_i} = \frac{A}{C_i} = \frac{A}{A+1} \frac{1}{C_f} \approx \frac{1}{C_f} \quad (A \gg 1)$$

# Charge sensitive amplifier (II)

So finally the fraction of charge signal measured by the amplifier is:

$$\frac{Q_i}{Q_s} = \frac{C_i v_i}{v_i (C_i + C_{\text{det}})} = \frac{1}{1 + C_{\text{det}} / C_i} \quad C_f \approx \frac{A}{C_i} \quad (A \gg 1)$$

Example:

$$A = 10^3$$

$$C_f = 1 \text{ pF}$$



$$C_i = 1 \text{ nF}$$

$$C_{\text{det}} = 10 \text{ pF}$$



$$Q_i/Q_s = 0.99$$

$$(C_i \gg C_{\text{det}})$$

$$C_{\text{det}} = 500 \text{ pF}$$

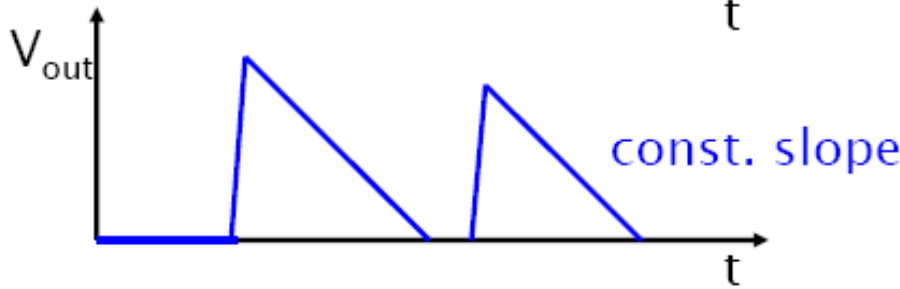
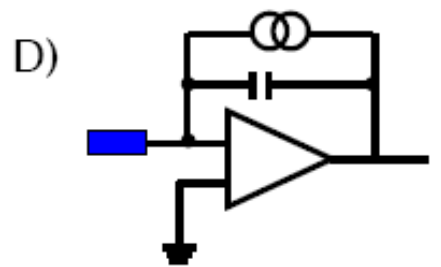
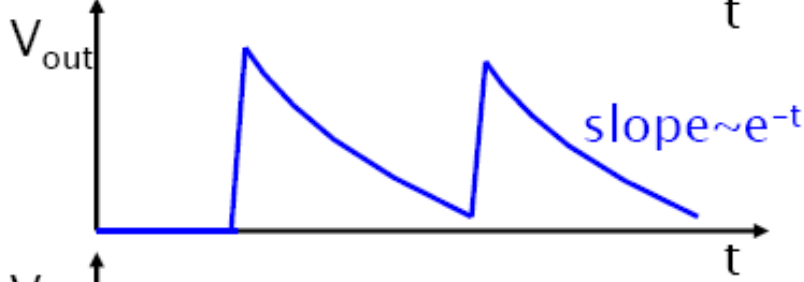
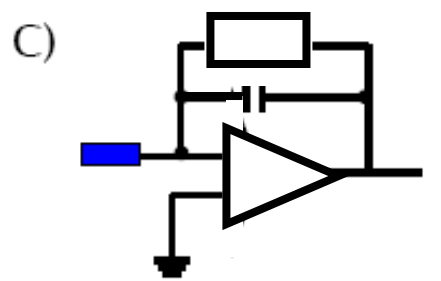
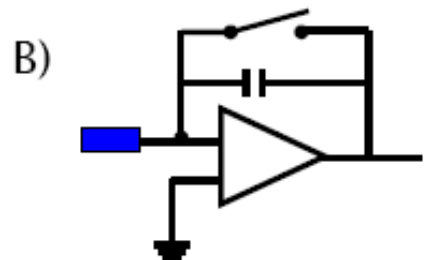
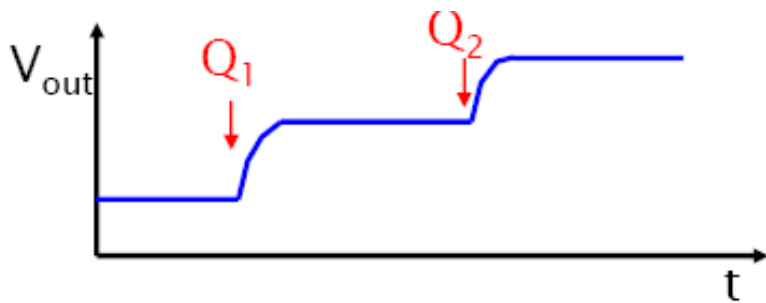
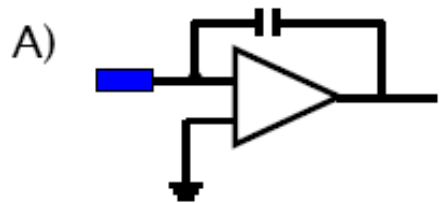


$$Q_i/Q_s = 0.67$$

$$(C_i \sim C_{\text{det}})$$

↑  
Si det: 50um thick, 500mm<sup>2</sup> area

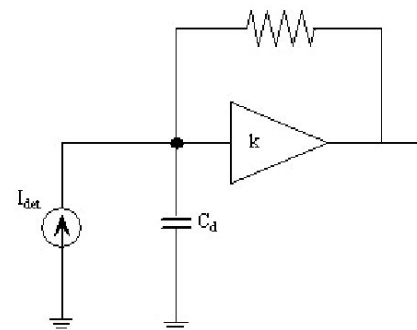
# Charge sensitive (pre)-amplifier



# Current preamplifiers :

## Transimpedance configuration

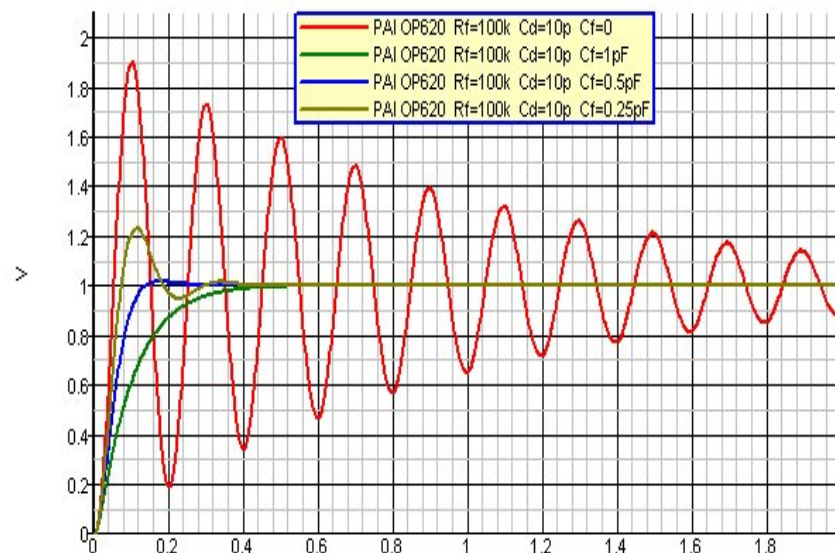
- $V_{out}(\omega)/i_{in}(\omega) = -R_f / (1+Z_f/GZ_d)$
- Gain =  $R_f$
- High counting rate
- Typically optical link receivers



## Easily oscillatory

- Unstable with capacitive detector
- Inductive input impedance  
 $L_{eq} = R_f / \omega_C$
- Resonance at :  $f_{res} = 1/2\pi \sqrt{L_{eq} C_d}$
- Quality factor :  $Q = R / \sqrt{L_{eq}/C_d}$ 
  - $Q > 1/2 \rightarrow$  ringing
- Damping with capacitance  $C_f$ 
  - $C_f = 2 \sqrt{C_d/R_f G_0 \omega_0}$
  - Easier with fast amplifiers

Current sensitive preamp



Step response of current sensitive preamp

# Charge vs Current preamps

## Charge preamps

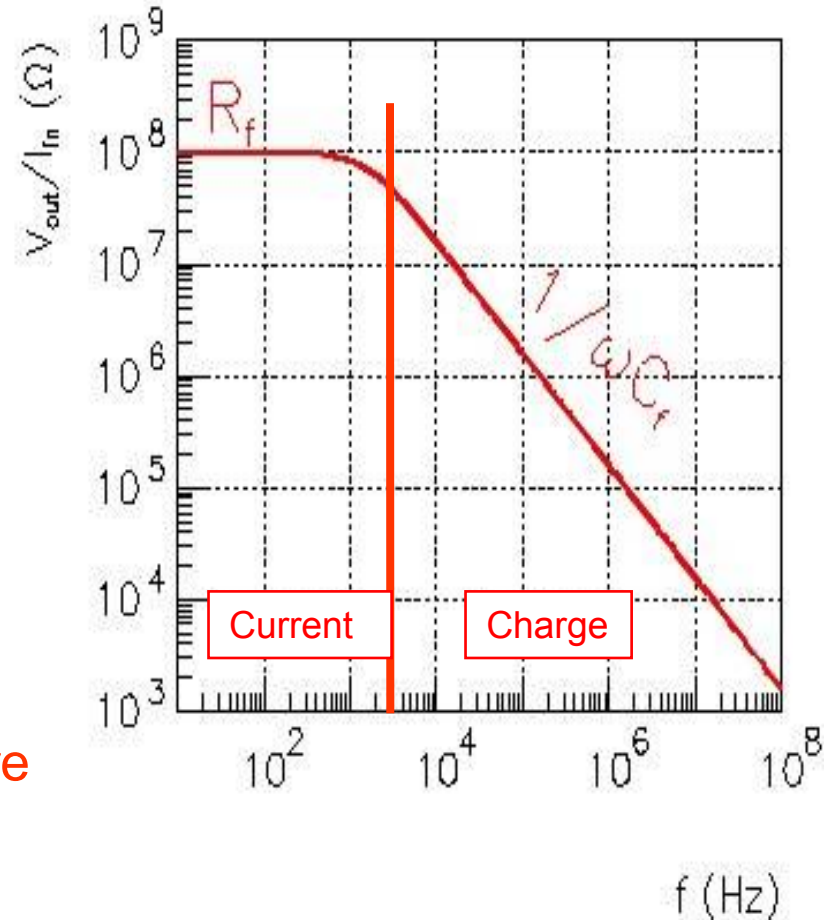
- Best noise performance
- Best with short signals
- Best with small capacitance

## Current preamps

- Best for long signals
- Best for high counting rate
- Significant parallel noise

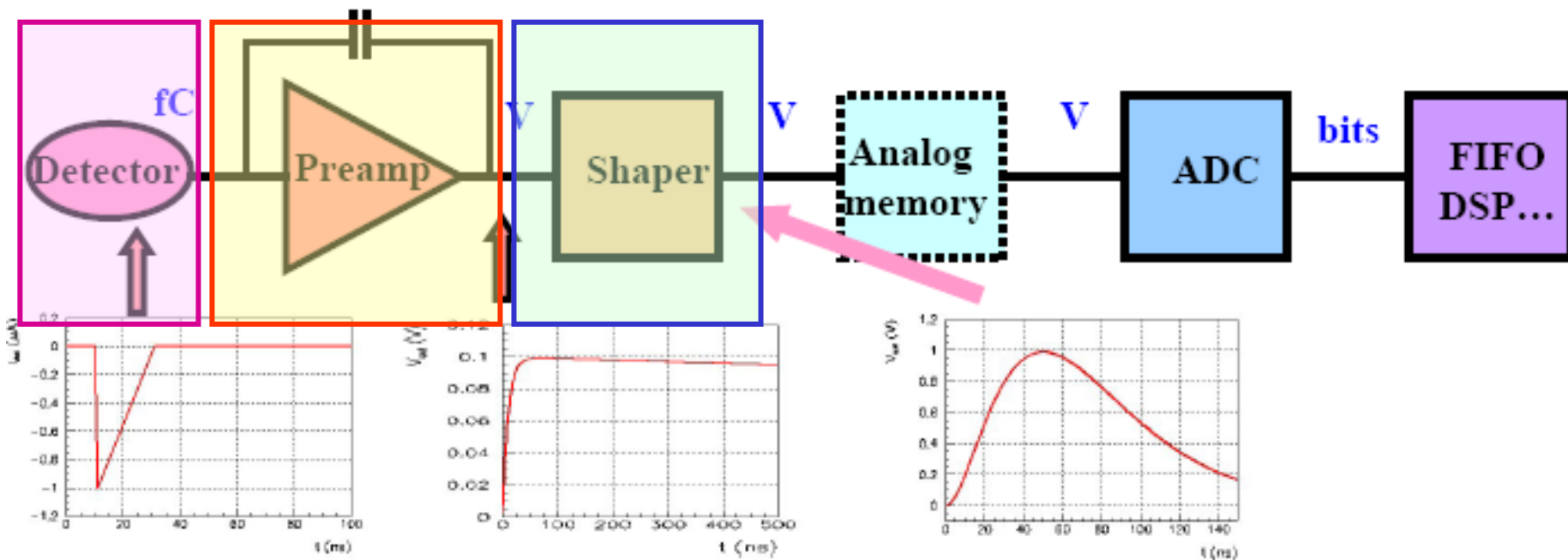
Charge preamps are not slow, they are long

Current preamps are not faster, they are shorter (but easily unstable)



# Overview of readout electronics

Most front-ends follow a similar architecture

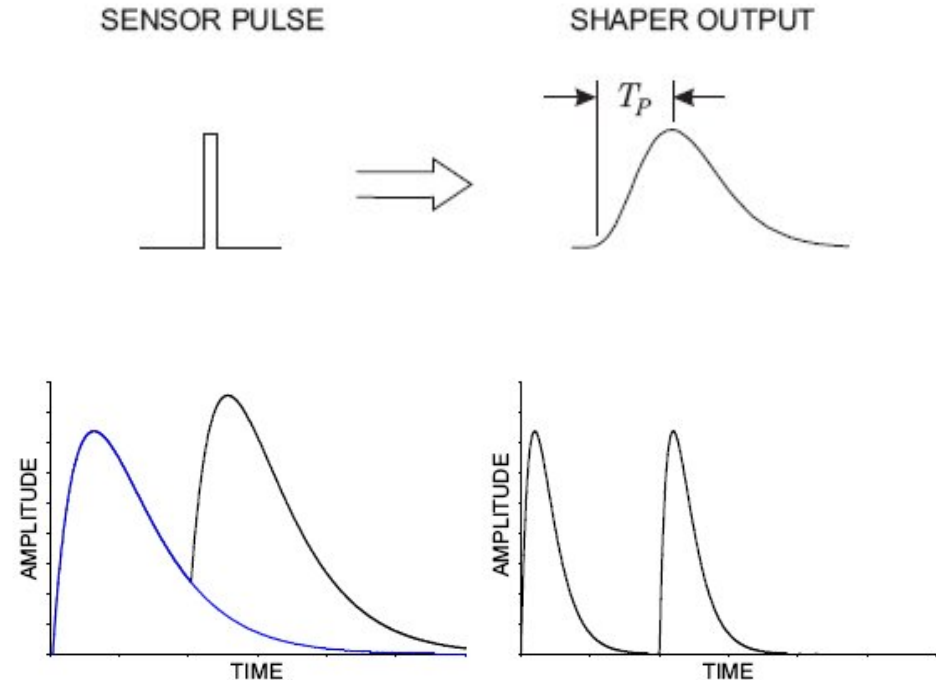


- Very small signals (fC) -> need **amplification**
- Measurement of **amplitude** and/or **time** (ADCs, discriminator, TDCs)
- Thousands to millions of channels

# Pulse shaping

Two conflicting objectives:

- 1) Limit the bandwidth to match the measurement time.  
→ too large bandwidth increases the noise
- 2) Contain the pulse width so that successive signal pulses can be measured without overlap (pile-up)  
→ Short pulse duration increases the allowed signal rate but also noise



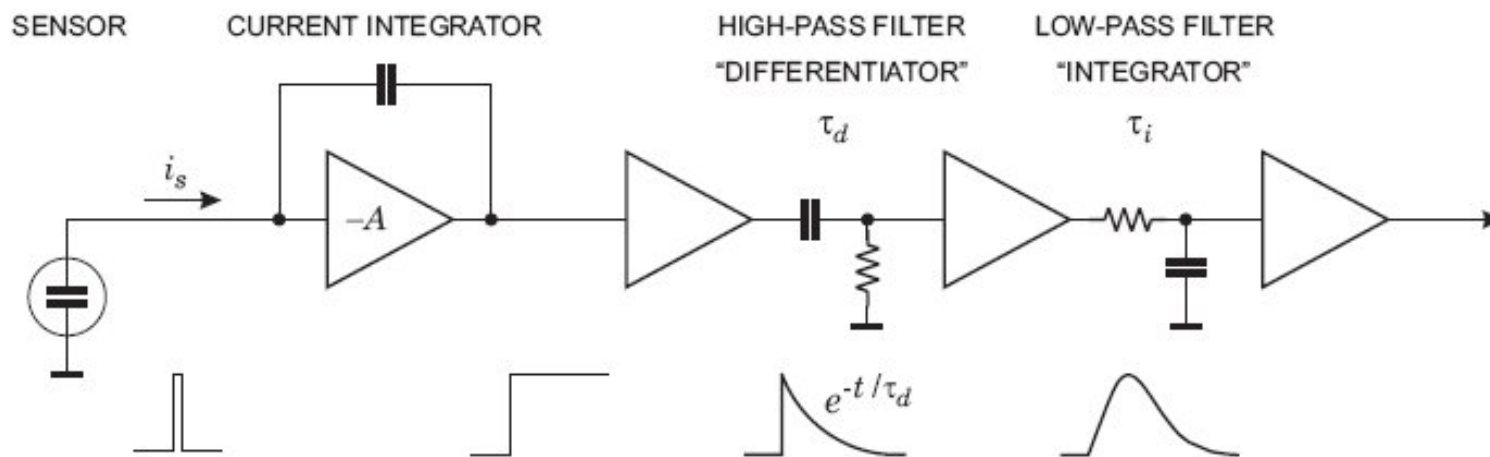


# CR-RC shaper

Example of a simple shaper: CR-RC

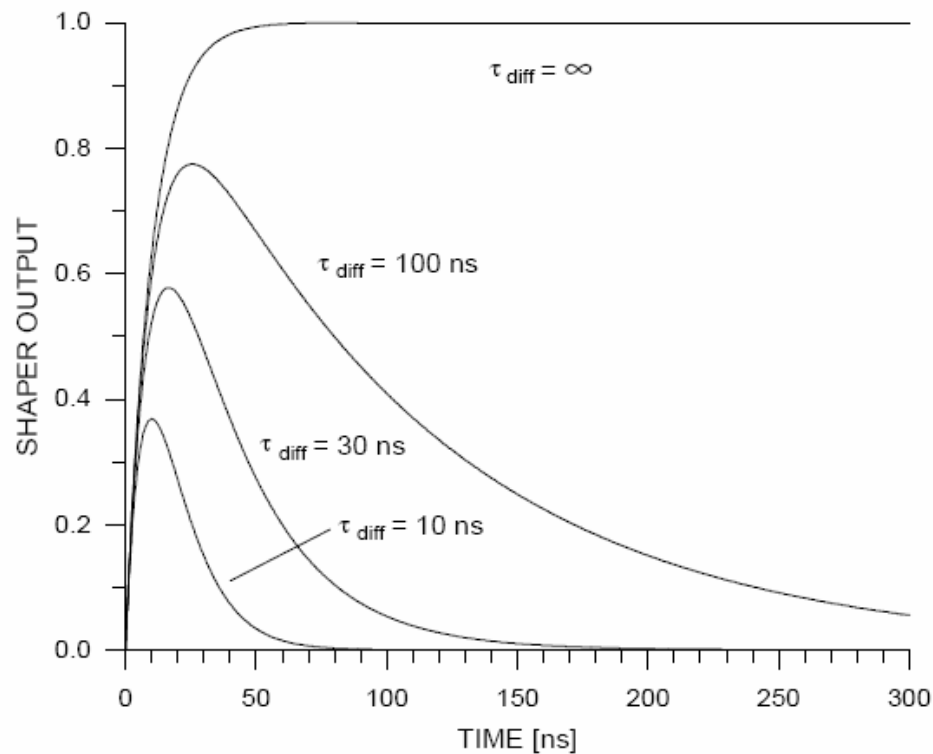
- the **high-pass filter** sets the duration of the pulse to have a decay time  $\tau_d$
- the **low-pass filter** increases the rise time to limit the noise bandwidth

**key design parameter: peaking time** → it dominates the noise bandwidth



# CR-RC shaper (II)

Effect of a CR-RC shaper with  
fix integrator time constant = 10ns  
and variable differentiator time constant

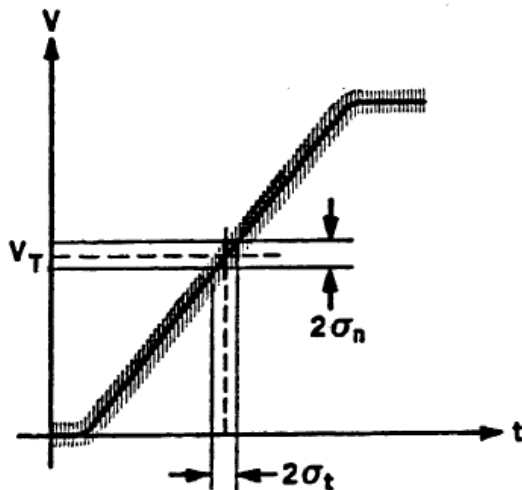


# Time measurements

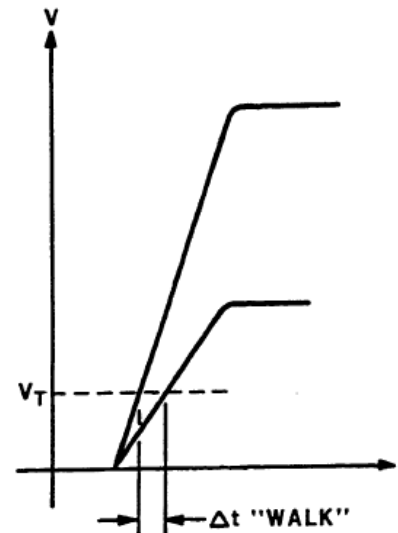
Time measurements are characterized by their **slope-to-noise ratio**

Two main effects contribute to the deterioration of a time measurement i.e. time of threshold crossing fluctuates due to:

jitter



time walk

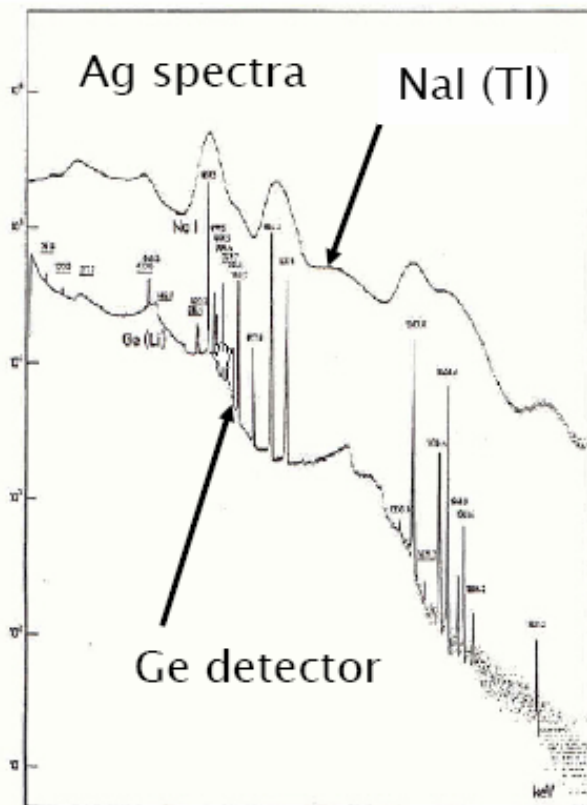


Often driven by the time constant of the shaper which determines rise time & amplifier bandwidth

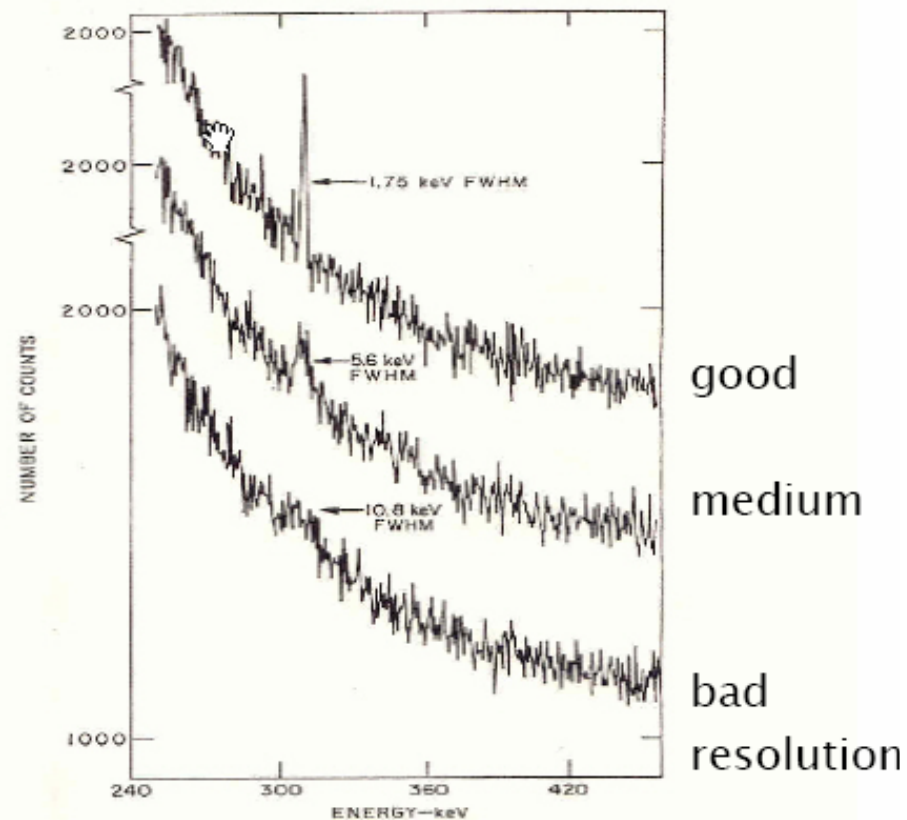
# Readout overview

Experiment	Shaping	tp	Technology	Dyn. Rge	Gains	ADC
ATLAS em	CRRC <sup>2</sup>	50 ns	BiCMOS 1.2 $\mu$	16 bits	1-10-100	12 bits 5 MHz
ATLAS had	Bessel 9	50 ns	Passive hybrid	16 bits	1-64	
BABAR	CRRC <sup>2</sup>	400 ns	BiCMOS 1.2 $\mu$	18 bits	1-4-32-256	10 bits 4 MHz
CMS em	RC <sup>2</sup>	50 ns	CMOS 0.25 $\mu$	16 bits	1-6-12	12 bits 40 MHz
CMS had	Gated int	25 ns				
DØ	CR	350 ns	Bipolar hyb	15 bits	1-8	12 bits
FLC	CRRC	150 ns	BiCMOS	16 bits	1-8-64	
KLOE	Bessel 3	200 ns	Bipolar hyb.	12 bits	1	
LHCb em	DLC	50 ns	BiCMOS 0.8 $\mu$	12 bits	1	12 bits 40 MHz
NA48	Bessel 8	70 ns	BiCMOS 1.2 $\mu$	14 bits	1-2.5-6-18	10 bits 40 MHz
Opera TT	CRRC <sup>2</sup>	150 ns	BiCMOS 0.8 $\mu$		1	

# Why bother about resolution and noise?



(J.C.I. Philippot, IEEE Trans. Nucl. Sci. NS-17/3 (1970) 446)



G.A. Armantrout *et al.*, IEEE Trans. Nucl. Sci. NS-19/1 (1972) 107

low noise improves the resolution and the ability to distinguish (signal) structures

low noise improves the signal to background ratio (signal counts are in fewer bins and thus compete with fewer background counts)

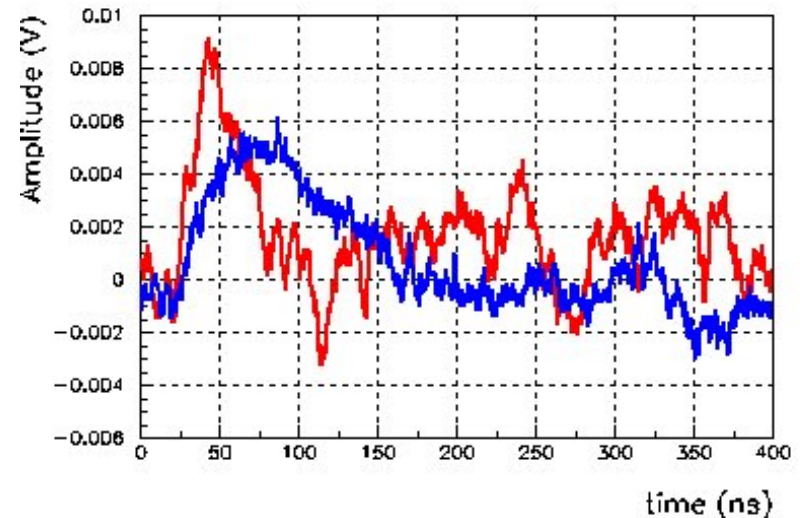
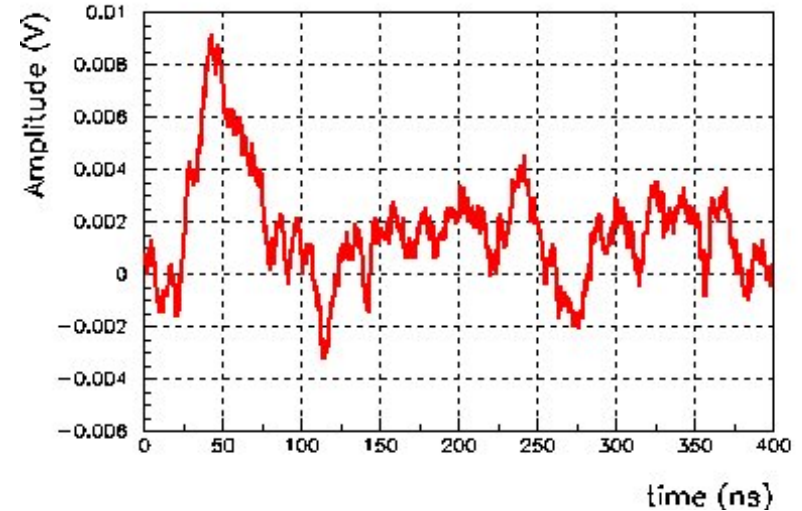
# Electronics noise

## Definition of Noise

- Random fluctuation superimposed to interesting signal
- Statistical treatment

## Three types of noise

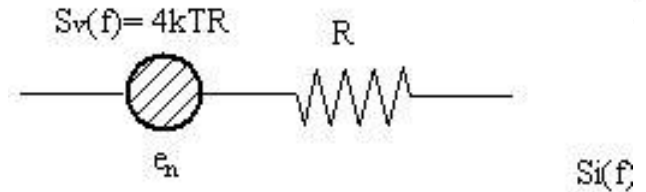
- Fundamental noise (Thermal noise, shot noise)
- Excess noise ( $1/f$  ...)
- Parasitic → EMC/EMI (pickup noise, ground loops...)



# Calculating electronics noise

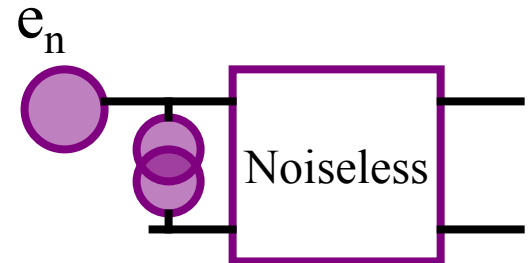
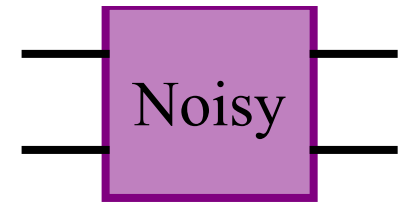
## Fundamental noise

- Thermal noise (**resistors**) :  $S_v(f) = 4kTR$
- Shot noise (**junctions**) :  $S_i(f) = 2qI$



## Noise referred to the input

- All noise generators can be referred to the input as **2** noise generators :
- A voltage one  $e_n$  in series : **series noise**
- A current one  $i_n$  in parallel : **parallel noise**
- Two generators : no more, no less...



■ **To take into account the Source impedance**

## ■ Golden rule

- **Always calculate the signal before the noise**  
**what counts is the signal to noise ratio**

Noise generators  
referred to the input

# Noise in charge pre-amplifiers

## 2 noise generators at the input

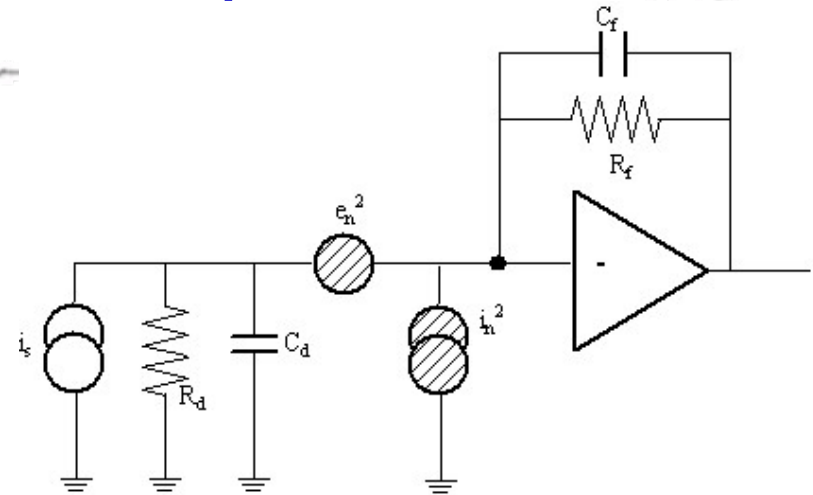
- Parallel noise : ( $i_n^2$ ) (leakage currents)
- Series noise : ( $e_n^2$ ) (preamp)

## Output noise spectral density :

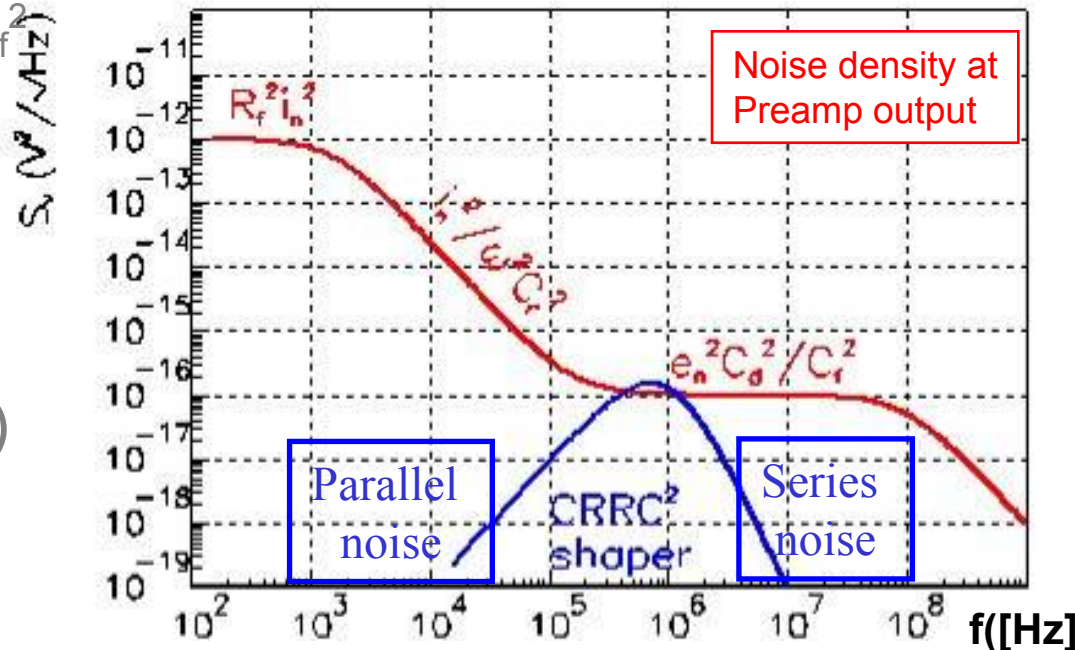
- $S_v(\omega) = (i_n^2 + e_n^2/|Z_d|^2) / \omega^2 C_f^2$   
 $= i_n^2 / \omega^2 C_f^2 + e_n^2 C_d^2 / C_f^2$
- Parallel noise in  $1/\omega^2$
- Series noise is flat, with a « noise gain » of  $C_d/C_f$

## rms noise $V_n$

- $V_n^2 = \int S_v(\omega) d\omega/2\pi \rightarrow \infty (!)$
- Benefit of shaping...



Noise generators in charge preamp



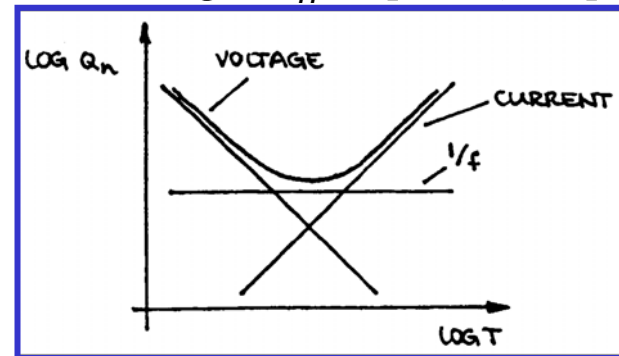


# Equivalent Noise Charge (ENC)

Two basic noise mechanisms: input noise current  $i_n$  [pA/ $\sqrt{\text{Hz}}$ ]  
 input noise voltage  $e_n$  [nV/ $\sqrt{\text{Hz}}$ ]

Equivalent Noise Charge:

$$Q_n^2 = \underbrace{i_n^2}_{\text{from Front End}} \underbrace{T_s F_i}_{\text{from Shaper}} + \underbrace{C_i^2 v_n^2}_{\text{from Front End}} \underbrace{\left(\frac{F_v}{T_s}\right)}_{\text{from Shaper}}$$



where  $T_s$  = characteristic shaping time (e.g. peaking time)  
 $F_i, F_e$  "Form Factors" that are determined by the shape of the pulse (calculated in the frequency or time domain)  
 $C_i$  = total capacitance at the input node (detector capacitance + input capacitance of preamplifier + stray capacitance + ... )

- Current noise contribution increases with  $T$
- Voltage noise contribution decreases with increasing  $T$   
 only for "white" voltage & current noise sources + capacitive load  
 "1/f" voltage noise contribution constant in  $T$

# Preamps overview

Experiment	Detector	Q/I	Technology		Power	Noise : $e_n$
ATLAS em	LAr	I	Bipolar	Hybrid	50 mW	0.4 nV/ $\sqrt{\text{Hz}}$
ATLAS had	Tiles + PMT	Q	None			
ATLAS HEC	LAr	I	GaAs	ASIC	108mW	0.8 nV/ $\sqrt{\text{Hz}}$
BABAR	CsI + PD	Q	JFET	Hybrid	50 mW	0.6 nV/ $\sqrt{\text{Hz}}$
CMS em	PbWO <sub>4</sub> +APD	Q	CMOS	ASIC	50 mW	0.9 nV/ $\sqrt{\text{Hz}}$
CMS had	Tiles + HPD	Q	BiCMOS	ASIC		
DØ	LAr	Q/I	JFET	Hybrid	270 mW	0.5 nV/ $\sqrt{\text{Hz}}$
FLC	W/Si	Q	BiCMOS		3 mW	1 nV/ $\sqrt{\text{Hz}}$
KLOE	CsI + PD	Q	Bipolar	Hybrid	60 mW	
LHCb em	PMT	Q	None			
NA48	LKr	I	JFET	Hybrid	80 mW	0.4 nV/ $\sqrt{\text{Hz}}$
Opera TT	PMTMA	Q	BiCMOS	ASIC	5 mW	

# Coherent noise in a multi-channel system

## Coherent noise problem :

- Noise adds linearly instead of quadratically
- Particularly sensitive in calorimetry as sums are performed to reconstruct jets or  $E_t^{\text{miss}}$

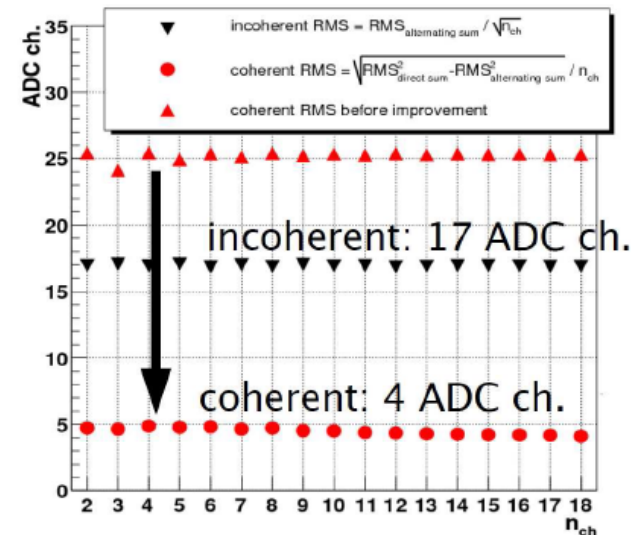
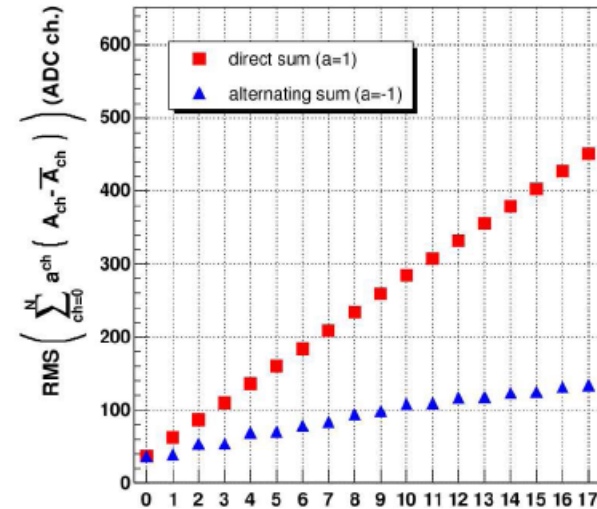
$$\sum a_i^2 = n \sigma_{\text{incoh}}^2 + n^2 \sigma_{\text{coh}}^2 \quad (i=\text{channels})$$

## Coherent noise estimation

- Perform **Direct** and **Alternate** sums to extract coherent noise
- $SD^2 = \sum a_i^2$
- $SA^2 = \sum (-1)^i a_i^2$
- $SA^2 = n \sigma_{\text{incoh}}^2$
- Incoherent & coherent noise :
  - $\sigma_{\text{incoh}}^2 = SA^2/n$
  - $\sigma_{\text{coh}}^2 = (SD^2-SA^2)/n^2$

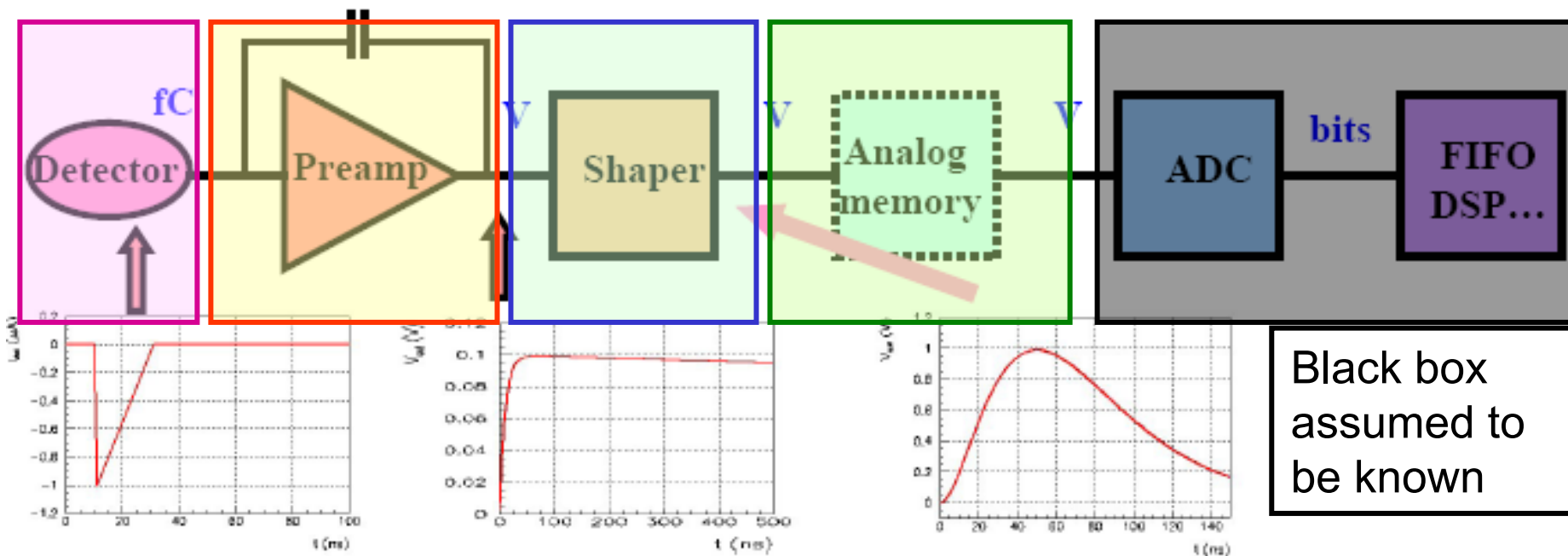
Usually  $\sigma_{\text{coh}} / \sigma_{\text{incoh}} < \sim 20 \%$

Chip 11 - RMS of direct and alternating sums



# Overview of readout electronics

Most front-ends follow a similar architecture

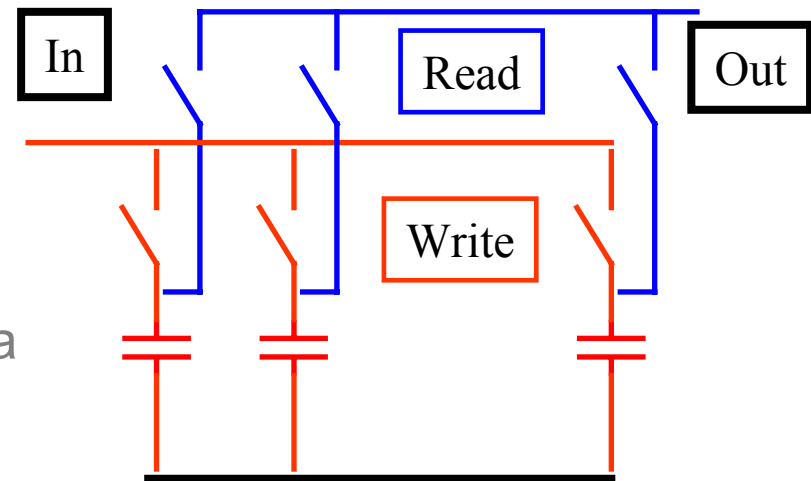


- Very small signals ( $fC$ ) -> need **amplification**
- Measurement of **amplitude** and/or **time** (ADCs, discriminator, TDCs)
- Thousands to millions of channels

# Analog memories

## Switched Capacitor Arrays (SCAs)

- Store signal on capacitors ( $\sim$ pF)
- Fast write ( $\sim$  GHz)
- Slower read ( $\sim$ 10MHz)
- Dynamic range : 10-13 bits
- depth : 100-2000 caps
- Insensitive to absolute value of capa (voltage write, voltage read)
- **Low power**
- Possible loss in signal integrity (droop, leakage current)



Principle of a « voltage-write, voltage-read » analog memory

The base of 90% of digital oscilloscopes !

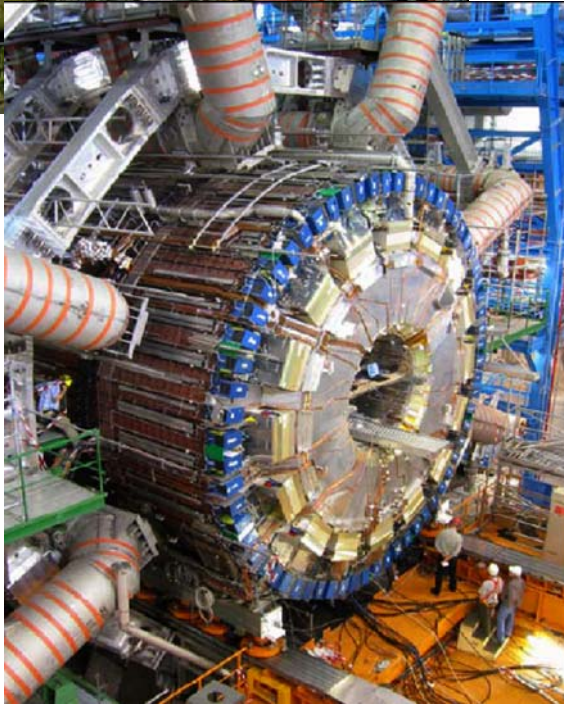
# Ionization calorimeters



Examples:  
DØ(LAr)  
NA48 (LKr)  
ATLAS (LAr)  
H1



Stable, Linear  
Easy to calibrate (!)  
Moderate resolution



# ATLAS LAr: Front End boards

**Amplify, shape, store and digitize LAr signals**

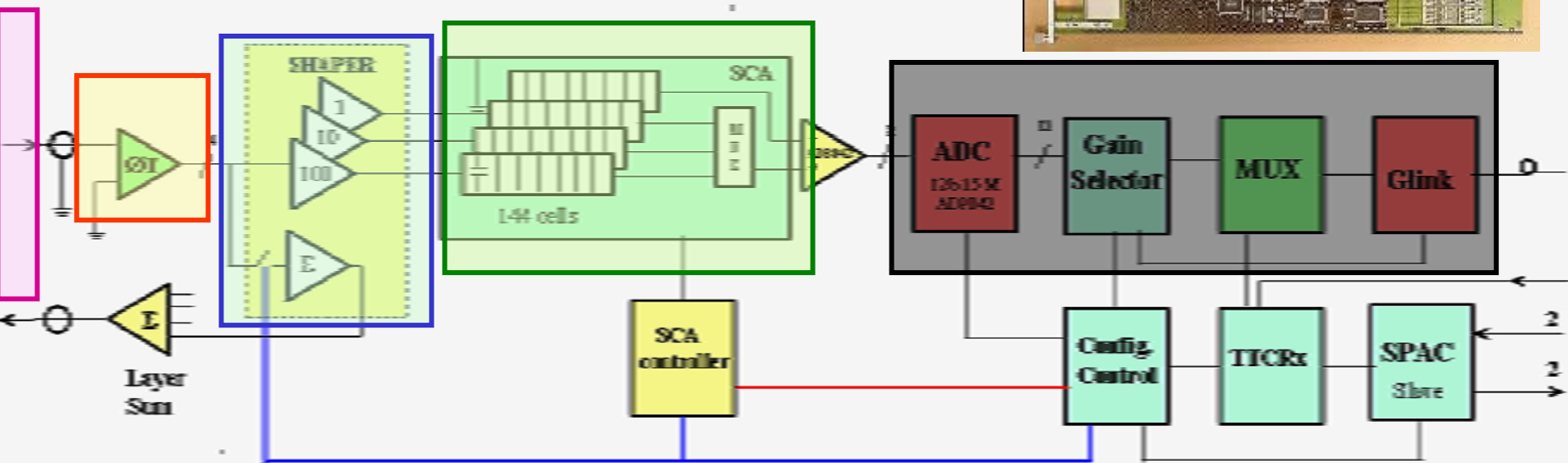
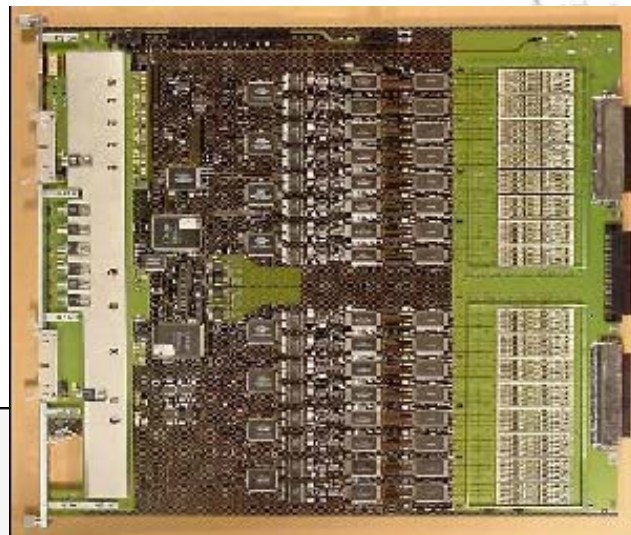
128 warm preamps

128 tri-gain shapers

128 quad pipelines

32 ADCs (12bits 5 MHz)

1 optical output (Glink)



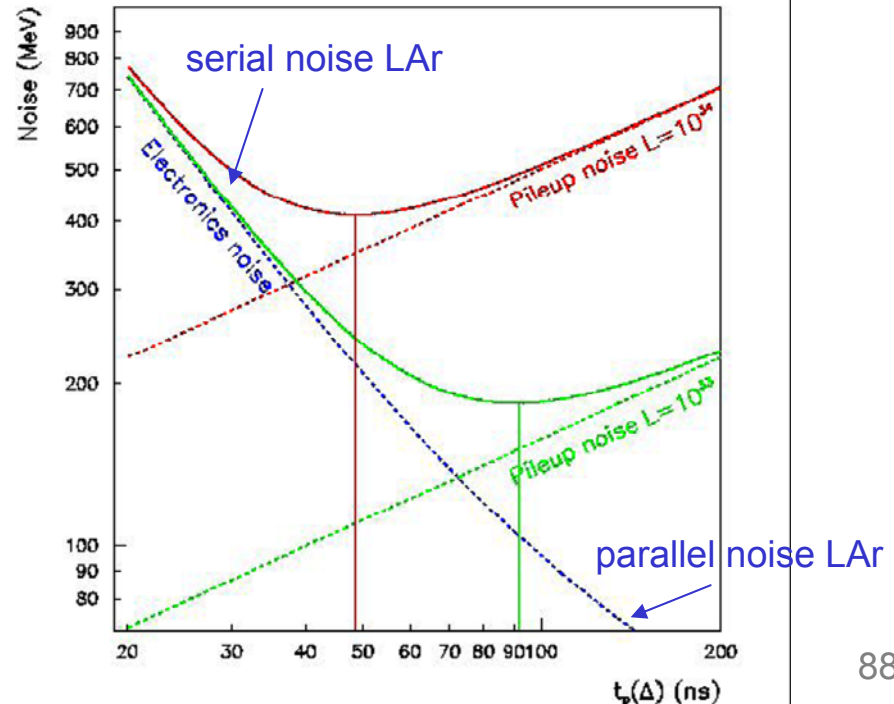
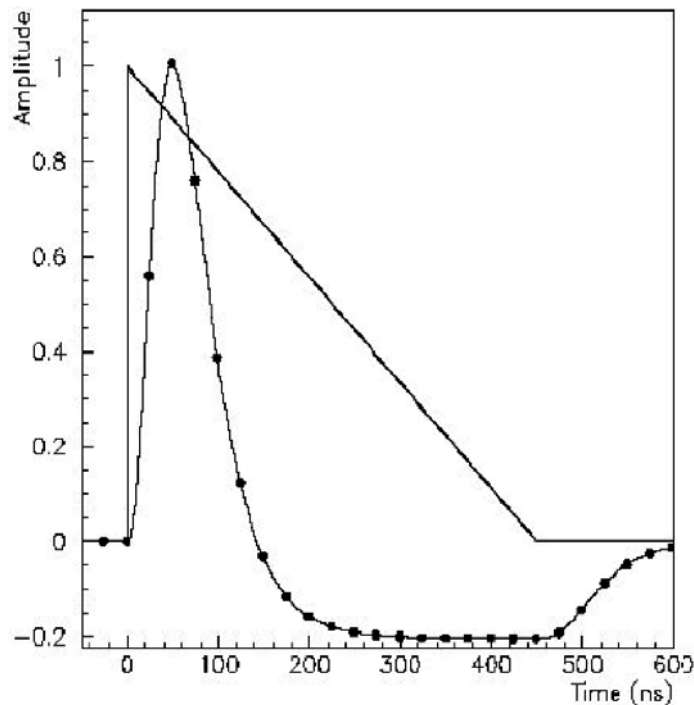
# ATLAS : LAr shaper

**Goal : optimize signal to noise ratio between electronics noise and pileup noise**

Ionization signal  $\sim 500\text{ns} = 20$  LHC bunch Xings

Reduced to 5 bunch Xings with fast shaper  $\rightarrow$  worse S/N due to loss of charge

Choice of peak time varies with luminosity  $\rightarrow 45\text{ns}$  at  $L=10^{34}\text{cm}^{-2}\text{s}^{-1}$

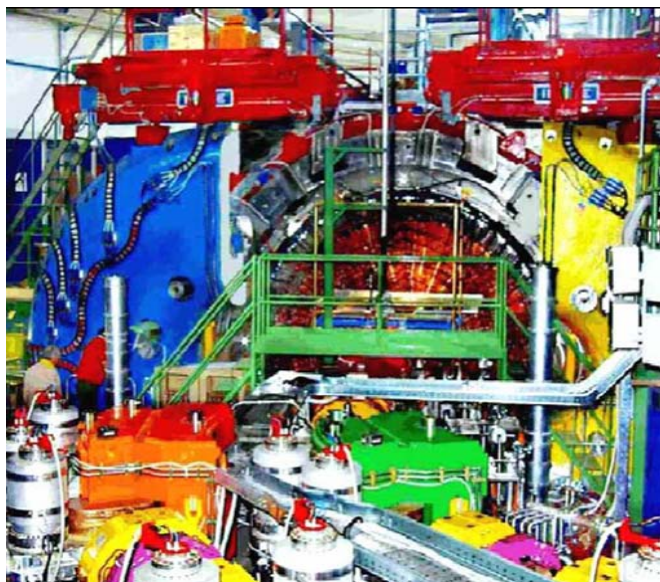




# Crystal calorimeters



Babar(CsI)  
Kloe(CsI)  
CMS (PbWO<sub>4</sub>)  
L3, CLEO, Belle,  
ALICE



Fast  
Best resolution  
Difficult to calibrate  
expensive

# CMS: ECAL Electronics

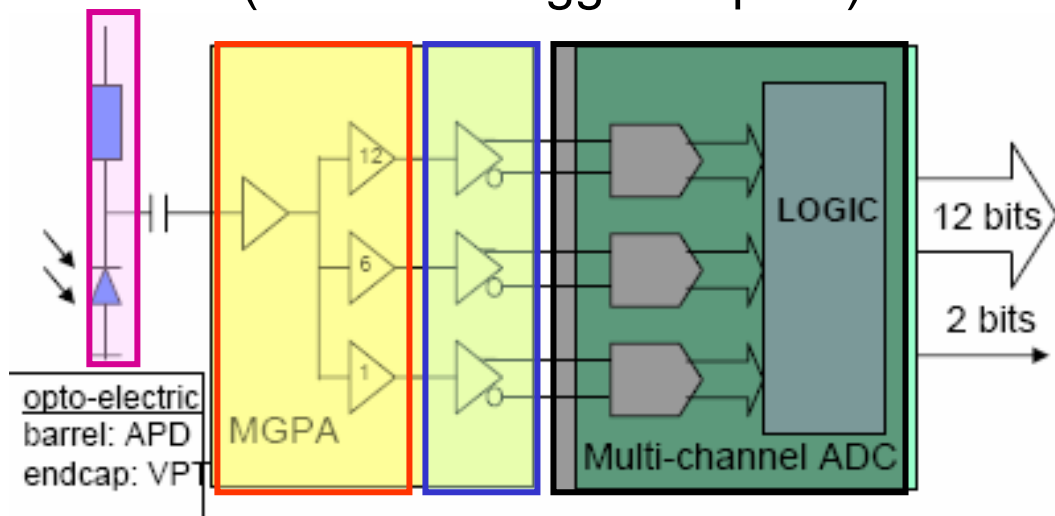
building block :

## Trigger Tower (25 channels)

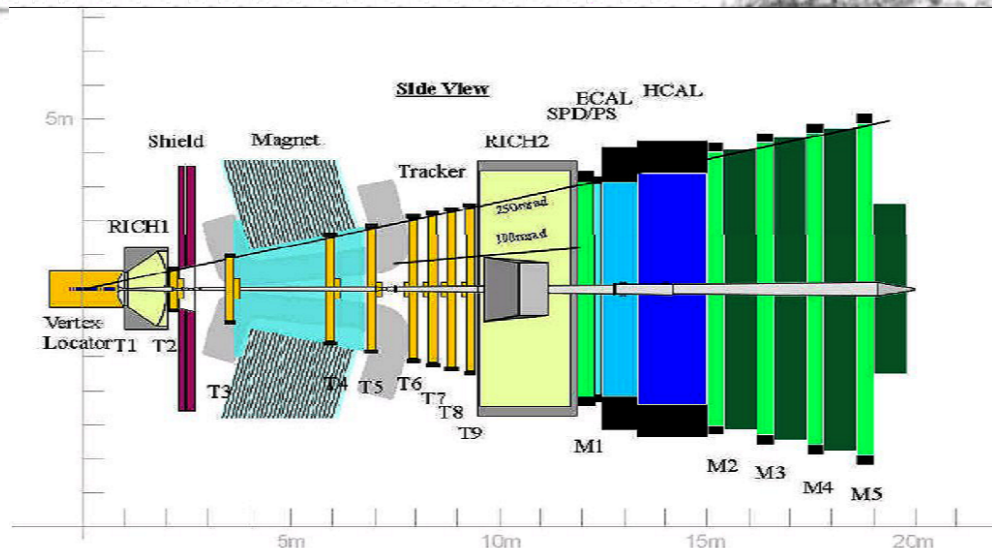
- 1 mother board
- 1 LV regulator board
- 5 VFE boards (5 channels each)
- 1 FE board

2 fibres per TT sending

- trigger primitives (every beam crossing)
- data (on level 1 trigger request)

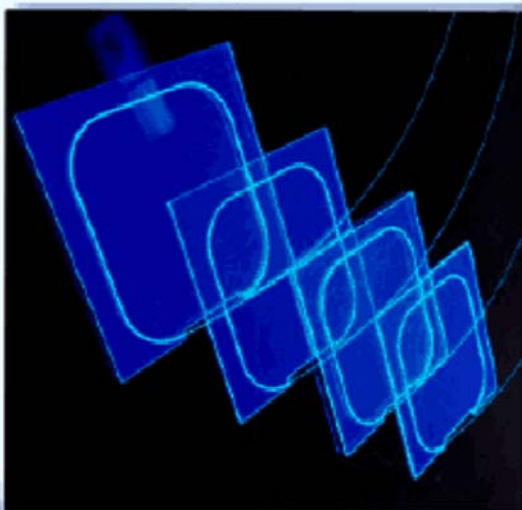


# Scintillating calorimeters



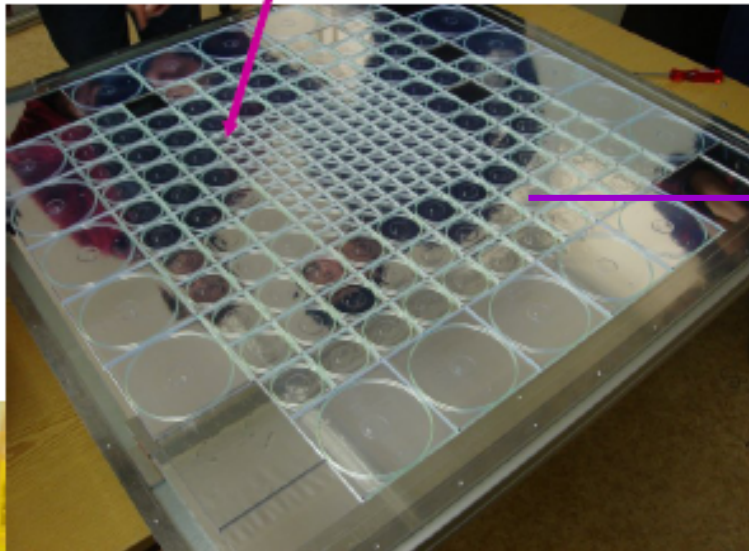
CMS hadronic  
LHCb  
OPERA  
ILC hadronic  
ILC em  
ATLAS hadronic

Fast  
Cheap  
Moderate resolution  
Difficult to calibrate

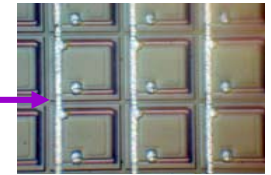
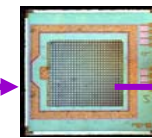
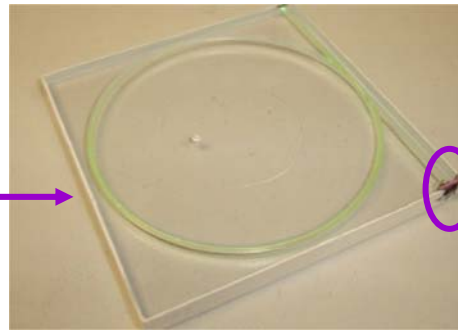


# ILC: hadronic calorimeter (CALICE)

Iron/**plastic(tiles)** sandwich



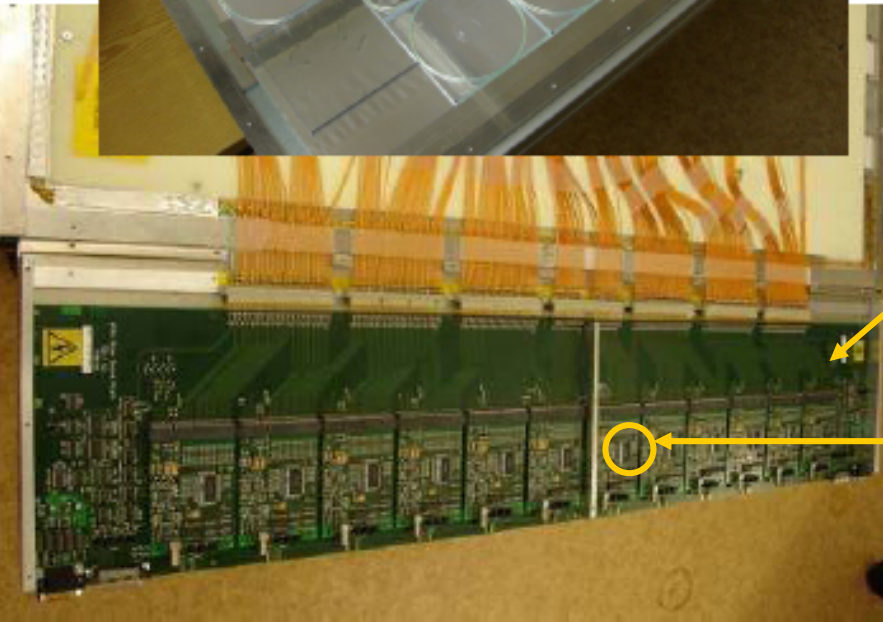
Single tile readout with WLS fiber + SiPM:  
pixel device operated  
in Geiger mode



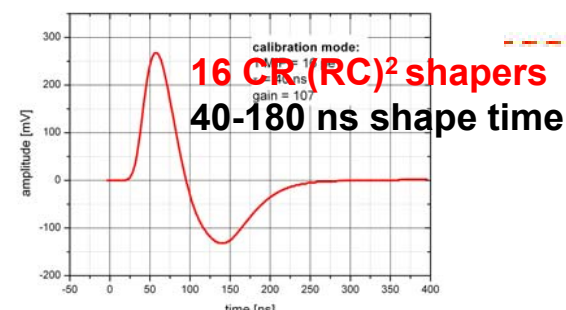
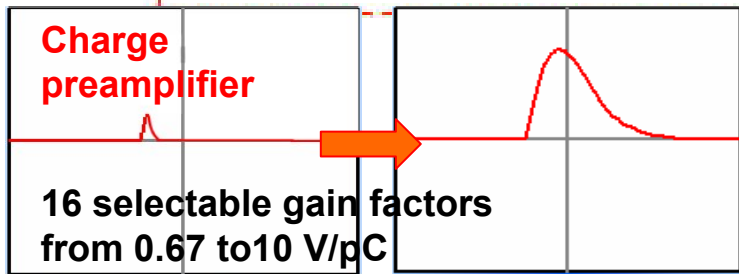
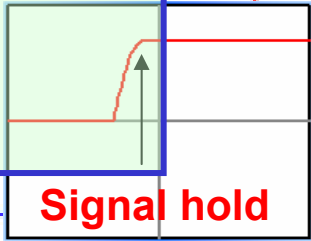
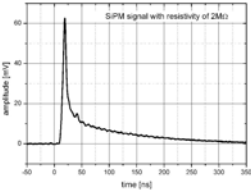
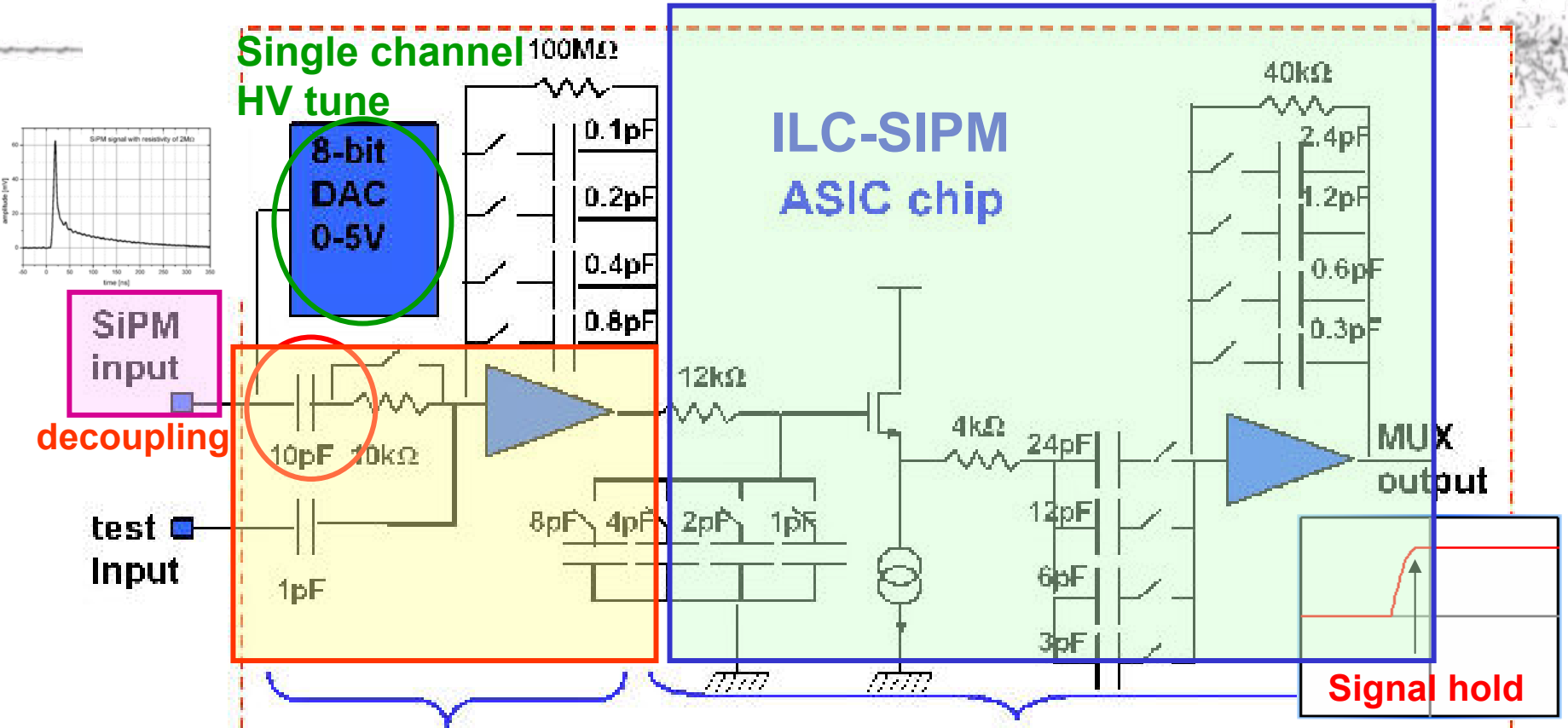
Read out 216 tiles/module  
38 sampling layers  
~8000 channels

VFE: control board for 12 ASICs / layer  
connect to SiPMs

ASIC: amplification + shaping +  
multiplexing (18 ch.)

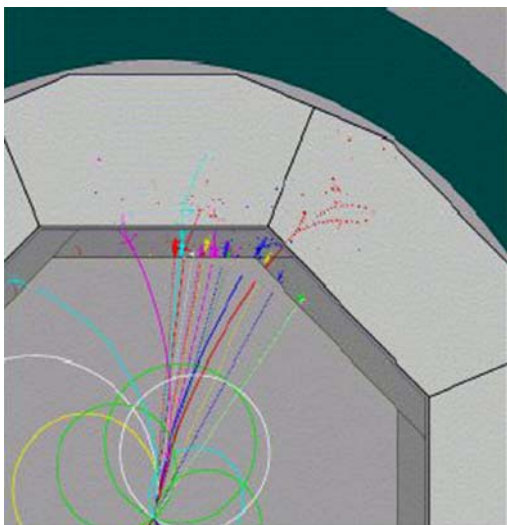
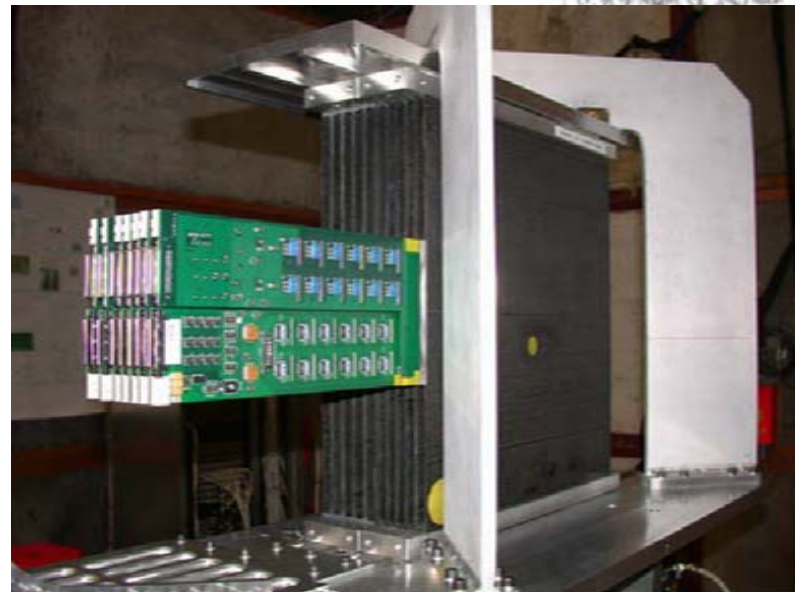
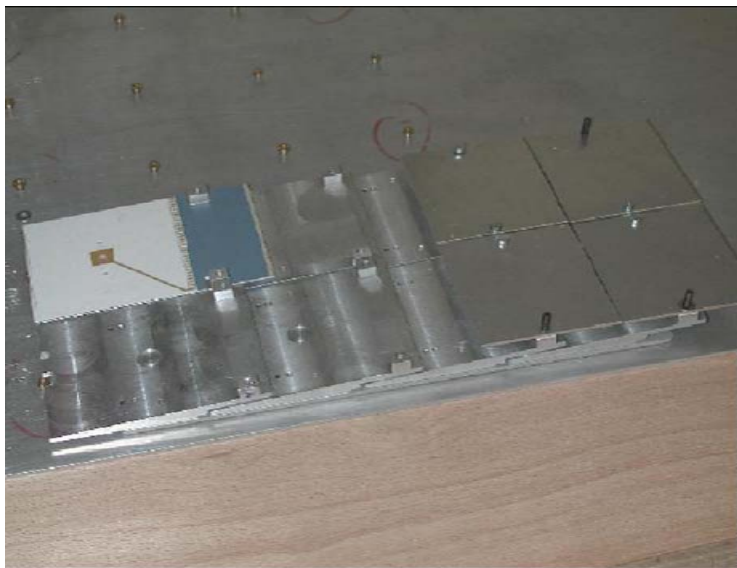


# ILC: HCAL readout chip



- The

# Semiconductors calorimeters

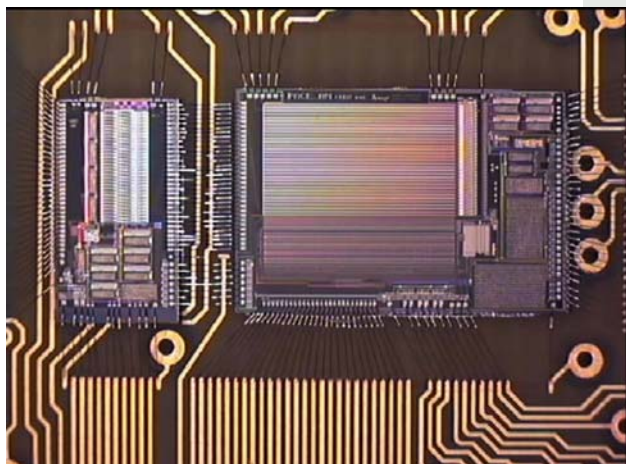
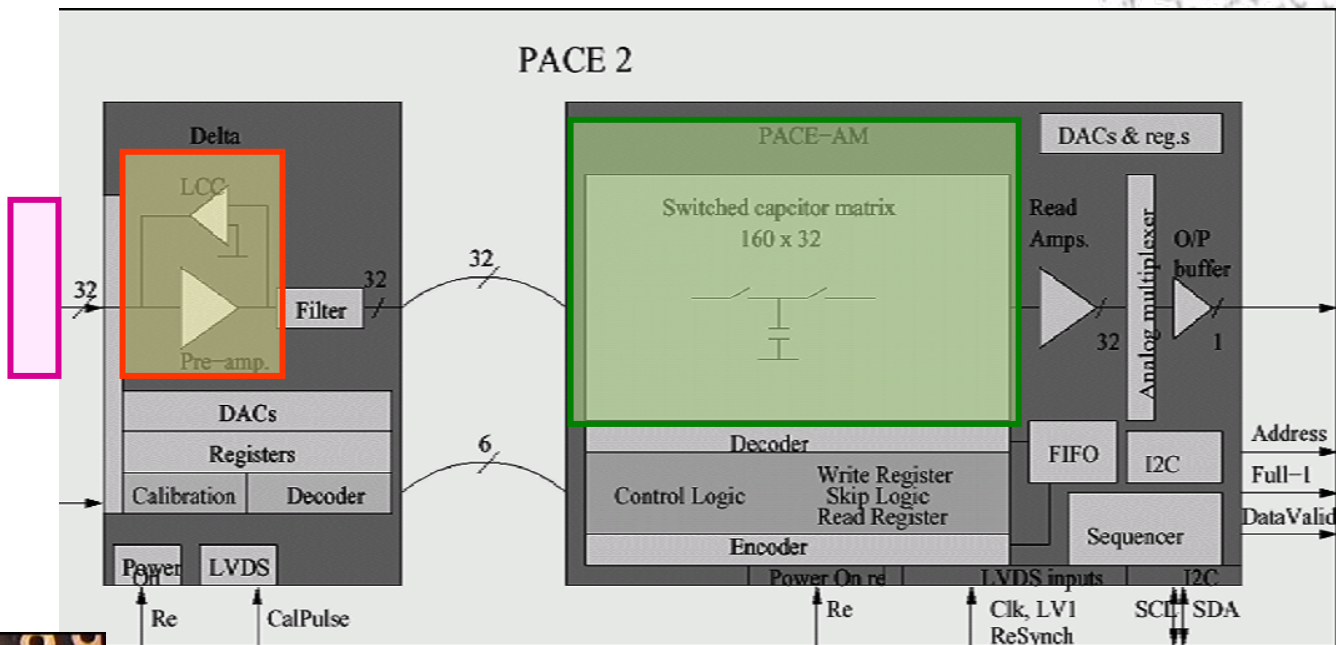


CMS pre-shower  
ILC CALICE ECAL

Highly granular  
Good resolution  
Expensive

# CMS PreShower : readout chip PACE2

- Aluminium
- Silicon sensor:  
 $1 \text{ cm}^2$   
 $V_{\text{depl}} = (60 \pm 5) \text{ V}$   
 $I_{\text{leak}} = 100 \text{ nA}$



## MCM

Delta preamp  
 PACE analogue memory

## Requirements

- 10 bit dynamic range
- Low gain and high gain
- Leakage current comp

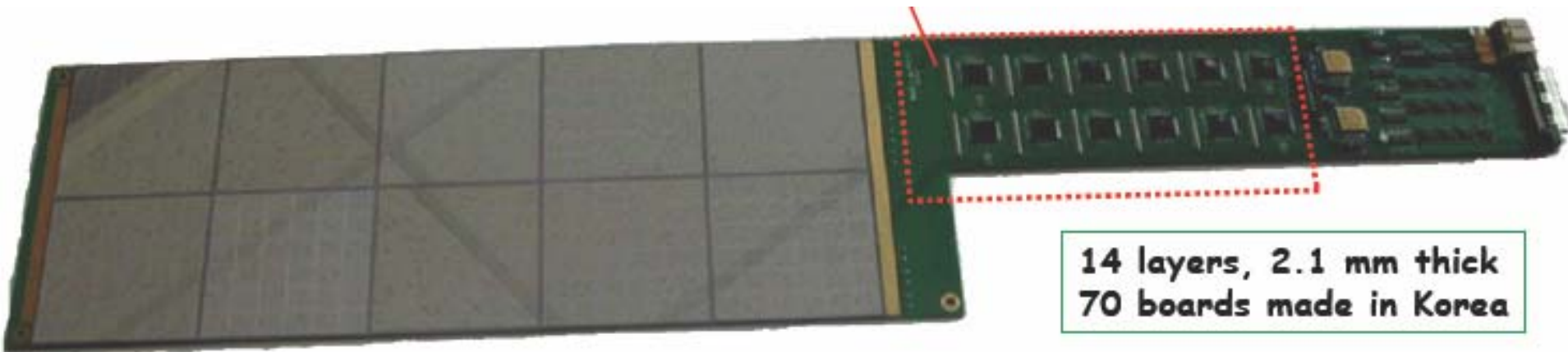
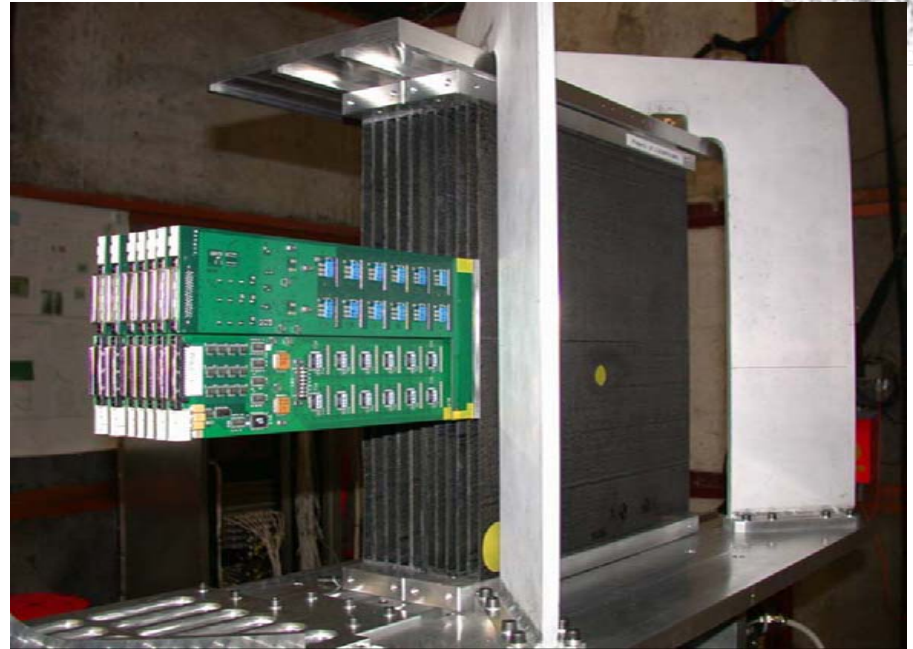
# ILC: W-Si em calorimeter (CALICE)

## “Imaging calorimeter”

30 layers W-Si

1 cm<sup>2</sup> SiPADs

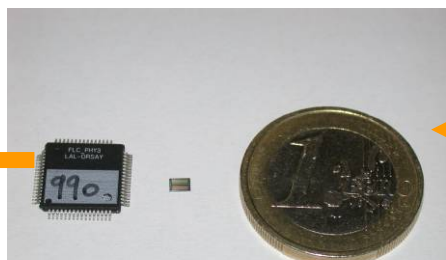
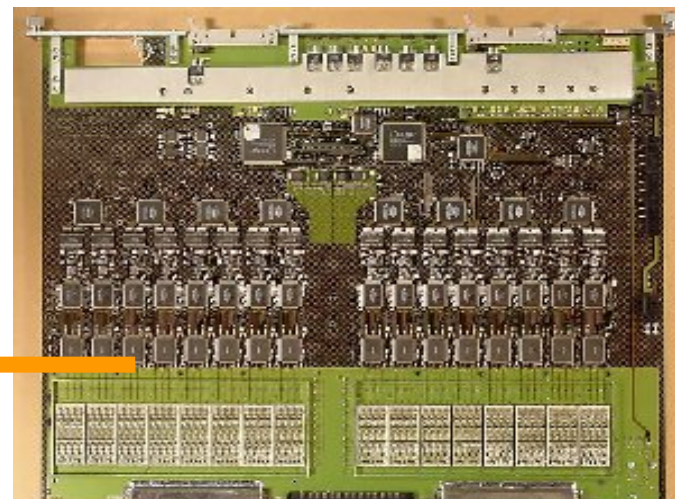
~10000 channels





# Trends & Future

- More channels / more functionality in the chip (analog+digital)
  - more integration
- Detector imbedded electronics
  - reduce cable volume = dead volume
  - ultra-low power consumption

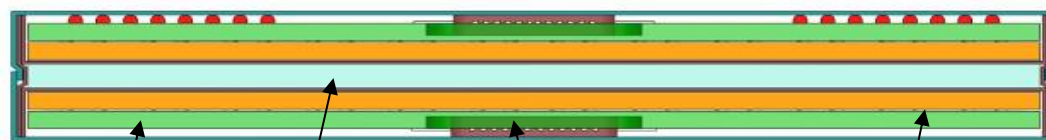
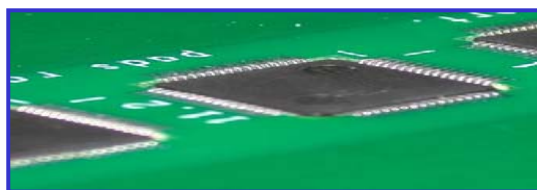


ILC : 100 $\mu$ W/ch

FLC\_PHY3 18ch 10\*10mm 5mW/ch

ATLAS LAr FEB 128ch 400\*500mm 1 W/ch

## Readout chip integrated in active layer (Si-W ECAL for ILC)



PCB (600 $\mu$ m)

Tungsten (1 mm)

FE chip (1mm)

Wafer (400 $\mu$ m)

# Imbedded electronics (ILC ECAL)

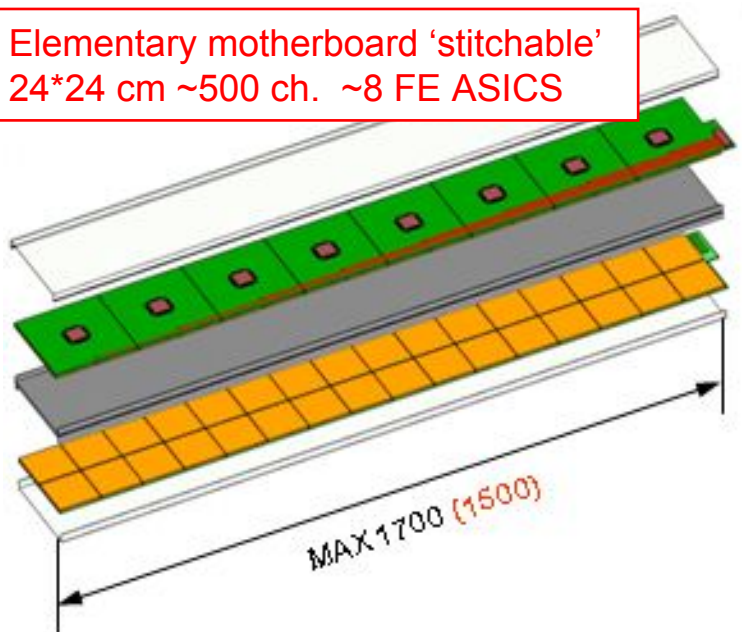
Front-end ASICs embedded in detector

- Very high level of integration
- Ultra-low power with **pulsed mode**
- **Target 0.35  $\mu\text{m}$  SiGe technology**

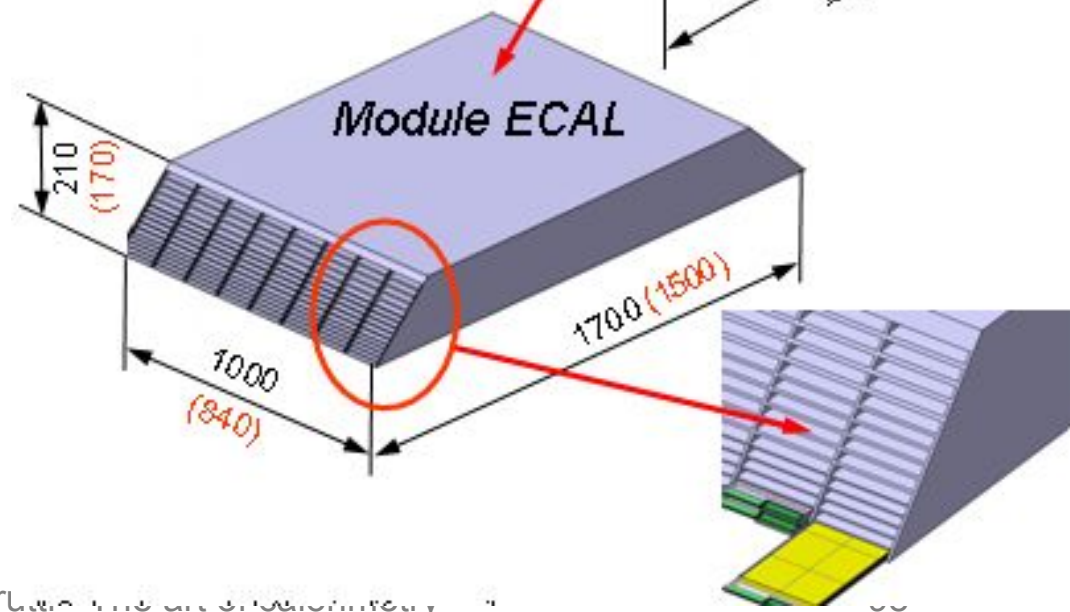
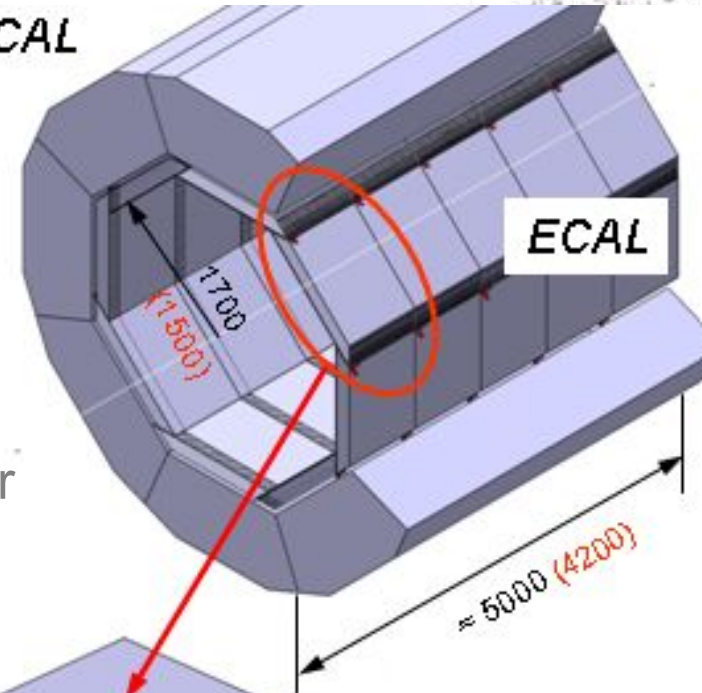
All communications via edge

- 4,000 ch/slab, minimal room, access, power
- small data volume ( $\sim$  few 100 kbyte/s/slab)

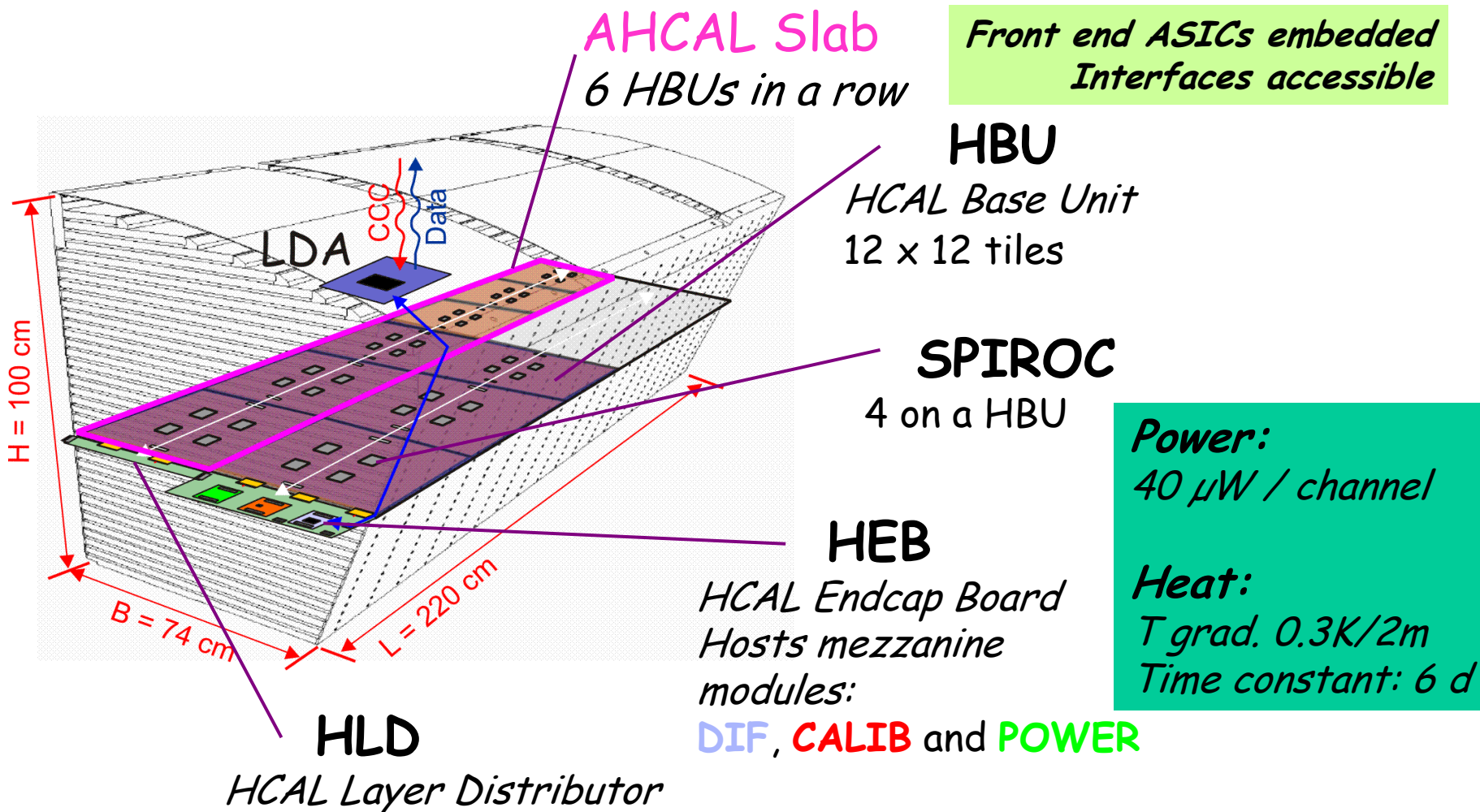
Elementary motherboard 'stitchable'  
24\*24 cm  $\sim$ 500 ch.  $\sim$ 8 FE ASICS



HCAL

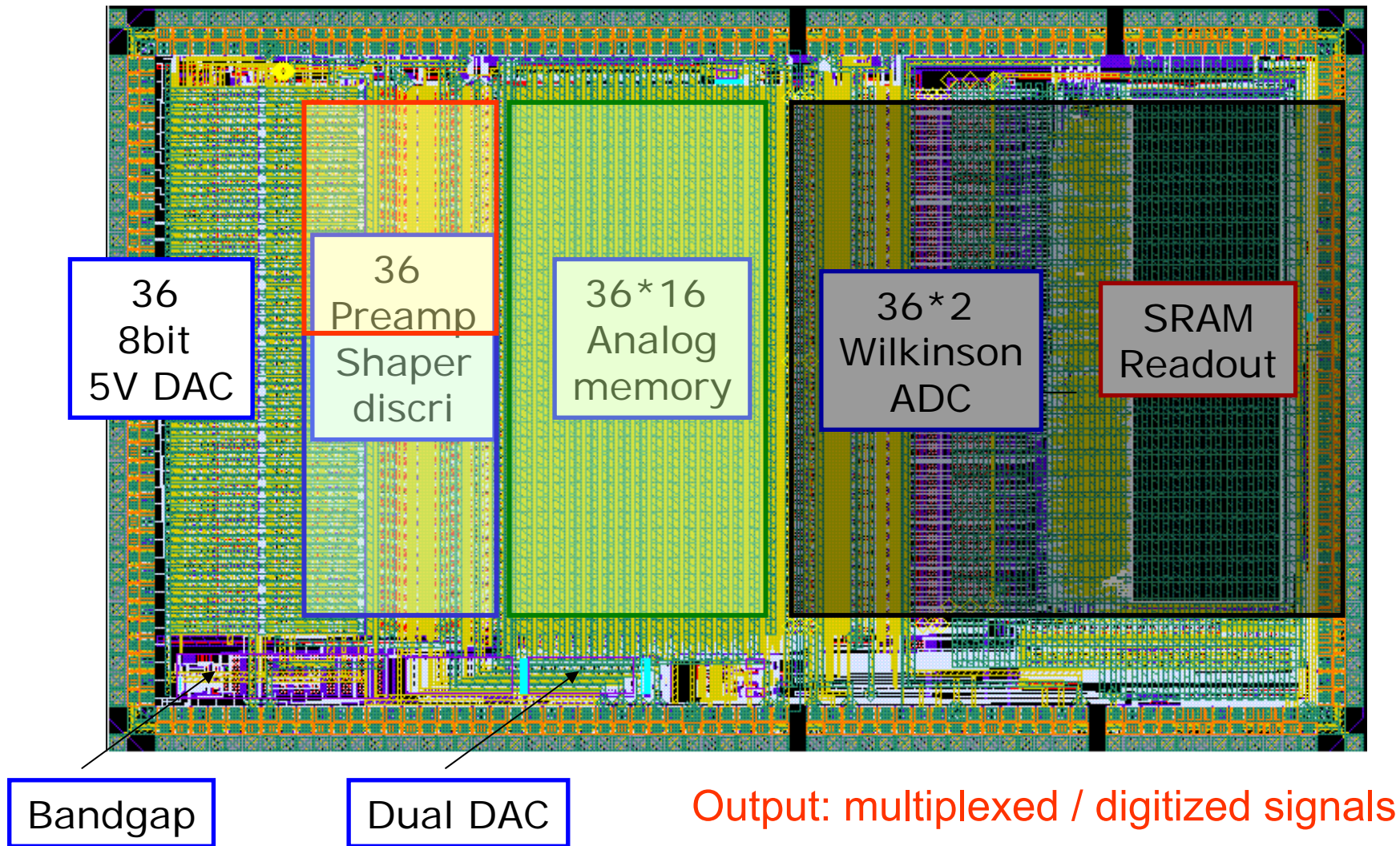


# Imbedded electronics (ILC HCAL)



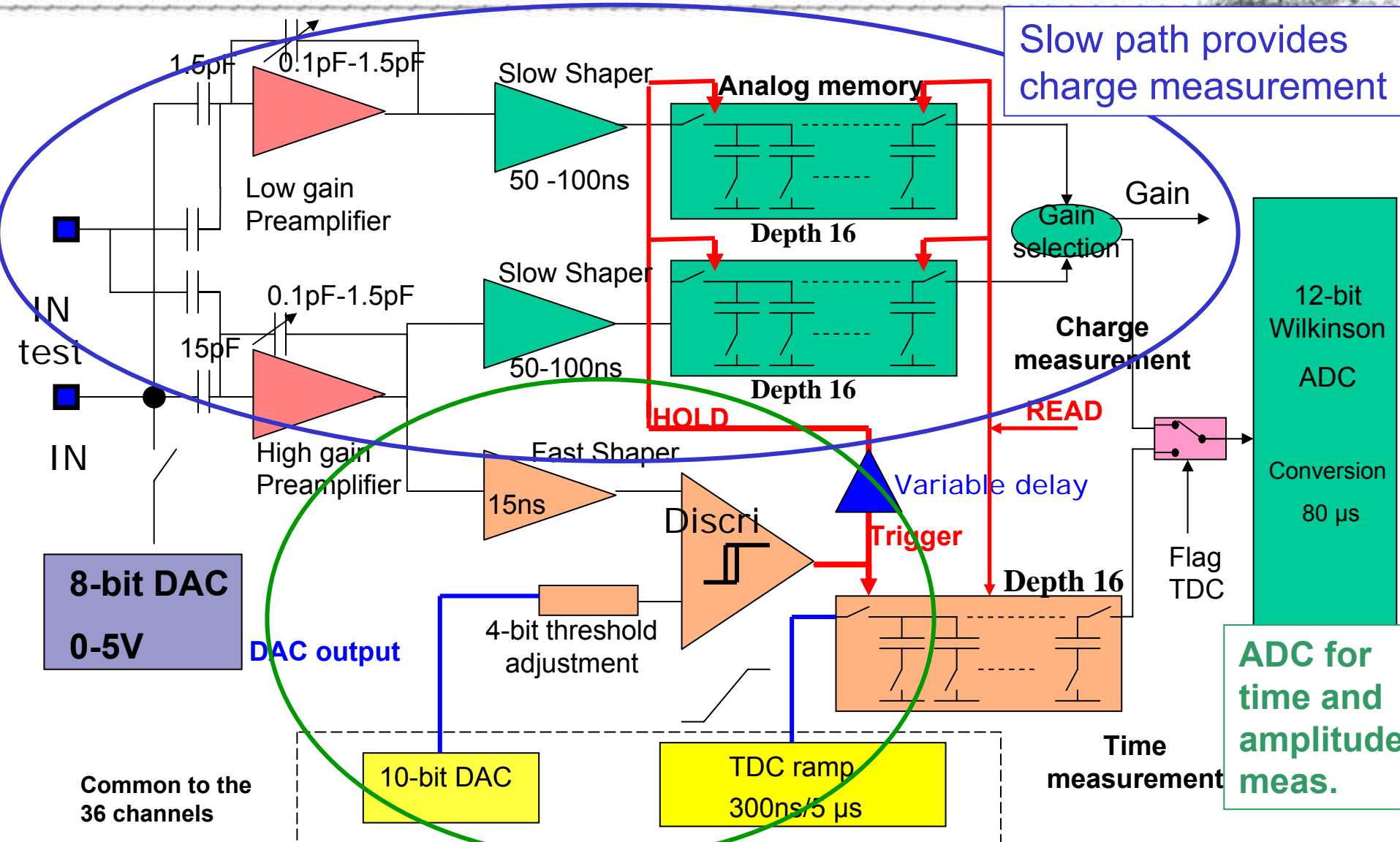
# More pixels / more functionality

SPIROC layout (CALICE chip for Analog HCAL readout)





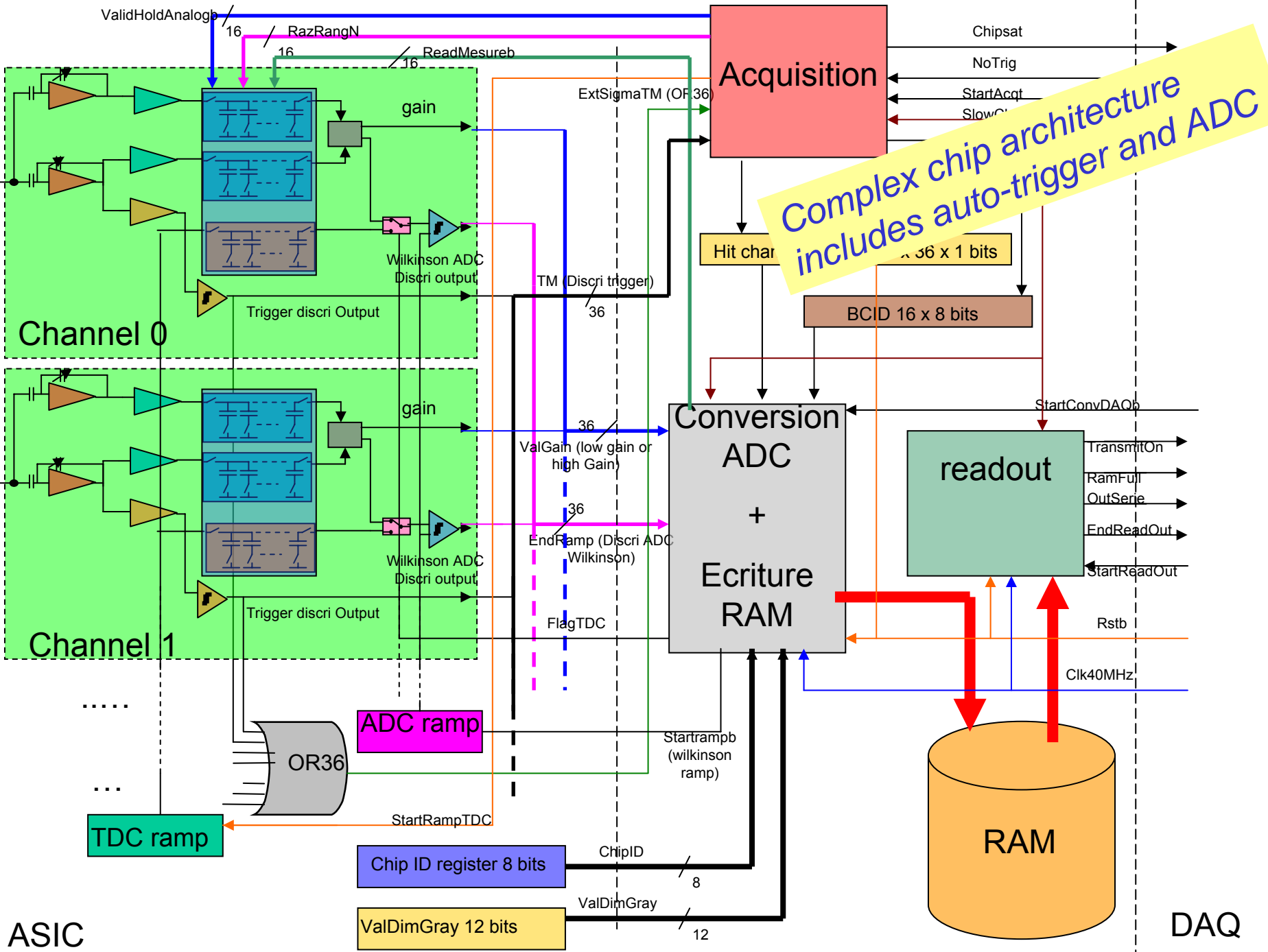
# SPIROC : One channel schematic



Slow path provides charge measurement

Fast path provides auto-trigger and time measurement

ADC for time and amplitude meas.



# Acknowledgments

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For these slides I have to give credit to the work of  
C. D'Ambrosio, Dieter Renker, H. Spieler, C. de La Taille  
from whom I have taken many plots and figures